

DIS and Hadronization systematic errors

0. Introduction
1. DIS Bodek-Yang parameterization
2. DIS differential cross section error
3. DIS A-scaling error
4. DIS PDF error
5. Low- W hadronization error
6. High- W hadronization error
7. Conclusion

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Teppei Katori and Shivesh Mandalia
NuWro Workshop, Dec. 3, 2017



2011



Kochom cię, Wrocław!

2017

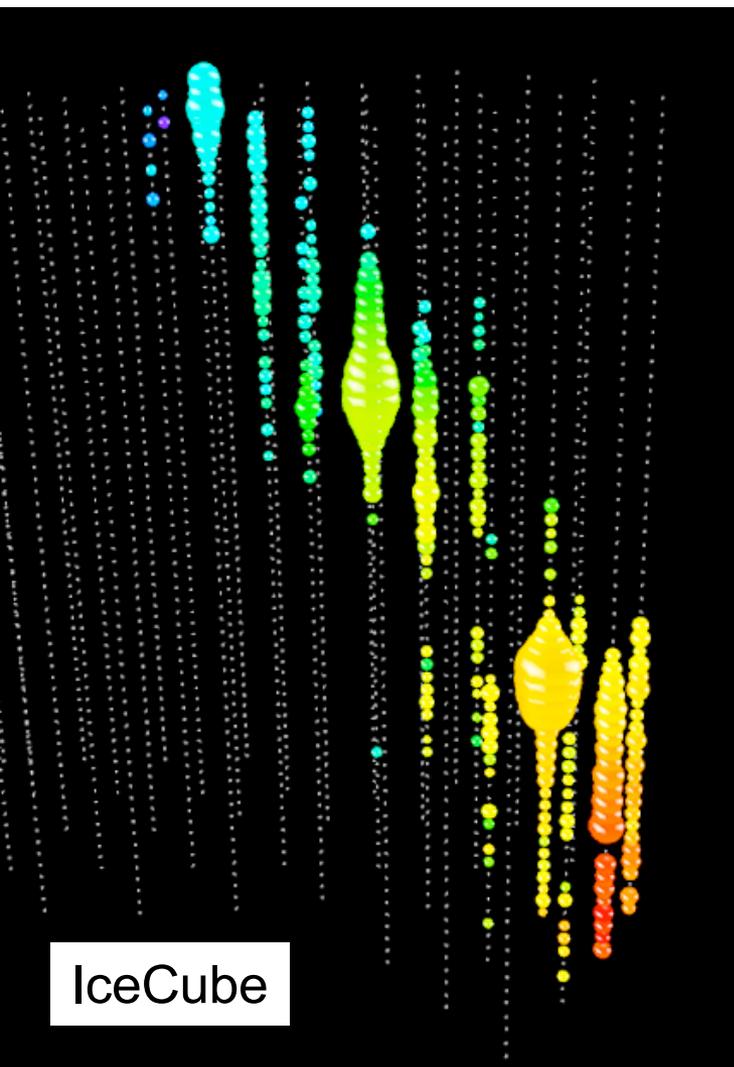
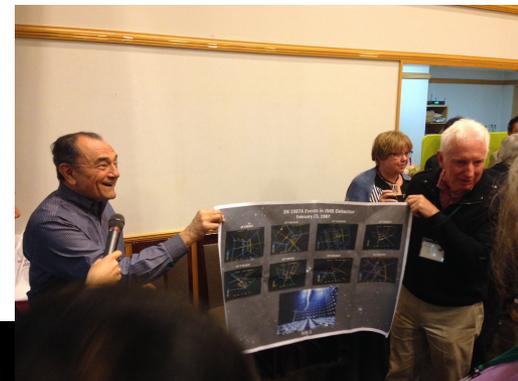


Supernova neutrino (1987, Feb. 23) 30 years celebration

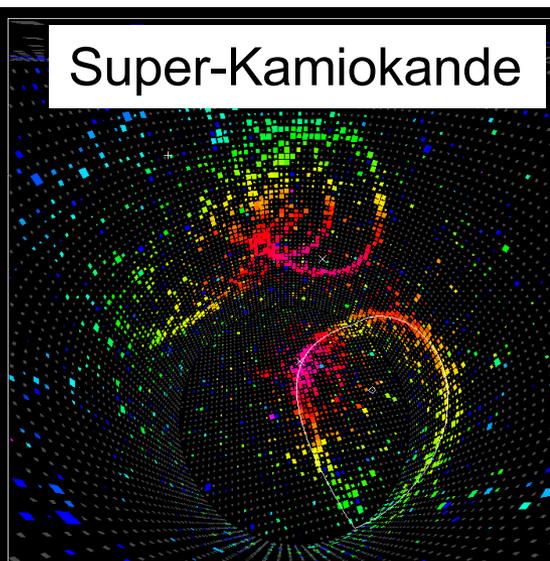
- 8 neutrinos are detector by IMB detector (Ohio, USA)

Prof. Hank Sobel (UC Irvine) and Prof. Bob Svoboda (UC Davis)

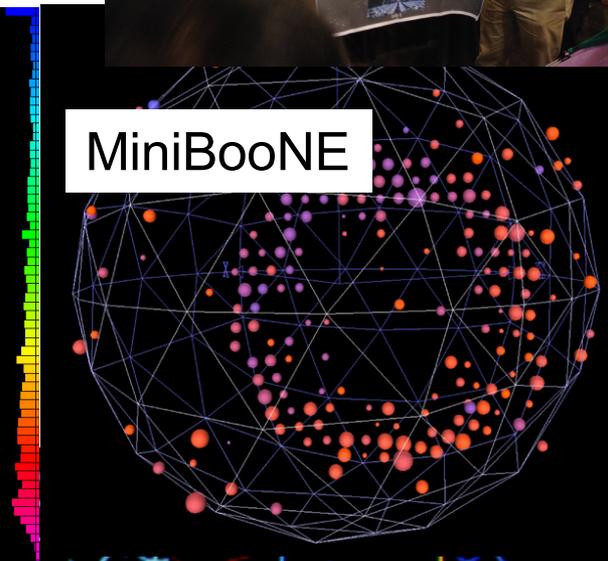
- They showed event display of 8 supernova neutrinos detected at IMB detector.



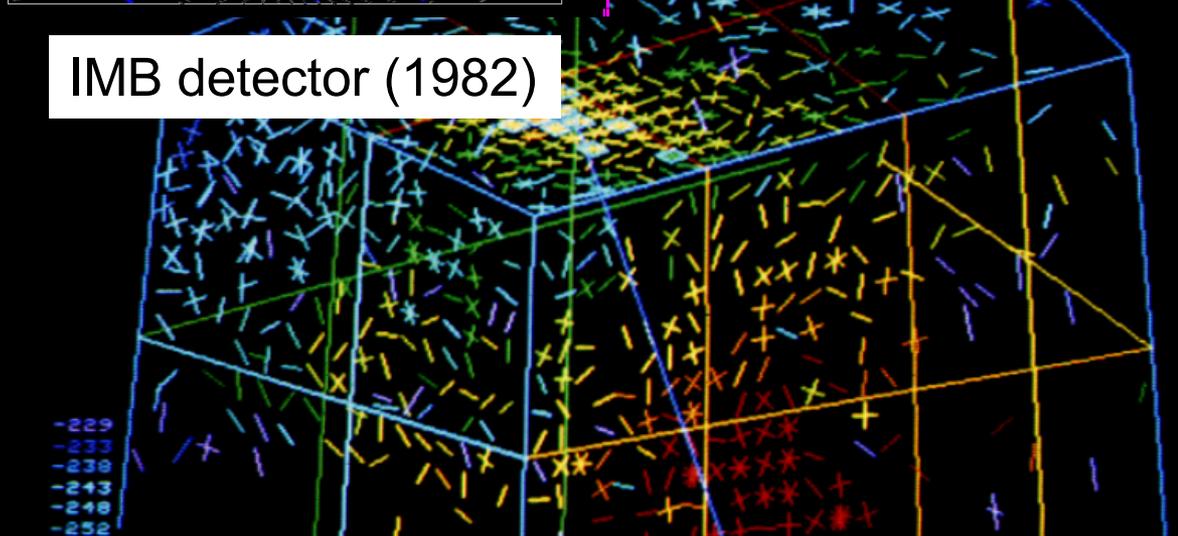
IceCube



Super-Kamiokande



MiniBooNE

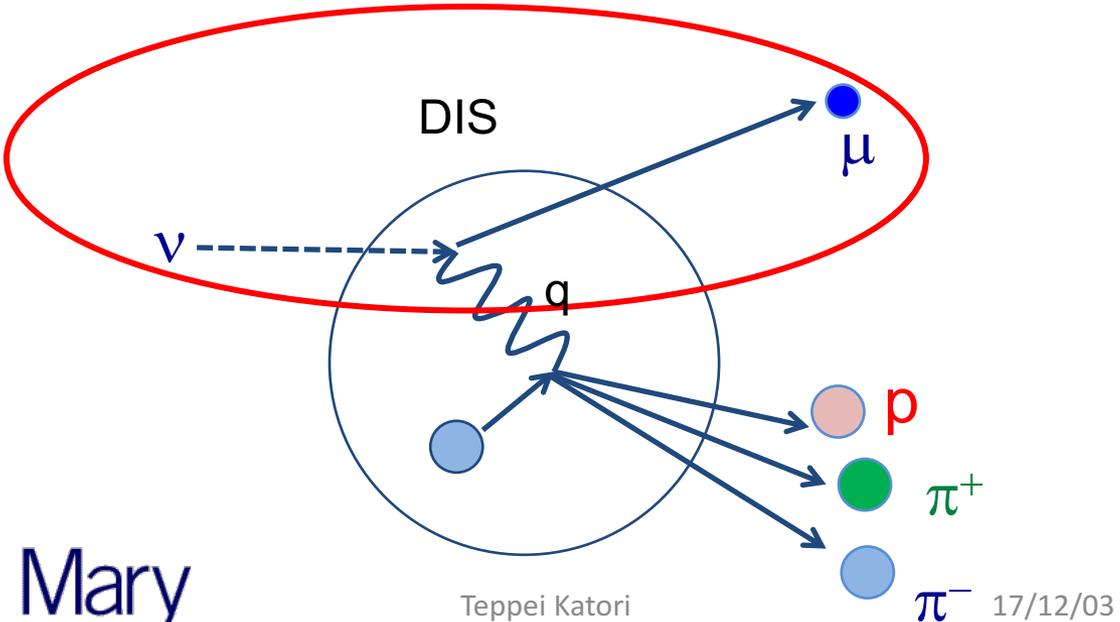


IMB detector (1982)

0. Neutrino cross section overview

Deep Inelastic Scattering (DIS)

- a process to scatter a charged lepton by an incident lepton with given energy
- DIS **differential cross section** is function of x and y
- DIS **total cross section** is function of E_ν , integrated in x and y



0. Neutrino cross section overview

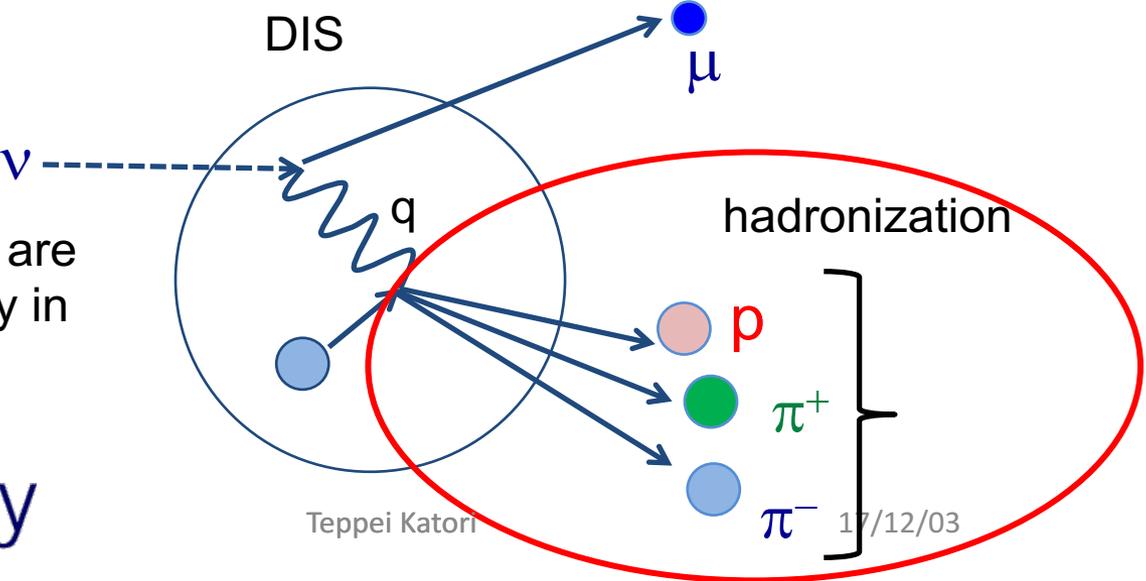
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Hadronization

- Hadronization is a process to generate hadrons from given Q^2 and W
- number of hadrons (multiplicity) and hadrons kinematics are computed.

DIS and Hadronization are modelled independently in simulation



0. Neutrino cross section overview

Deep Inelastic Scattering (DIS)

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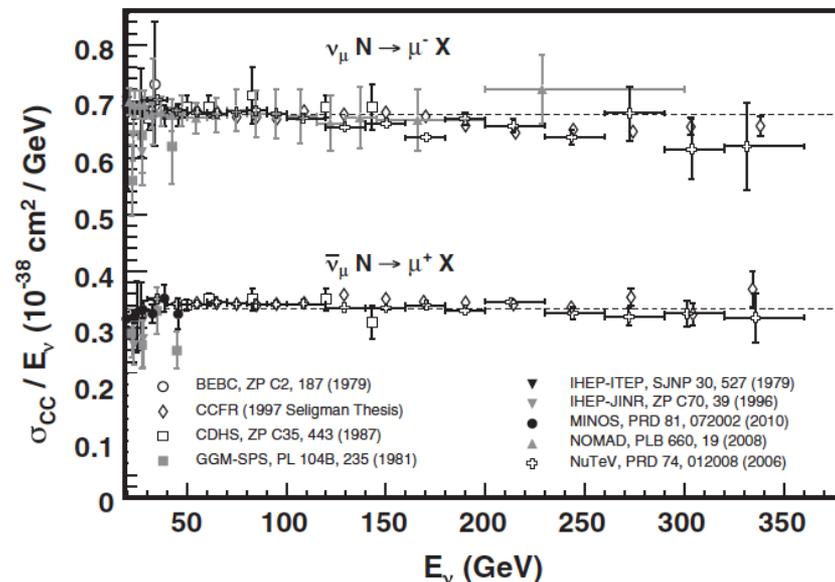
Hadronization

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- number of hadrons (multiplicity) and hadrons kinematics are computed.

$$- \sigma(\nu)/E = 0.677 \pm 0.014 \times 10^{-38} \text{ (cm}^2/\text{GeV)}$$

DIS total cross section error $\sim 2\%$?

- This is the error of CCDIS total cross section at 30 to 200 GeV
- Most of our analyses need errors of differential cross section error



0. SIS region physics

Basic ingredients

- $\Delta(1232)$ -resonance
- higher resonances
- non-resonant background

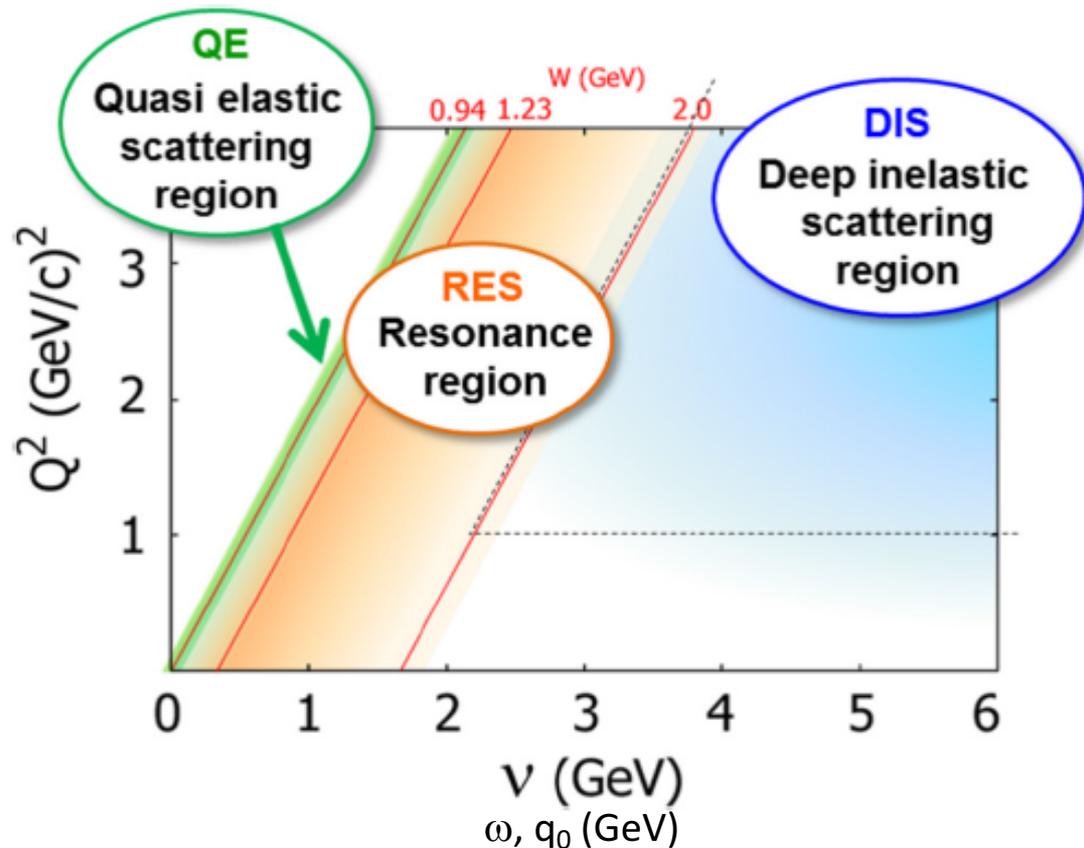
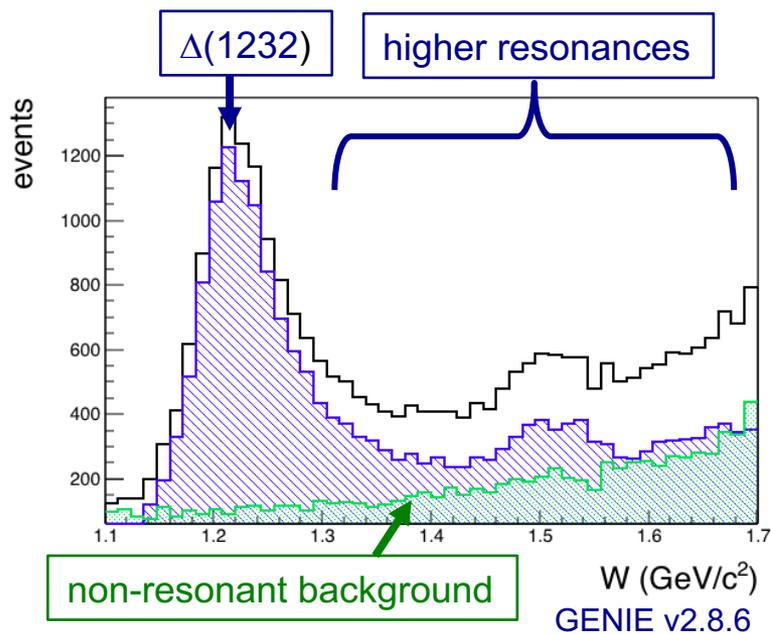


Figure 1. Kinematical regions of the neutrino-nucleus interaction relevant to the next-generation neutrino-oscillation experiments. The energy transfer to a nucleus and the squared four-momentum transfer are denoted by ν and Q^2 , respectively.

0. GENIE SIS model

Cross section

$W^2 < 2.9 \text{ GeV}^2$: RES

$W^2 > 2.9 \text{ GeV}^2$: DIS

Hadronization

$W^2 < 5.3 \text{ GeV}^2$: KNO scaling based model

$2.3 \text{ GeV}^2 < W^2 < 9.0 \text{ GeV}^2$: transition

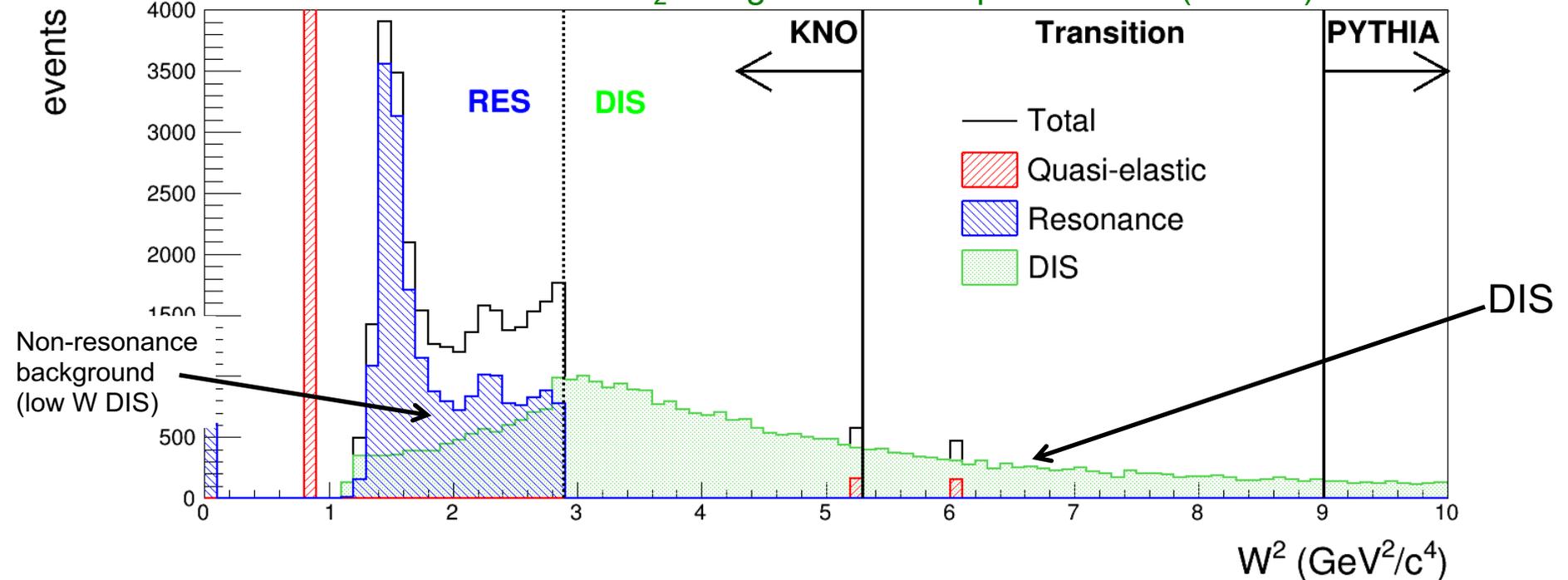
$9.0 \text{ GeV}^2 < W^2$: PYTHIA6

There are 2 kind of “transitions” in SIS region

- cross-section
- hadronization

W^2 distribution for H_2O target with atmospheric- ν flux (GENIE)

GENIE v2.8.0



0. NEUT SIS model

Cross section

$W^2 < 4 \text{ GeV}^2$: RES

$W^2 > 4 \text{ GeV}^2$: DIS

Hadronization

$W^2 < 4 \text{ GeV}^2$: KNO scaling based model

$4 \text{ GeV}^2 < W^2$: PYTHIA5

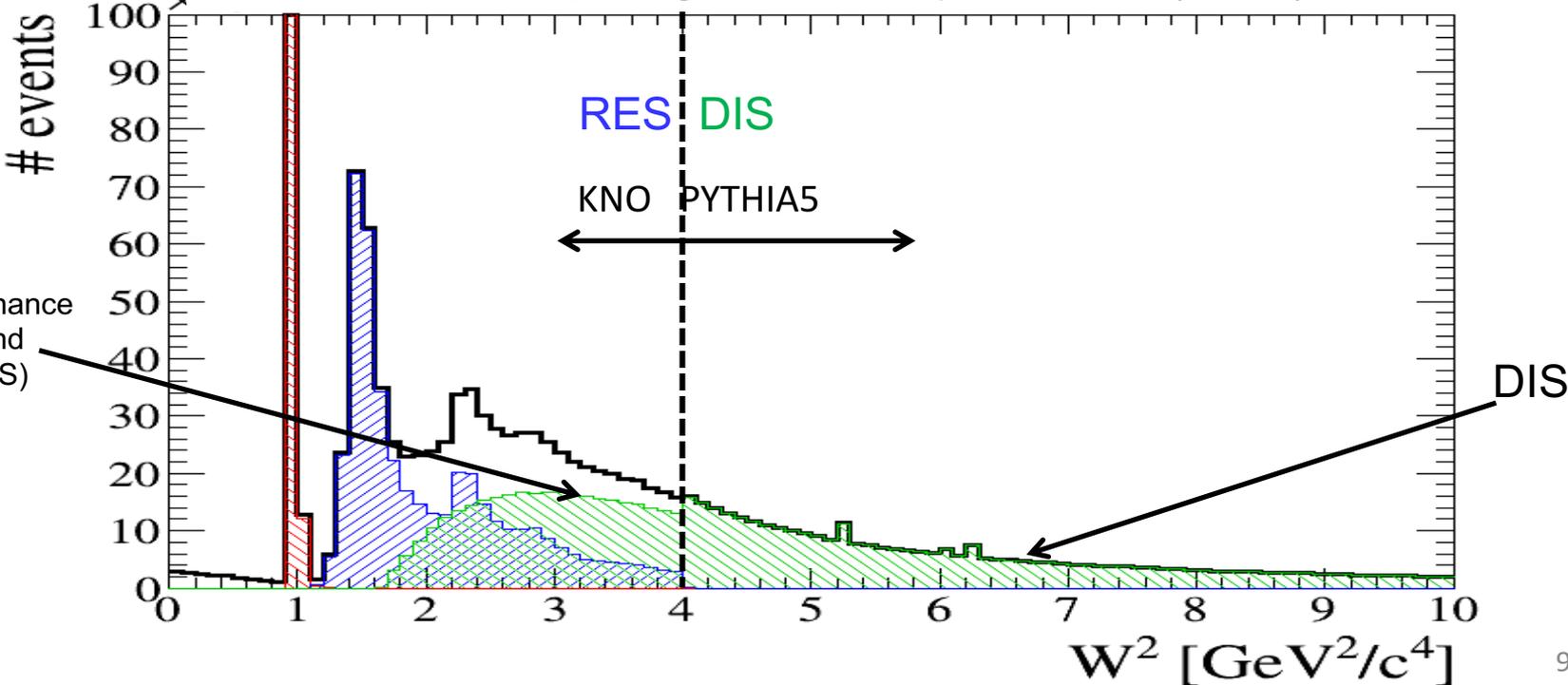
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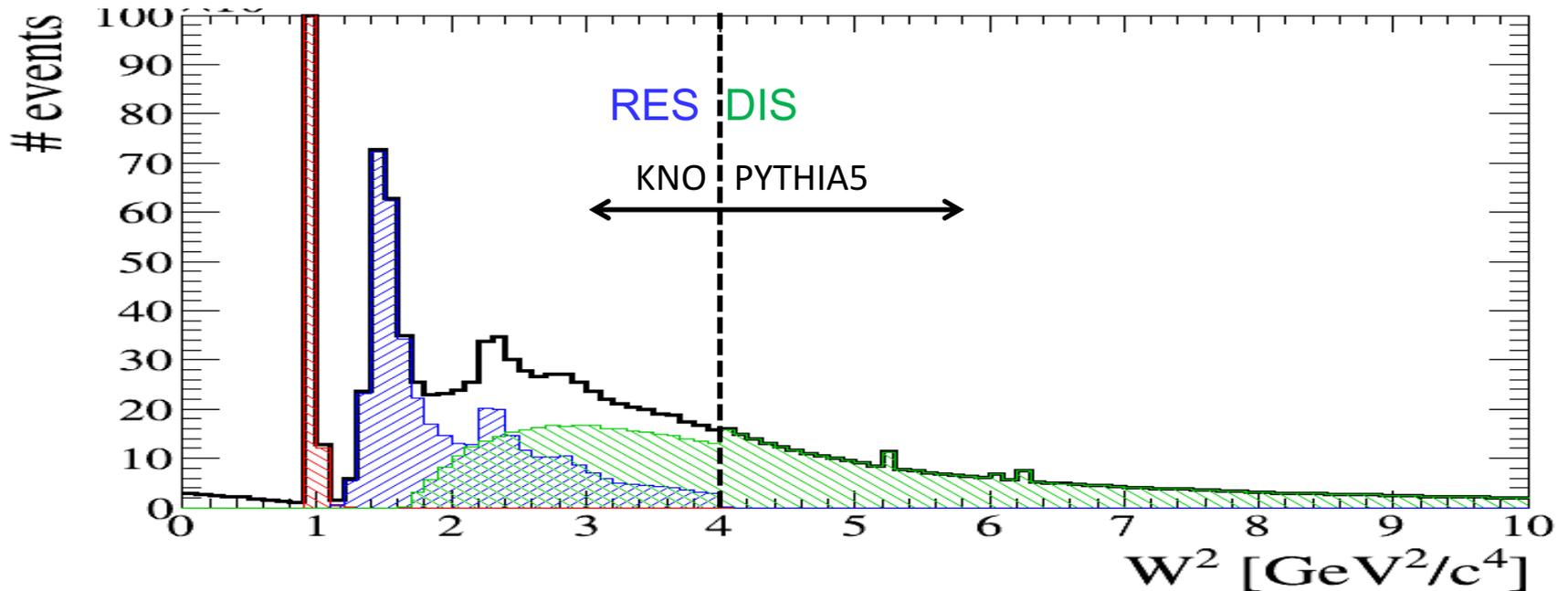
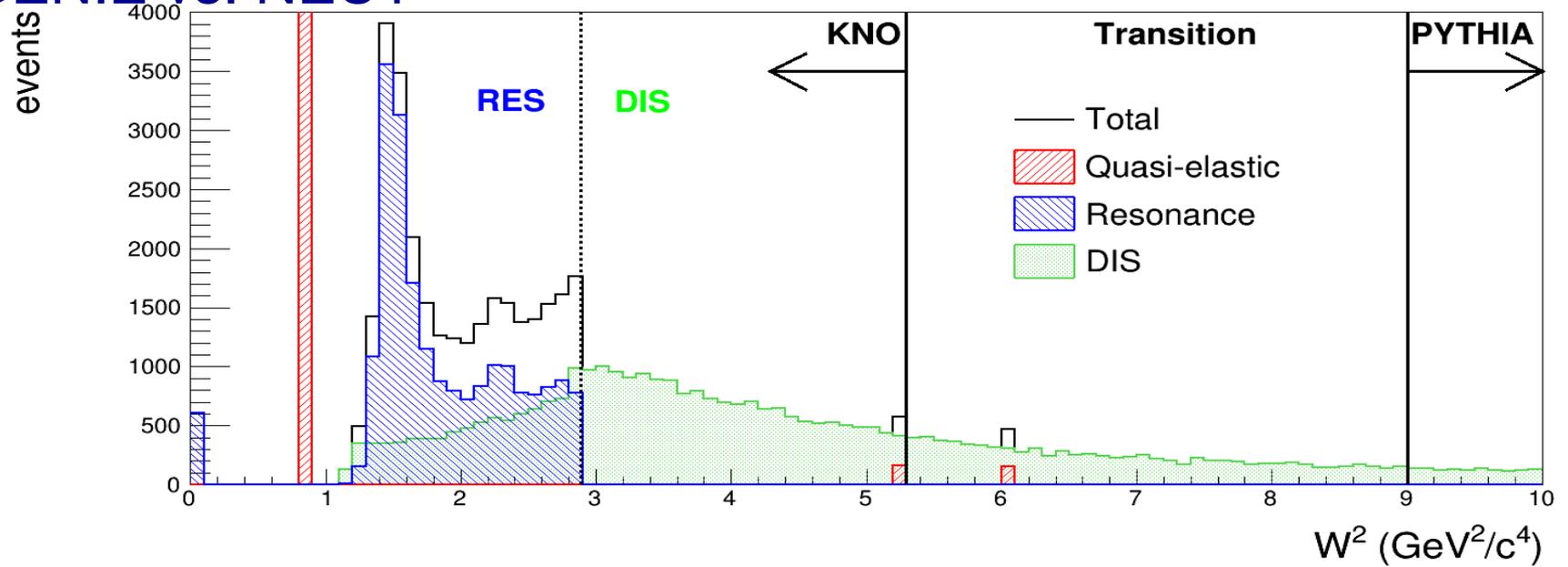
Christophe
Bronner
(IPMU)



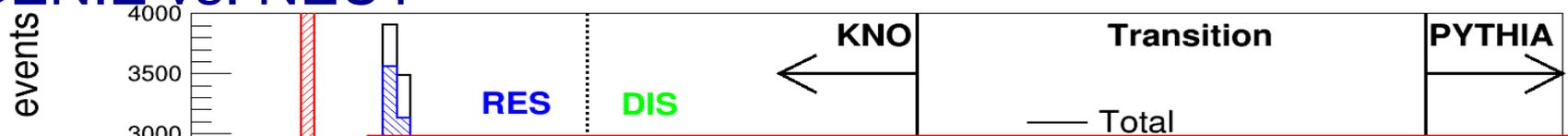
W^2 distribution for H_2O target with atmospheric- ν flux (NEUT)



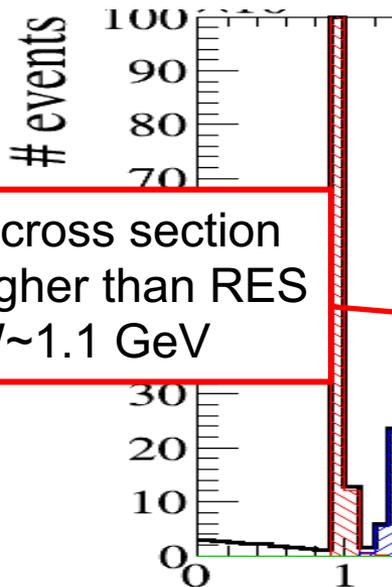
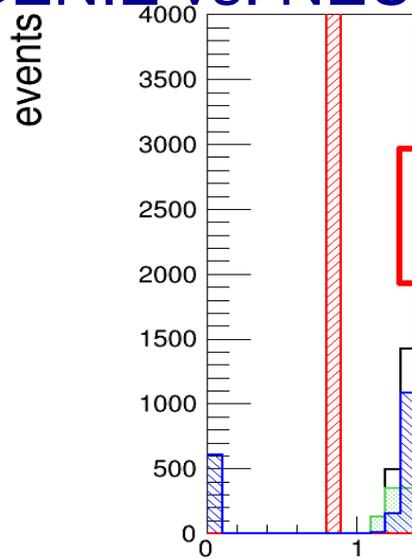
0. GENIE vs. NEUT



0. GENIE vs. NEUT



Fermilab wine and cheese seminar (Dec. 1)
 NOvA CC and NC π^0 production cross section measurement



DIS cross section
 is higher than RES
 at $W \sim 1.1$ GeV

JETP Seminar

First measurement of neutrino-induced production cross sections in the NOvA Near Detector

Duyang Hongyue
Daniel Pershey

Friday
December 1, 2017

Signal Composition

- Analysis signal has large contribution from both Res and D interactions
- There is a large multi- π component in ν_μ CC + π^0 events which included in the analysis signal

NOvA Simulation

$\nu_\mu + A \rightarrow \mu^- + \pi^0 + X$

NOvA π^0 Measurements

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H. Duyang & D. Pershey



0. DIS-hadronization error check list

- Goal is to make event weight with function of E_ν , x , y , etc, for IceCube oscillation program
- All errors are expected to be unimportant (?)

DIS or Hadronization	type of error	approach	size
DIS	Bodek-Yang correction	play with Bodek-Yang parameters (by eyes)	????
DIS	differential xs	NuTeV-GENIE comparison (bottom-up)	????
DIS	A-scaling	MINERvA-GENIE (bottom-up)	????
DIS	PDF	From nuclear PDF, CT10? nCTEQ? (top-down)	????
Hadronization	low W averaged charged hadron multiplicity	play with KNO parameters (by eyes)	????
Hadronization	high W averaged charged hadron multiplicity	bubble chamber-PYTHIA comparison (bottom-up)	????

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5. Low- W hadronization error

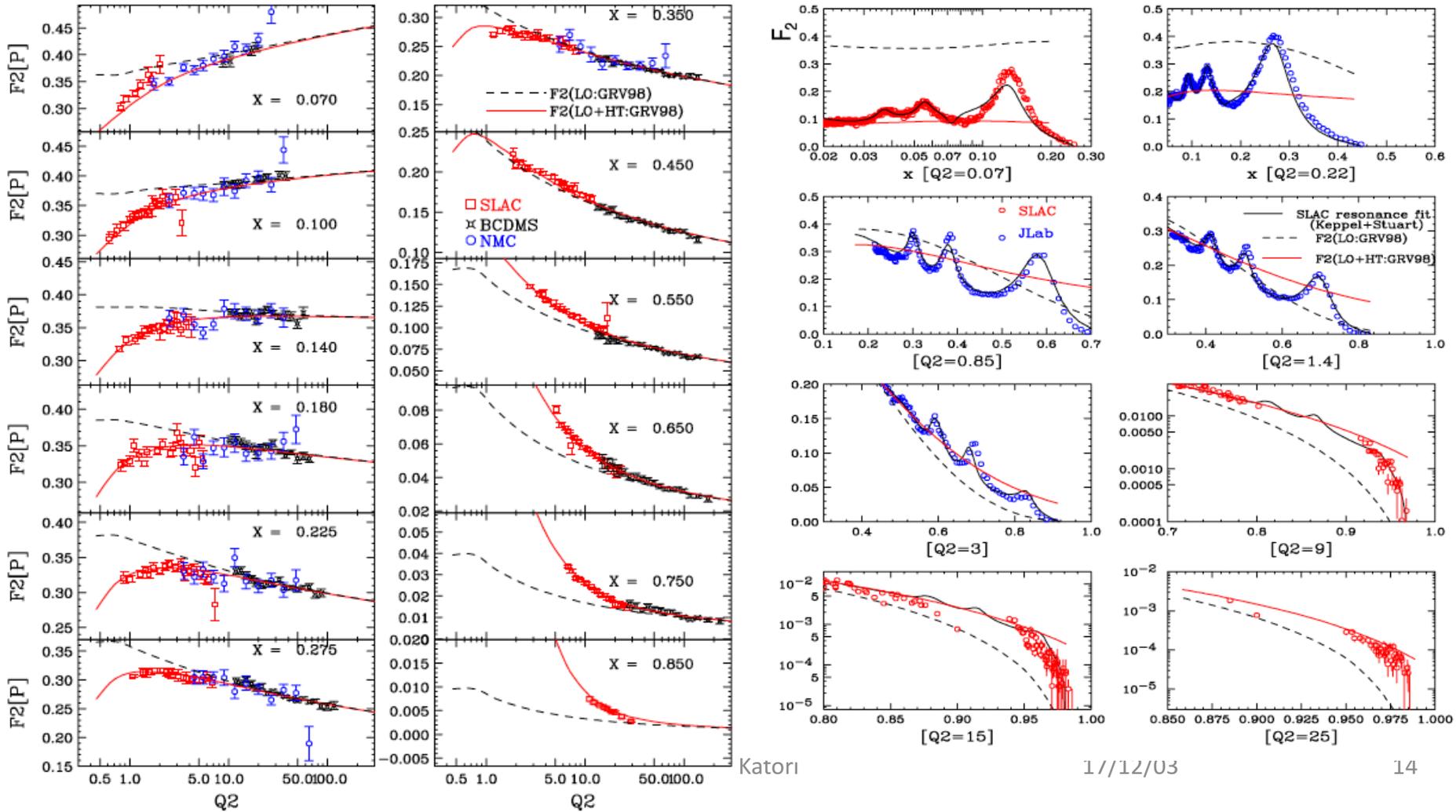
6. High- W hadronization error

7. Conclusion

1. Bodek-Yang correction for low Q^2 DIS

GRV98 is a PDF designed for low Q^2 region. Bodek-Yang correction makes GRV98 to work even lower Q^2 , or “duality” region by adding higher twist effect

Proton F_2 function GRV98-BY correction vs. data



Nachtmann variable $\xi = \frac{2x}{\left(1 + \sqrt{1 + \frac{4x^2 M^2}{Q^2}}\right)}$

1. Bodek-Yang correction for low Q² DIS

In GENIE, there are 11 parameters to control “Bodek-Yang correction” on GRV98 LO PDF

- A: high order twist correction
- B: quark transverse momentum
- Cvu1, Cvu2: valence u-quark PDF correction
- Cvd1, Cvd2: valence d-quark PDF correction
- Cs1u, Cs1d: sea u- and d-quark PDF correction
- x0, x1, x2: d(x)/u(x) correction

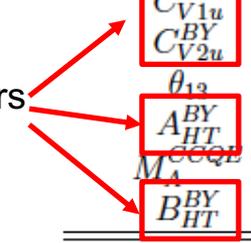
$$\xi \rightarrow \xi_\omega = \frac{2x \left(1 + \frac{M_f^2 + B}{Q^2}\right)}{\left(1 + \sqrt{1 + \frac{4x^2 M^2}{Q^2}}\right) + \frac{2Ax}{Q^2}}$$

$$K_{valence}(Q^2) = [1 - G_D^2(Q^2)] \cdot \left(\frac{Q^2 + C_{v2}}{Q^2 + C_{v1}}\right)$$

$$K_{sea}(Q^2) = \frac{Q^2}{Q^2 + C_{s1}}$$

parameter	impact (%)		
	1 year	3 year	5 year
hierarchy	100.0	100.0	100.0
Δm_{31}^2	38.8	37.9	37.6
Energy scale	21.2	21.4	21.7
A_{eff} scale	15.2	13.2	11.4
θ_{23}	3.4	4.8	5.7
ν_e/ν_μ ratio	0.5	1.7	2.6
$\nu_\mu/\bar{\nu}$ ratio	0.5	1.2	2.3
M_A^{RES}	1.2	2.0	1.7
C_{V1u}^{BY}	0.1	0.3	0.3
C_{V2u}^{BY}	0.0	0.0	0.2
θ_{12}	0.0	0.1	0.2
A_{HT}^{BY}	0.0	0.0	0.0
M_A^{CCQE}	0.0	0.0	0.0
B_{HT}^{BY}	0.0	0.0	0.0

DIS errors



PINGU Lol variations

Name	nominal value	uncertainty (%)
M_A^{CCQE}	0.99	-15, +25
M_A^{RES}	1.120	±20
A_{HT}^{BY}	0.538	±25
B_{HT}^{BY}	0.305	±25
C_{V1u}^{BY}	0.291	±30
C_{V2u}^{BY}	0.189	±30

1. Bodek-Yang correction errors

Parameter variations are defined

- errors A and B: I follow Joshua's choice
- errors on PDF correction: 30% for all
- errors on $d(x)/u(x)$: next page

Since no correlations of parameters are available, 9 BY-systematic study samples are made to maximize of parameter variation effects

BY-parameters	CV	error
A	0.538	$\pm 25\%$
B	0.305	$\pm 25\%$
CsU	0.363	$\pm 30\%$
CsD	0.621	$\pm 30\%$
Cv1U	0.291	$\pm 30\%$
Cv2U	0.189	$\pm 30\%$
Cv1D	0.202	$\pm 30\%$
Cv2D	0.255	$\pm 30\%$
X0	-0.00817	$+0.00817$
X1	0.0506	-0.0506
X2	0.0798	-0.0798

sample	sample
1	default
2	A+ δ A, B- δ B
3	A- δ A, B+ δ B
4	CsU+ δ CsU, CsD- δ CsD
5	CsU- δ CsU, CsD+ δ CsD
6	Cv1U+ δ Cv1U, Cv2U- δ Cv2U
7	Cv1U- δ Cv1U, Cv2U+ δ Cv2U
8	Cv1D+ δ Cv1D, Cv2D- δ Cv2D
9	Cv1D- δ Cv1D, Cv2D+ δ Cv2D
10	X0=0, X1=0, X2=0

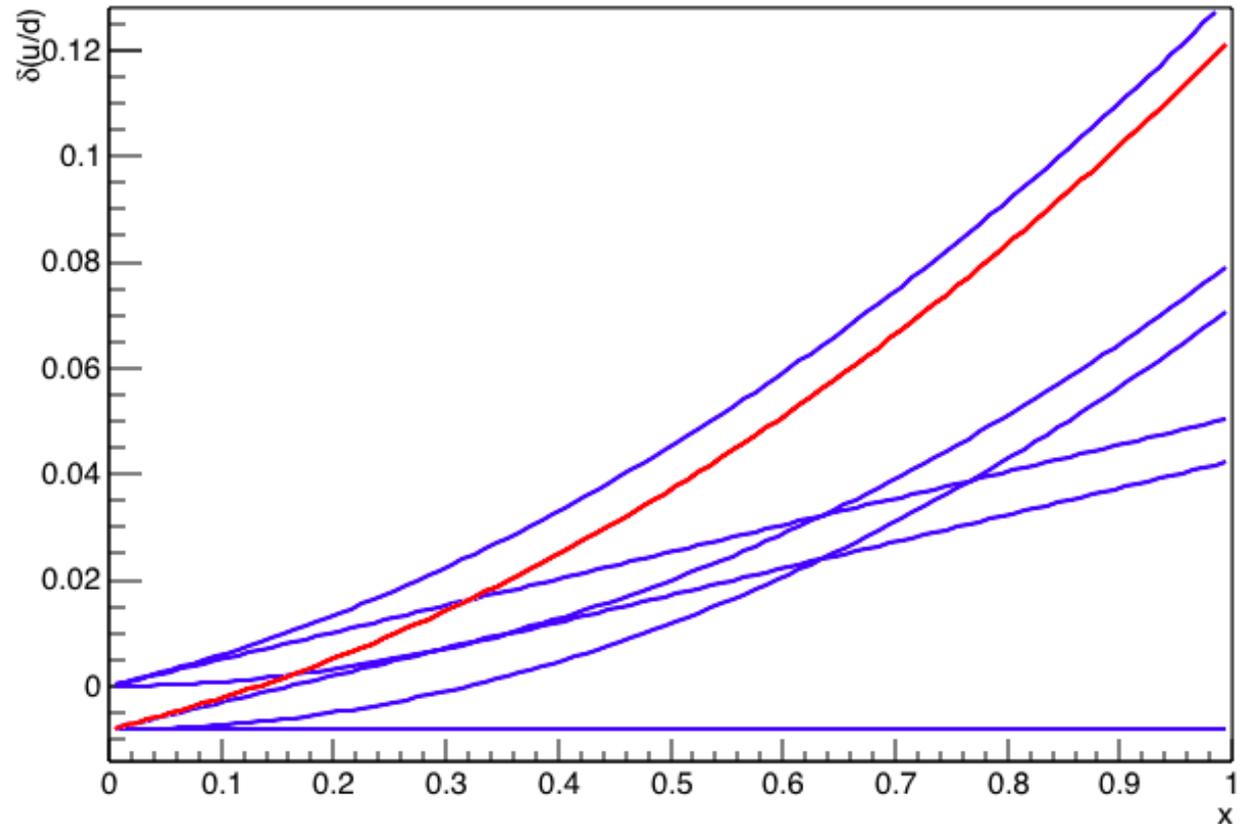


1. $d(x)/u(x)$ variation study

$$\delta(d(x)/u(x)) = X_0 + X_1 \cdot x + X_2 \cdot x^2$$

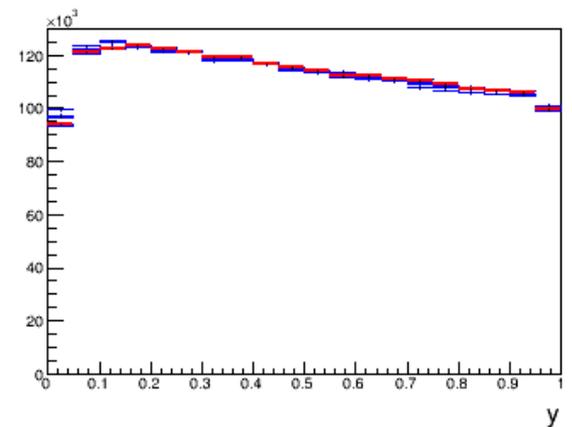
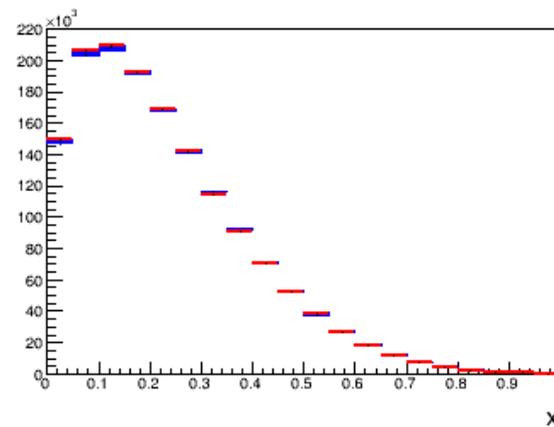
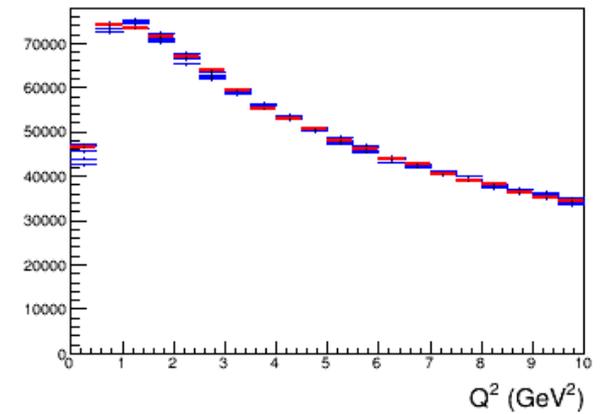
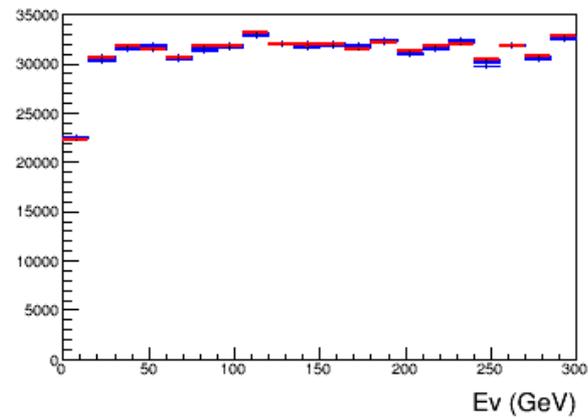
- 2nd order polynomial describe this error, ~10% effect at large x
- A reasonable choice of envelope is when the function is 0.

BY u/d ratio correction, $0.05 < x < 0.75$



1. Results

BY parameter variation make small variations in $E\nu$, Q^2 , x , y .

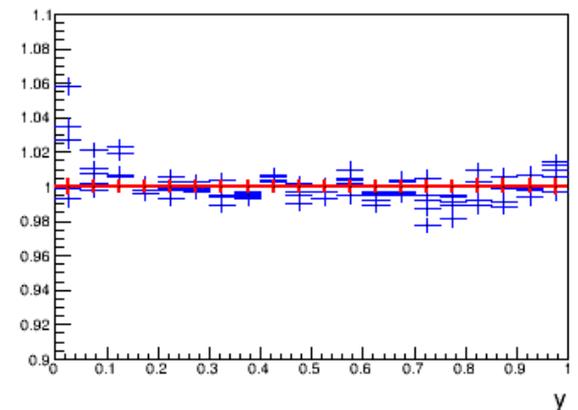
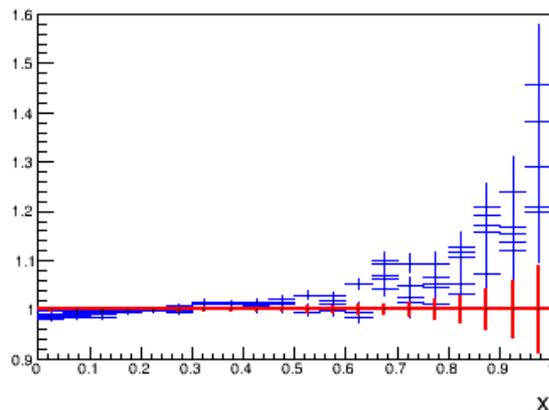
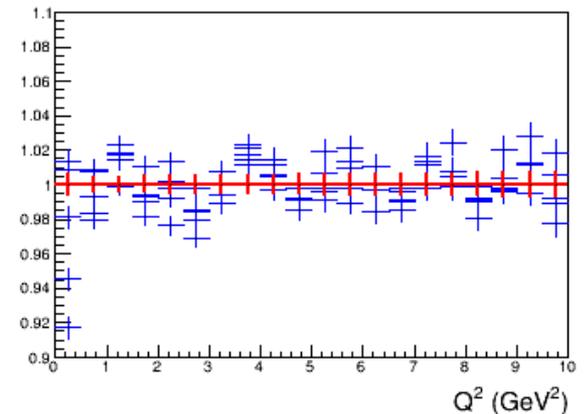
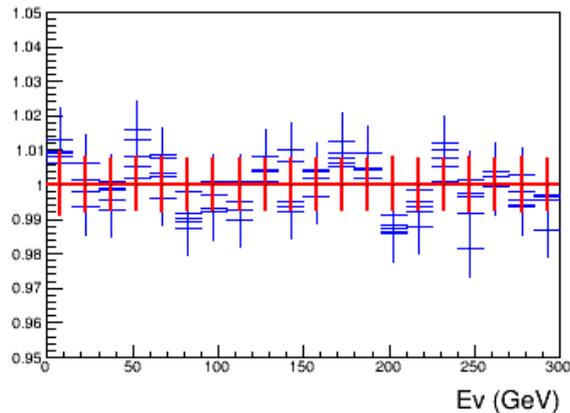


1. Results

BY parameter variation make small variations in $E\nu$, Q^2 , x , y .

- $E\nu$: <2% variation in all region
- Q^2 : ~8% variation at $Q^2=0.5 \text{ GeV}^2$
- x : ~50% variation at $x\sim 1$
- y : ~6% variation at $y\sim 0$

In general, variation can be large by assuming correlations on parameters



1. DIS Bodek-Yang correction error

- Goal is to make event weight with function of E_ν , x , y , etc, for IceCube oscillation program
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DIS	A-scaling	MINERvA-GENIE (bottom-up)	????
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3. DIS A-scaling error

4. DIS PDF error

5. Low-W hadronization error

6. High-W hadronization error

7. Conclusion



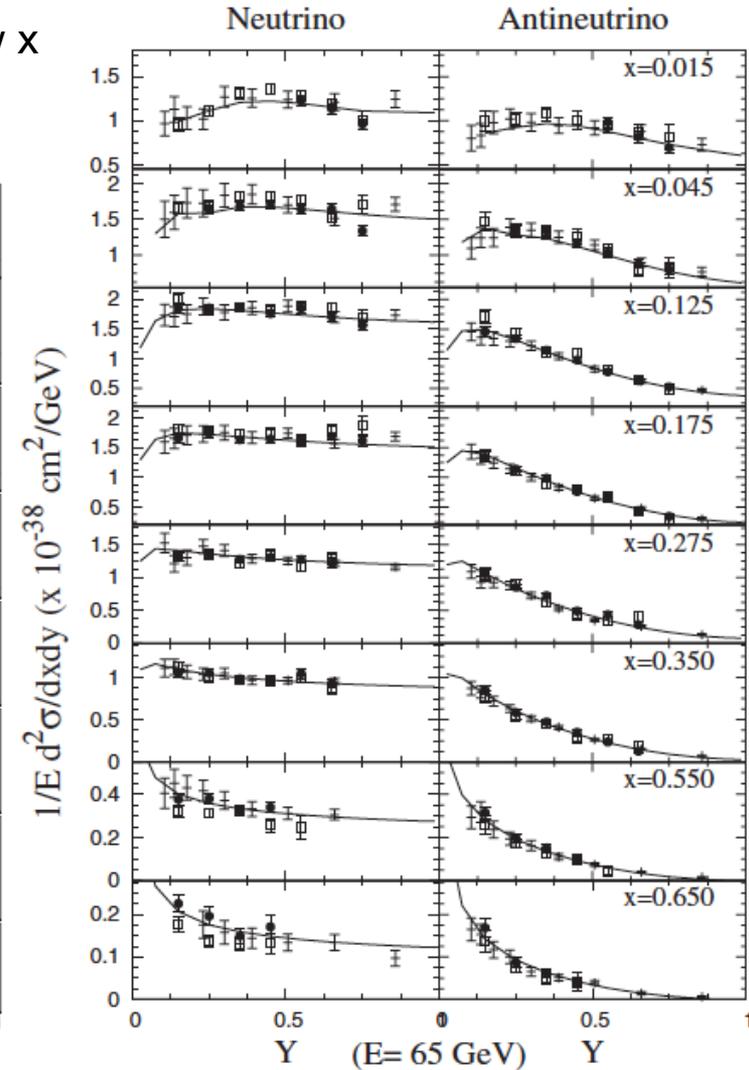
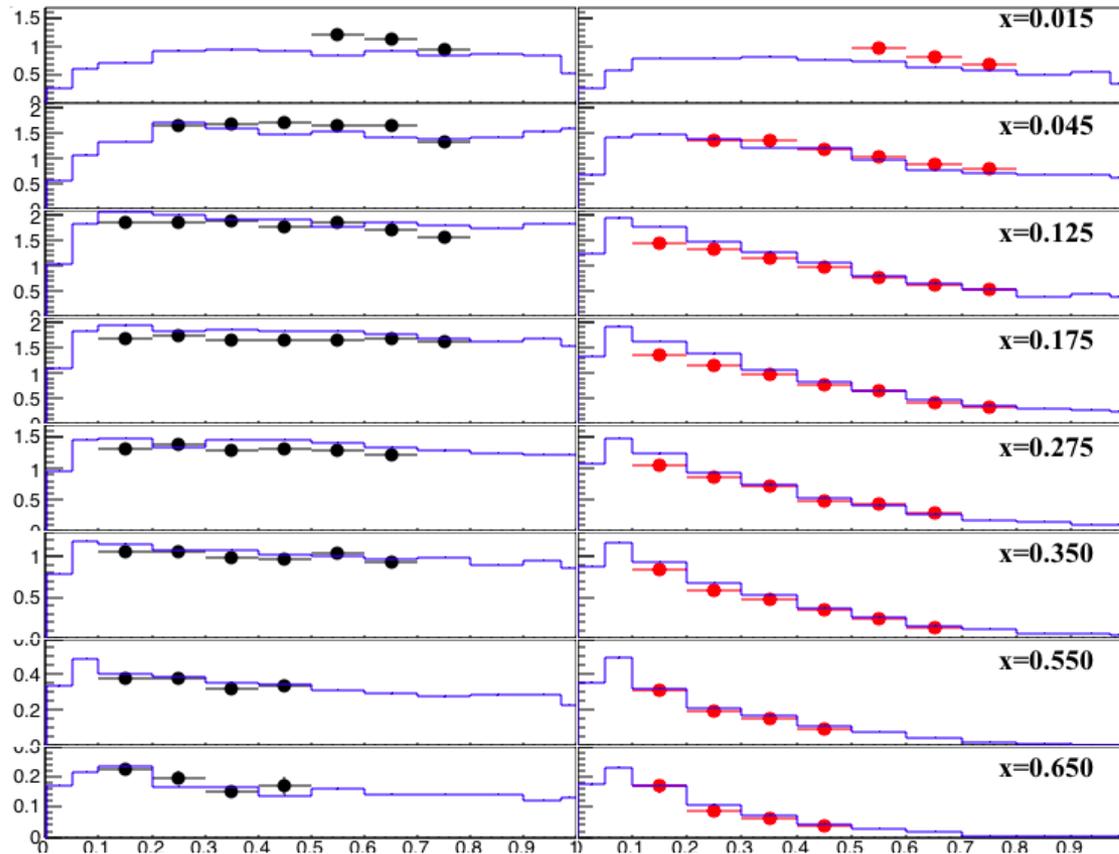
2. GENIE-NuTeV comparison

GENIE v2.10.6

Seems GENIE reproduce NuTeV data except very low x

65 GeV

NuTeV ν -Fe and anti- ν -Fe
differential cross section ($x, y, E\nu$)



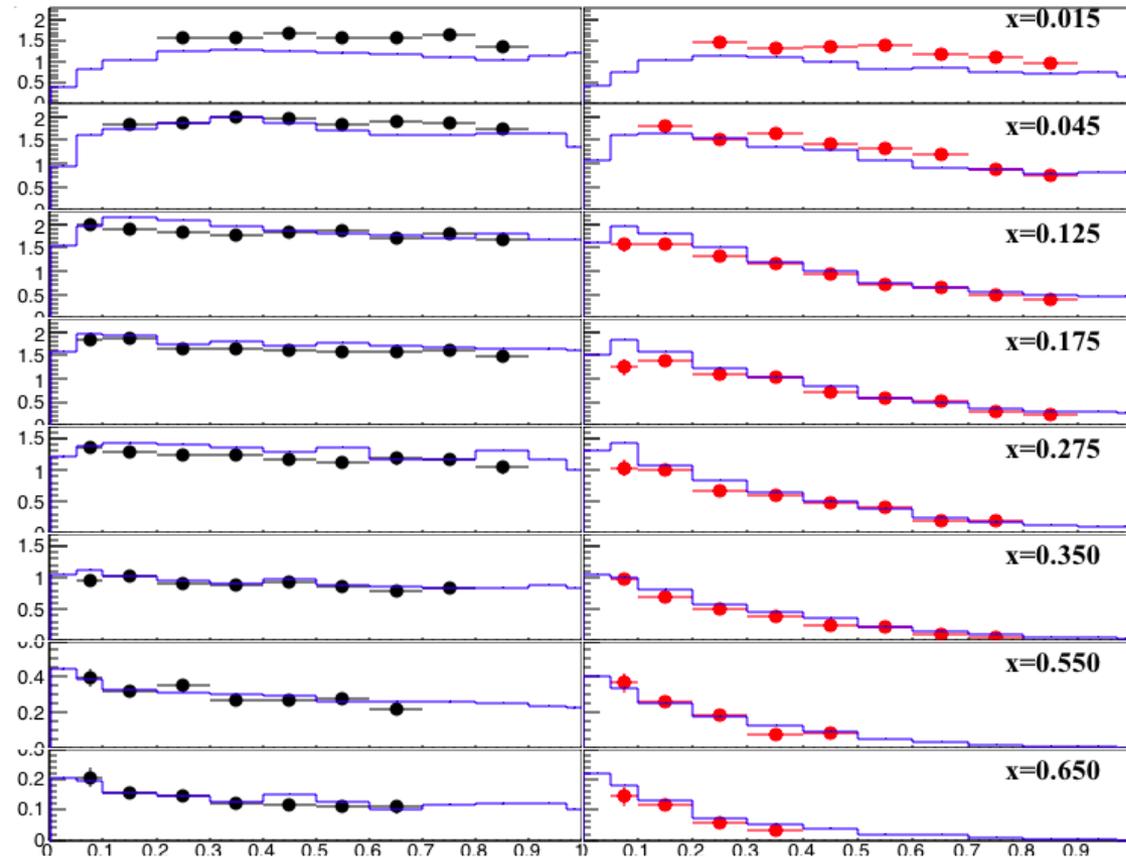


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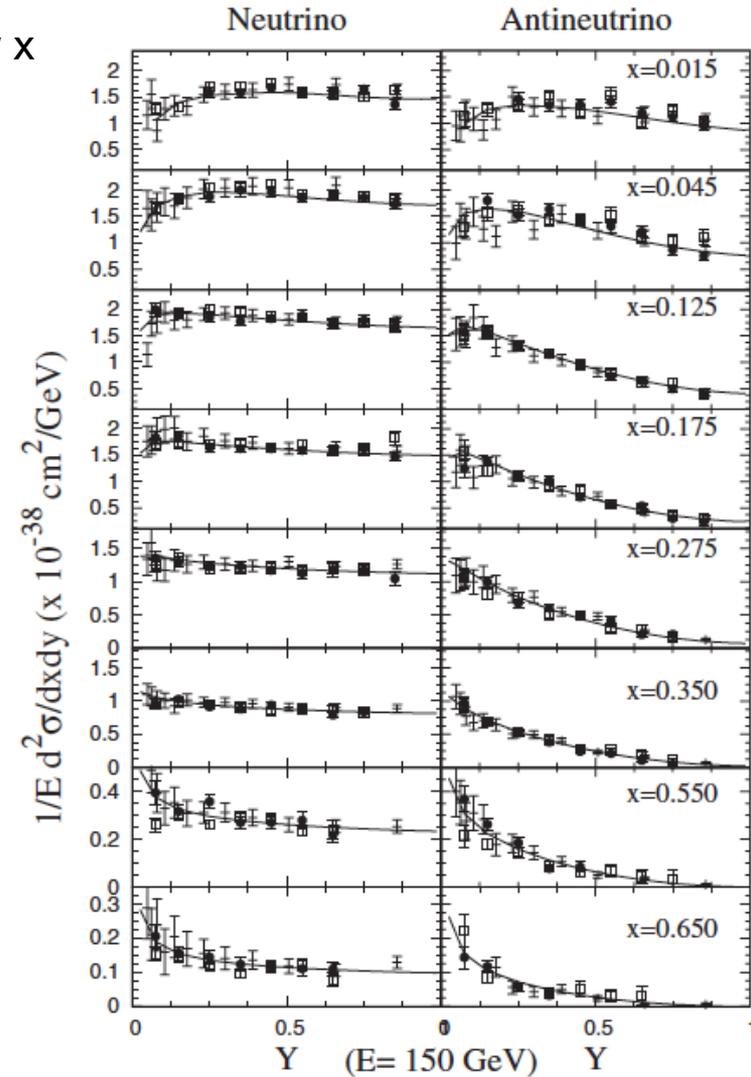
GENIE v2.10.6

Seems GENIE reproduce NuTeV data except very low x

150 GeV



NuTeV ν -Fe and anti- ν -Fe
differential cross section (x, y, E)



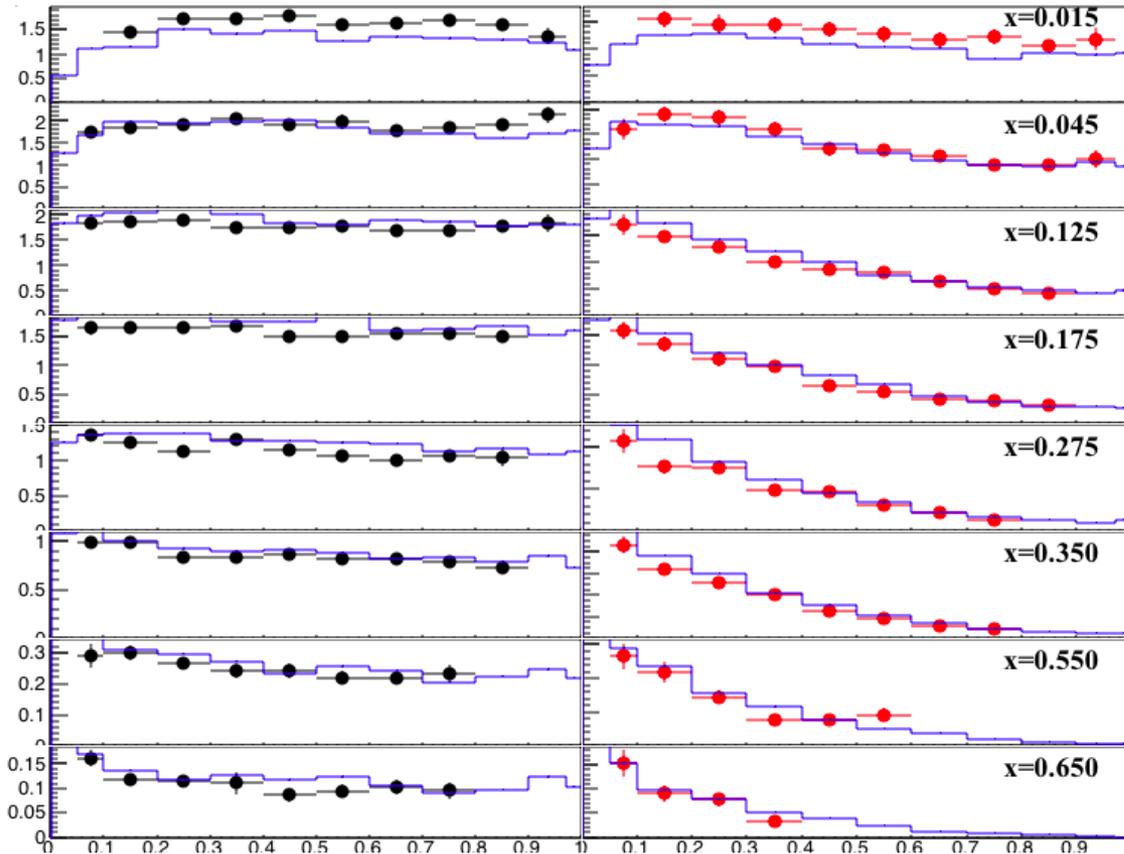


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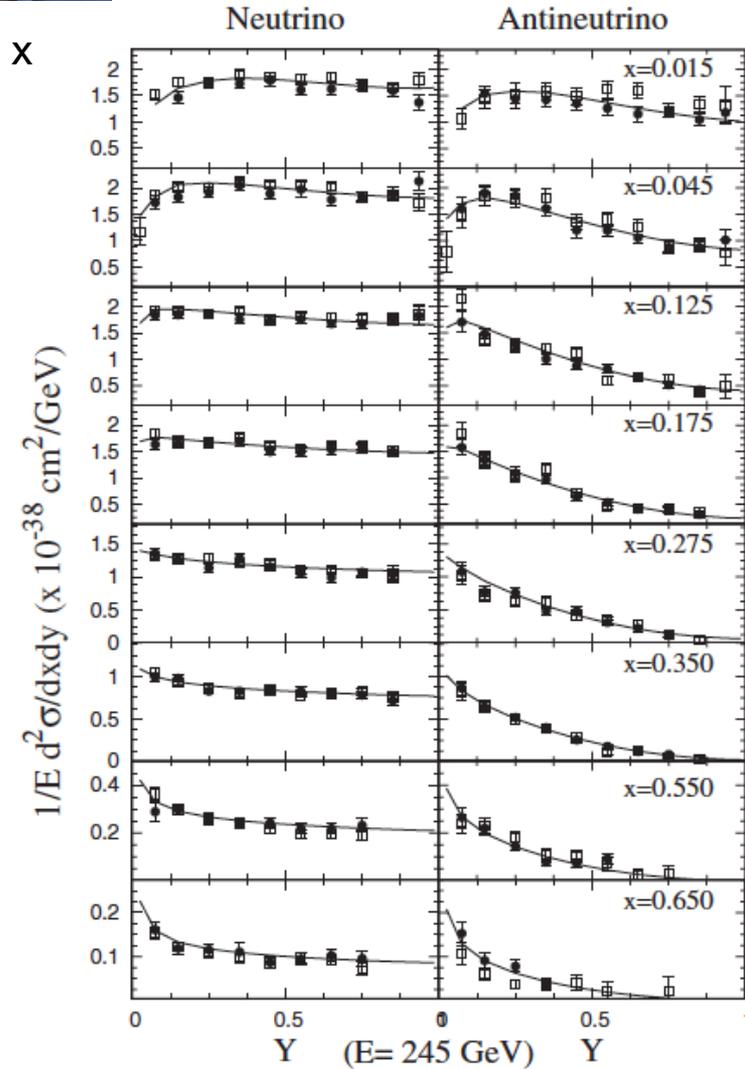
GENIE v2.10.6

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245 GeV



NuTeV ν -Fe and anti- ν -Fe
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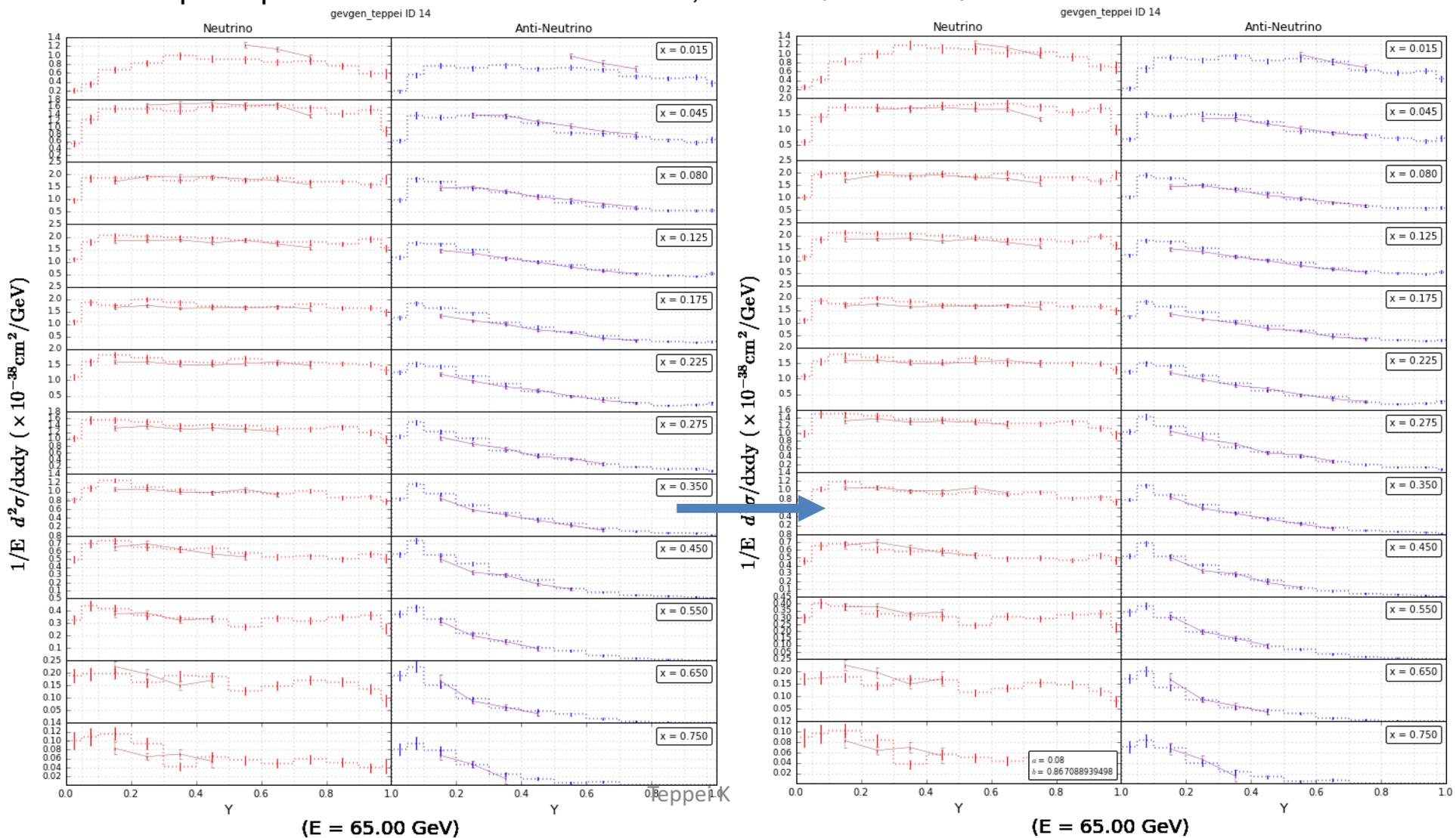


2. DIS differential cross section error

$$F(x, y) = bx^{-a}$$

GENIE-NuTeV comparison

- simple 2-parameter model with $a=0.08$, $b=0.87$ (for a trial)



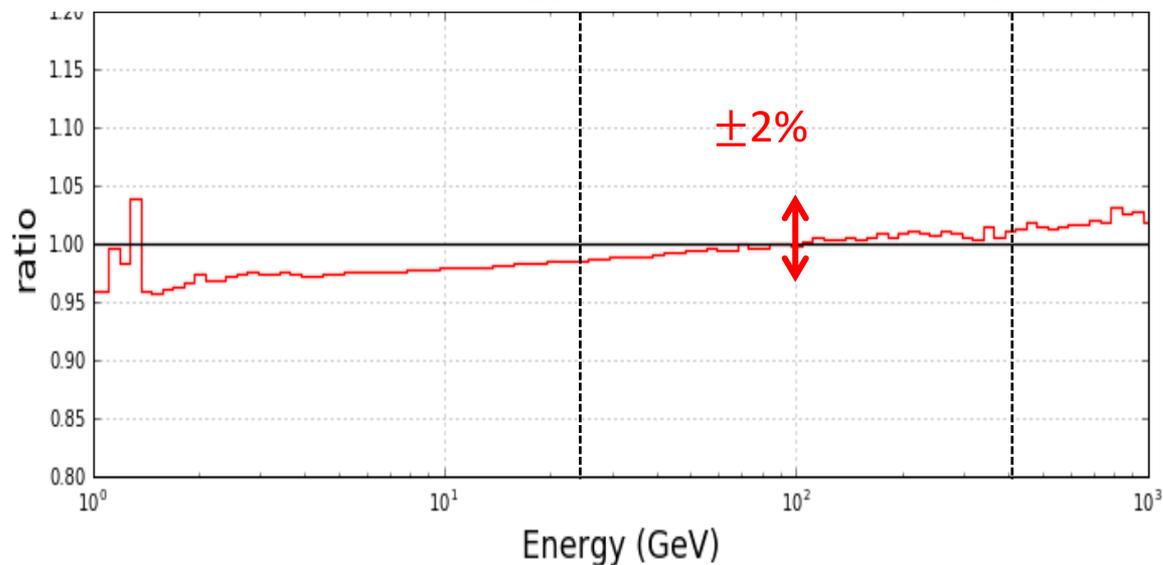
2. DIS differential cross section error

$$F(x, y) = bx^{-a}$$

GENIE-NuTeV comparison

- simple 2-parameter model with $a=0.08$, $b=0.87$ (for a trial)
- it has 2-3% shift of energy spectrum in 30-200 GeV
- However, the shift (\sim error) is larger than $\pm 2\%$ at $<10\text{GeV}$ and $>300\text{ GeV}$

Impact of low energy sample DIS re-weighting



2. DIS differential cross section error

- Goal is to make event weight with function of E_ν , x , y , etc, for IceCube oscillation program
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DIS	differential xs	NuTeV-GENIE comparison (bottom-up)	2-3% by GENIE study
DIS	A-scaling	MINERvA-GENIE (bottom-up)	????
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3. DIS A-dependent error



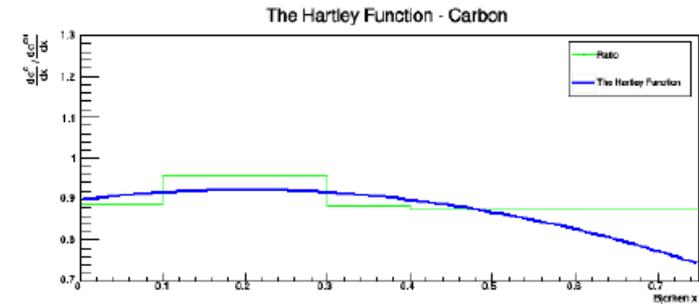
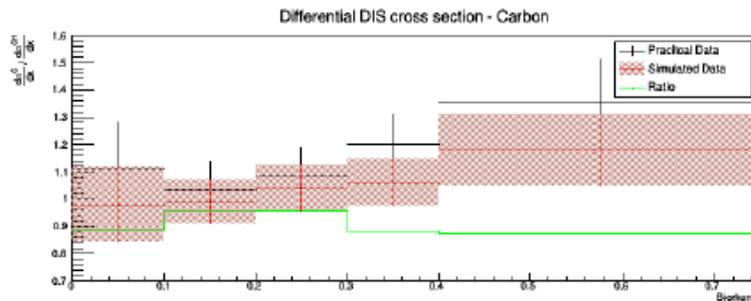
Liam Hartley
(Queen Mary)

GENIE-MINERvA comparison

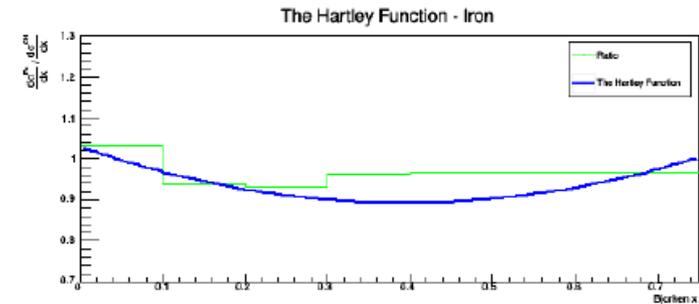
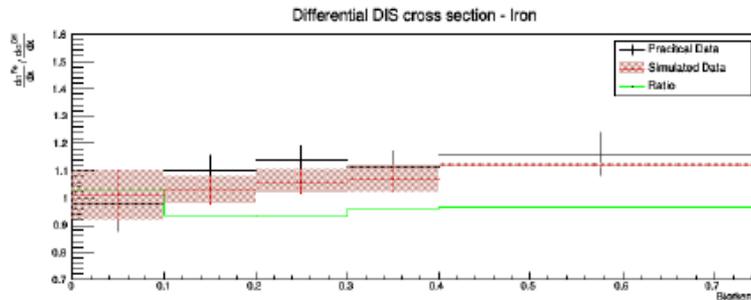
- Make a polynomial scaling function in A from data-MC ratio.
- Weight GENIE with function of x
- Bottom-up A-dependent DIS correction in x

$$\frac{d\sigma^A}{dx} / \frac{d\sigma^{CH}}{dx} = \frac{10A}{(-0.0084A^2 + 9.9A + 16)} + \frac{0.95(15 - A)}{A}x + \frac{0.95(A - 13.25)}{(A - 10)}x^2$$

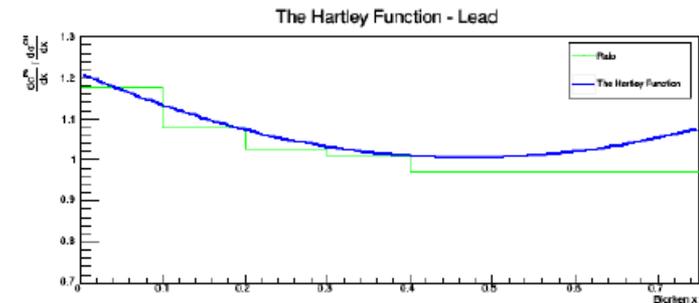
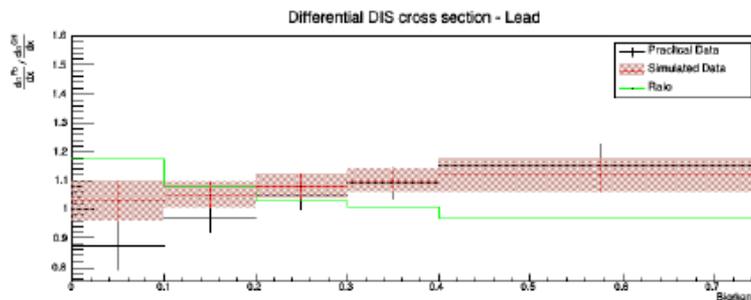
v-C



v-Fe



v-Pb

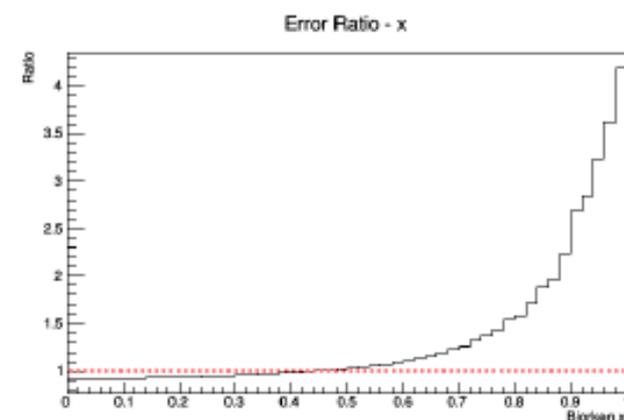
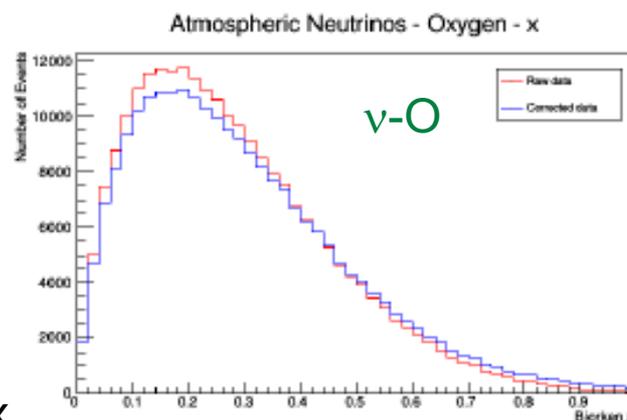


3. DIS A-dependent error

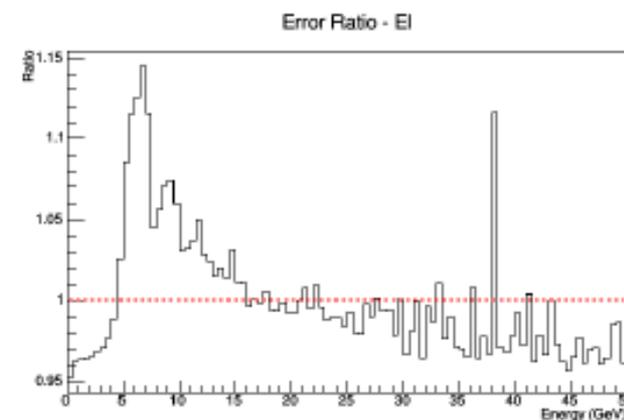
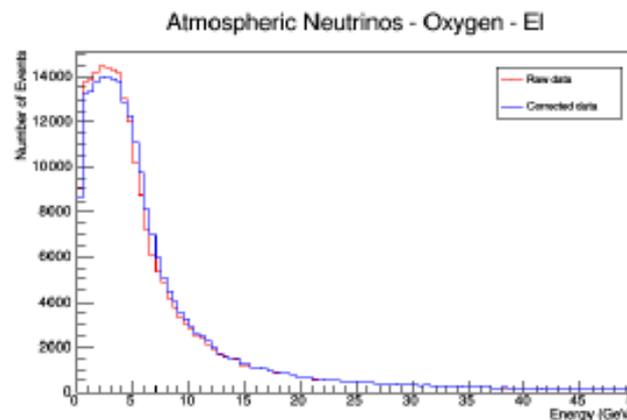
GENIE-MINERvA comparison

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- Weight GENIE with function of x
- Bottom-up A-dependent DIS correction in x
- Make prediction of correction in any targets, for example oxygen

$$\frac{d\sigma^A}{dx} / \frac{d\sigma^{CH}}{dx} = \frac{10A}{(-0.0084A^2 + 9.9A + 16)} + \frac{0.95(15 - A)}{A}x + \frac{0.95(A - 13.25)}{(A - 10)}x^2$$



Reasonably large variation (~10-20%) in x (under investigation)



3. DIS A-dependent error

- Goal is to make event weight with function of E_ν , x , y , etc, for IceCube oscillation program
- All errors are expected to be unimportant (?)

DIS or Hadronization	type of error	approach	size
DIS	Bodek-Yang correction	play with Bodek-Yang parameters (by eyes)	maybe large?
DIS	differential xs	NuTeV-GENIE comparison (bottom-up)	2-3% by GENIE study
DIS	A-scaling	MINERvA-GENIE (bottom-up)	maybe large?
DIS	PDF	From nuclear PDF, CT10? nCTEQ? (top-down)	????
Hadronization	low W averaged charged hadron multiplicity	play with KNO parameters (by eyes)	????
Hadronization	high W averaged charged hadron multiplicity	bubble chamber-PYTHIA comparison (bottom-up)	????

0. Introduction

1. DIS Bodek-Yang parameterization

2. DIS differential cross section error

3. DIS A-scaling error

4. DIS PDF error

5. Low-W hadronization error

6. High-W hadronization error

7. Conclusion



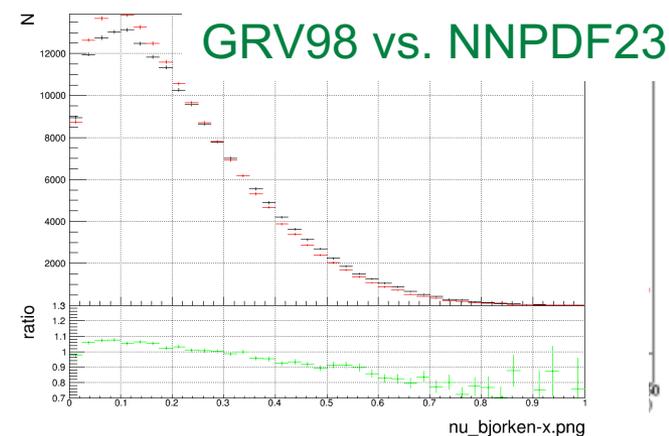
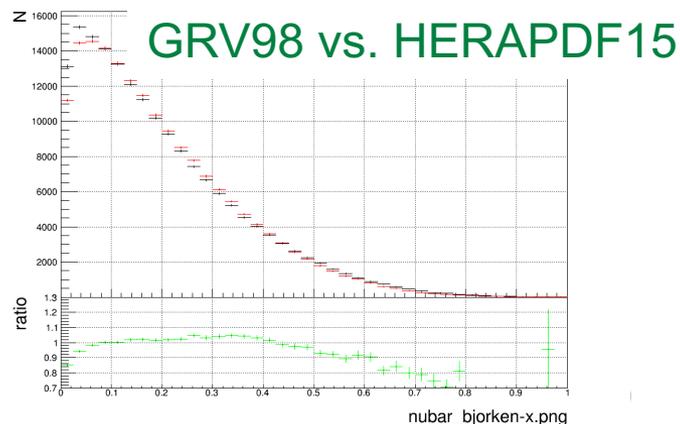
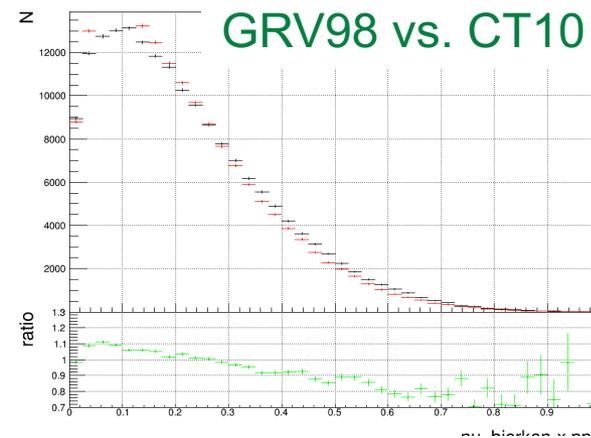
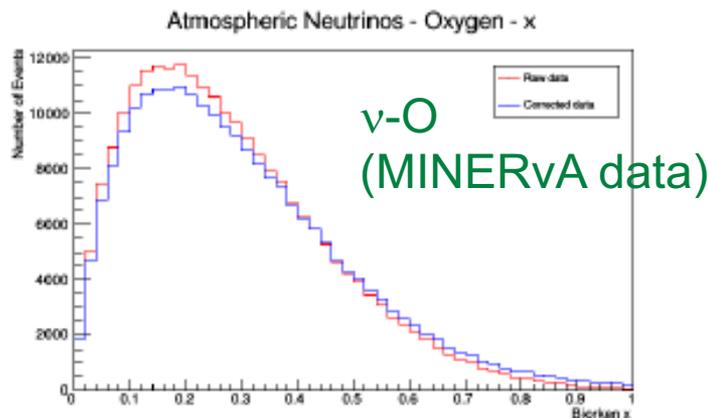
4. DIS PDF error

We tried to use couple of PDF from LHA PDF

- CT10 (NLO)
- HERAPDF15 (NLO)
- NNPDF23 (NLO)

As expected(?), PDF variation (top-down error) is smaller and well-controlled.

→ If we use a better PDF, variation would be few %



4. DIS PDF error

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6. High- W hadronization error

7. Conclusion



5. Low-W hadronization model

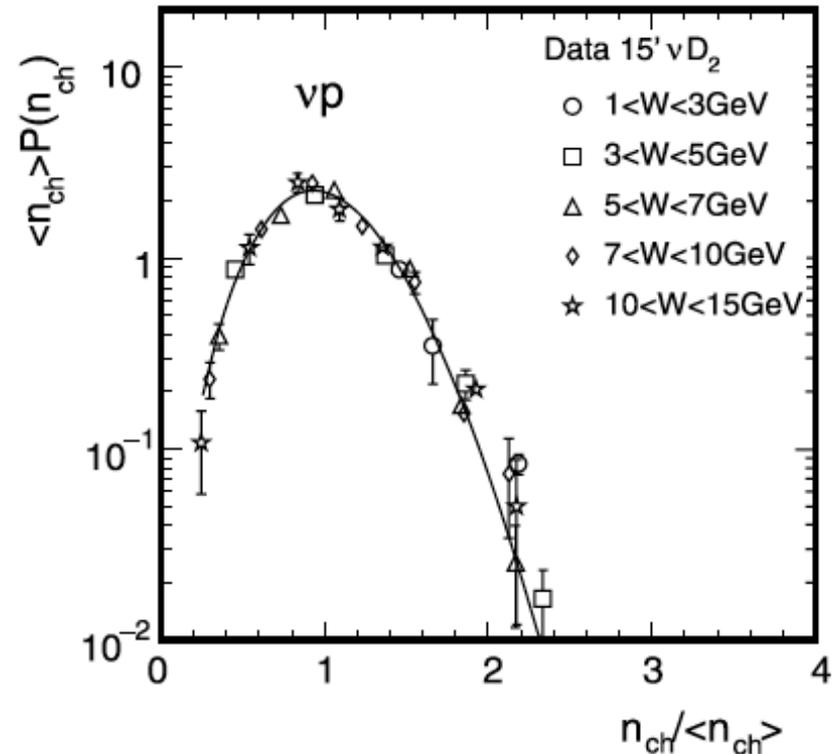
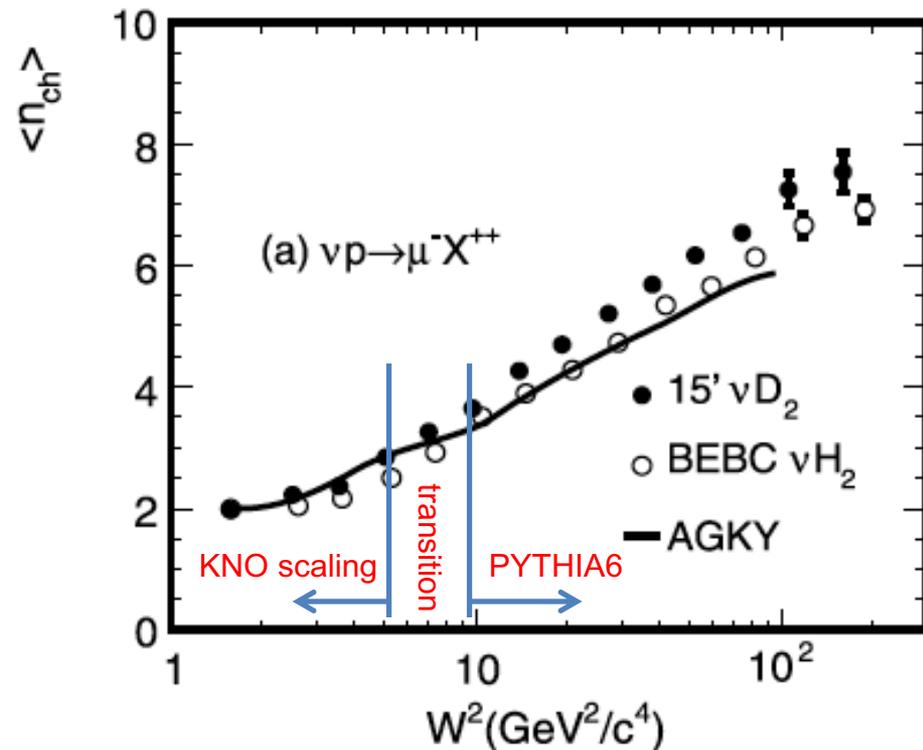
In AGKY model, hadronization model is a combination of 2 models.

KNO-scaling based model (low W hadronization)

- Data-driven model (agree with bubble chamber data, by construction)
- Averaged charged hadron multiplicity $\langle n_{ch} \rangle$ is chosen from data, with empirical function
- Averaged neutral hadron multiplicity is chosen from isospin.
- Then variance of multiplicity is chosen from KNO-scaling law.

$$\langle n_{ch} \rangle = a_{ch} + b_{ch} \cdot \ln(W^2)$$

$$\langle n \rangle \cdot P(n) = \frac{2e^{-c} c^{cn/\langle n \rangle + 1}}{\Gamma(cn/\langle n \rangle + 1)}$$



5. Low- W hadronization model

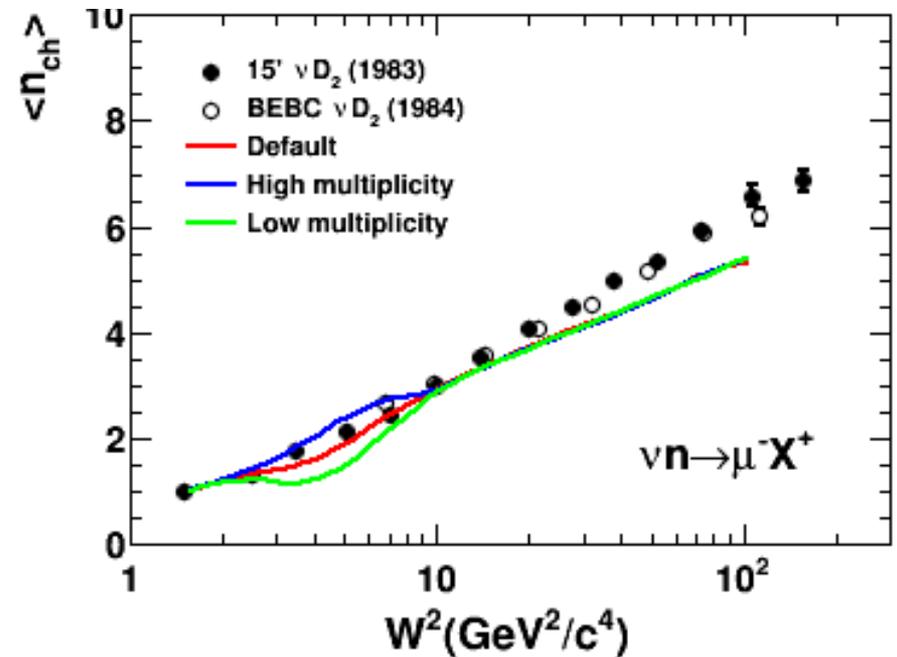
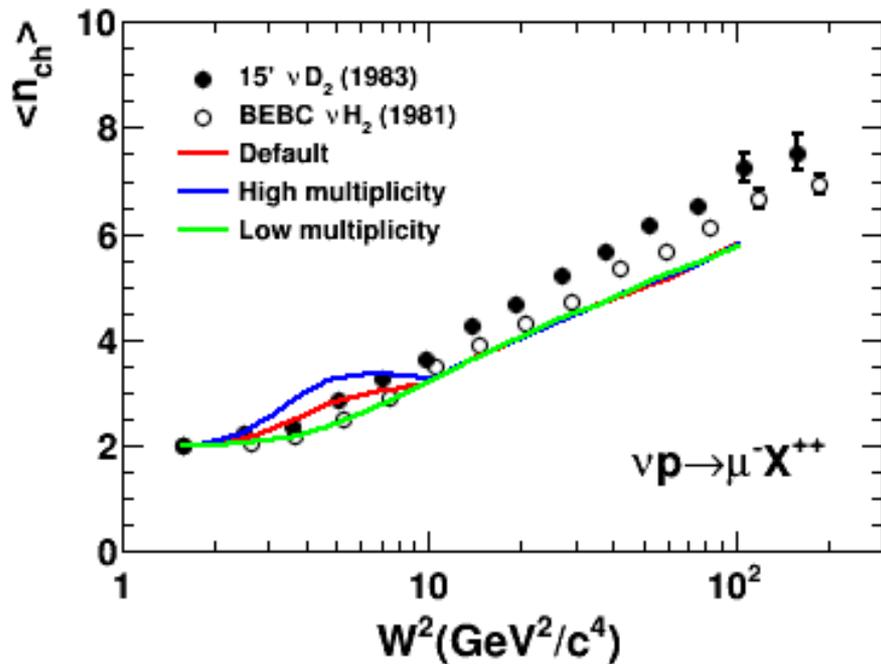
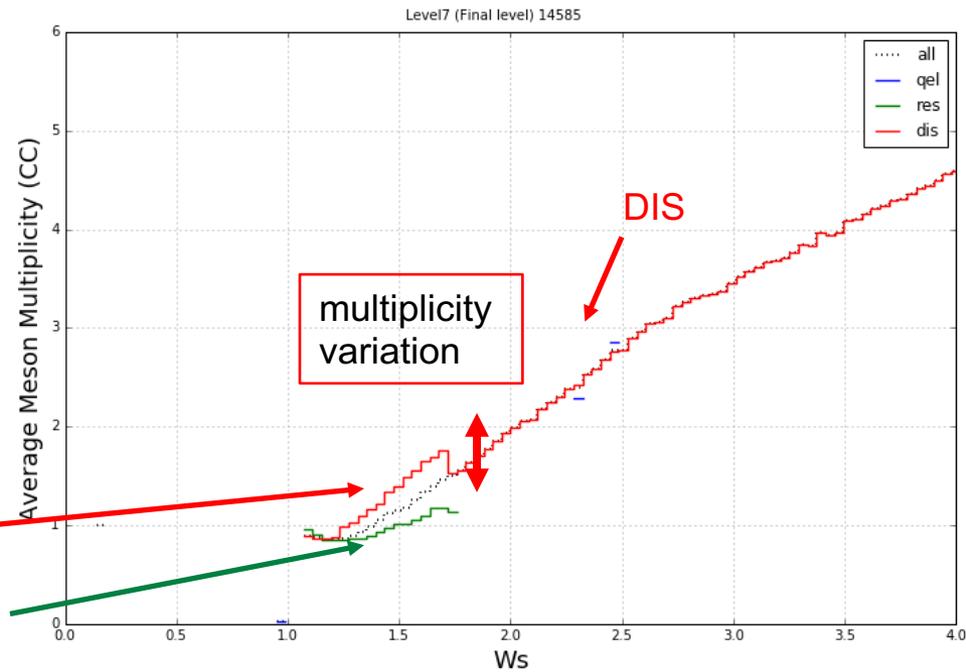
I added $\pm 25\%$ variation on “ b_{ch} ” parameters
 → This imitate discontinuity of hadron multiplicity

It looks to cover existing data variation

Then translate this variation in terms of hadron visible energy

$$y^{\text{eff}} = \frac{E_h^{\text{vis}}}{E_h^{\text{vis}} + E_\mu} \cdot E_h^{\text{vis}} = \sum_{E_h > E_{th}^i} T_h^i + \sum E_\gamma^i$$

non-resonance background
resonance contribution



5. Low-W hadronization model error

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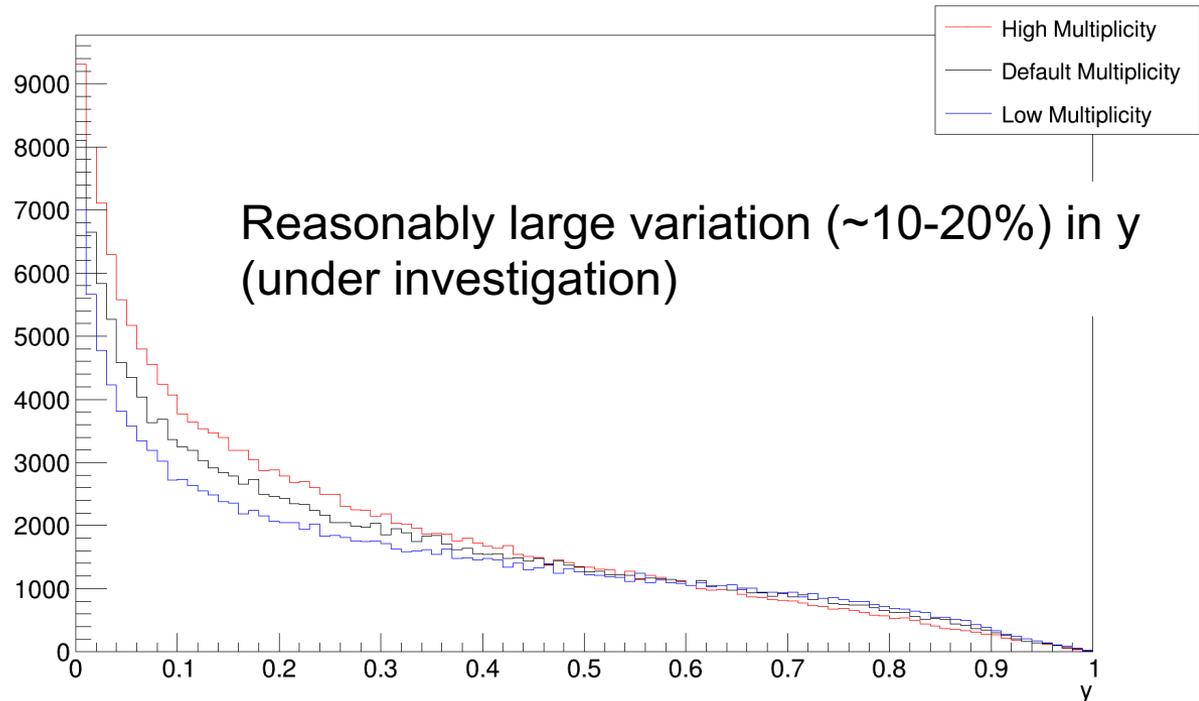
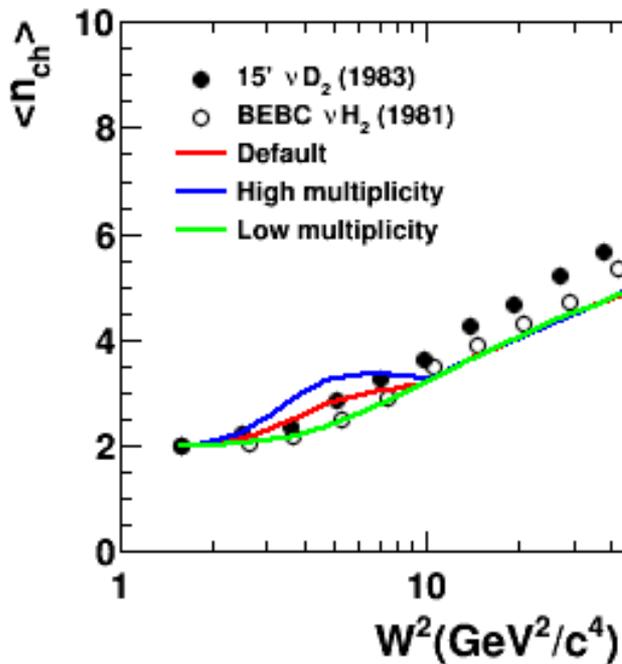
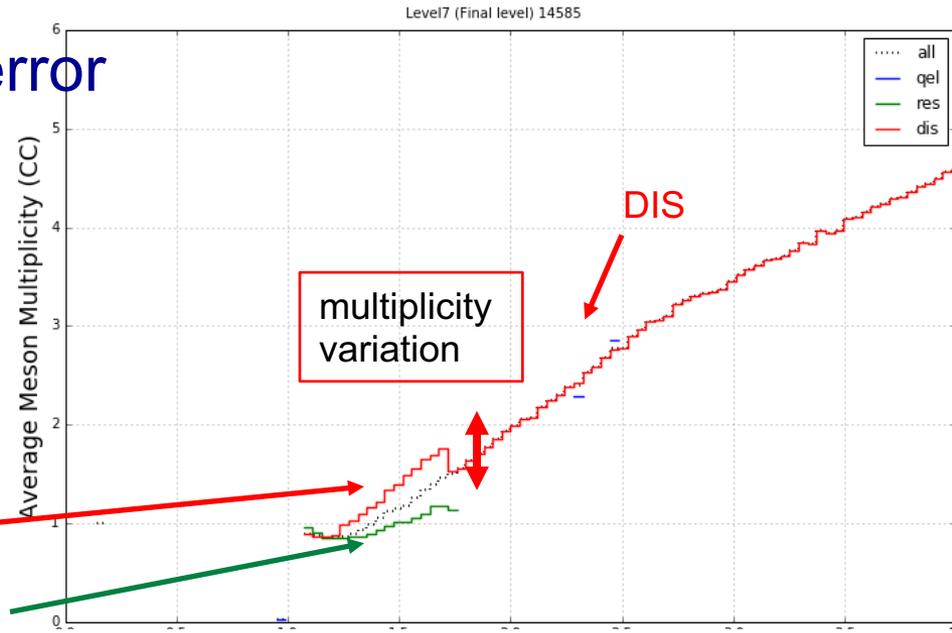
It looks to cover existing data variation

Then translate this variation in terms of hadron visible energy

$$y^{\text{eff}} = \frac{E_h^{\text{vis}}}{E_h^{\text{vis}} + E_\mu} \cdot E_h^{\text{vis}} = \sum_{E_h^i > E_{th}^i} T_h^i + \sum E_\gamma^i$$

non-resonance background

resonance contribution



5. Low-W hadronization error

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Hadronization	high W averaged charged hadron multiplicity	bubble chamber-PYTHIA comparison (bottom-up)	????

0. Introduction

1. DIS Bodek-Yang parameterization

2. DIS differential cross section error

3. DIS A-scaling error

4. DIS PDF error

5. Low- W hadronization error

6. High- W hadronization error

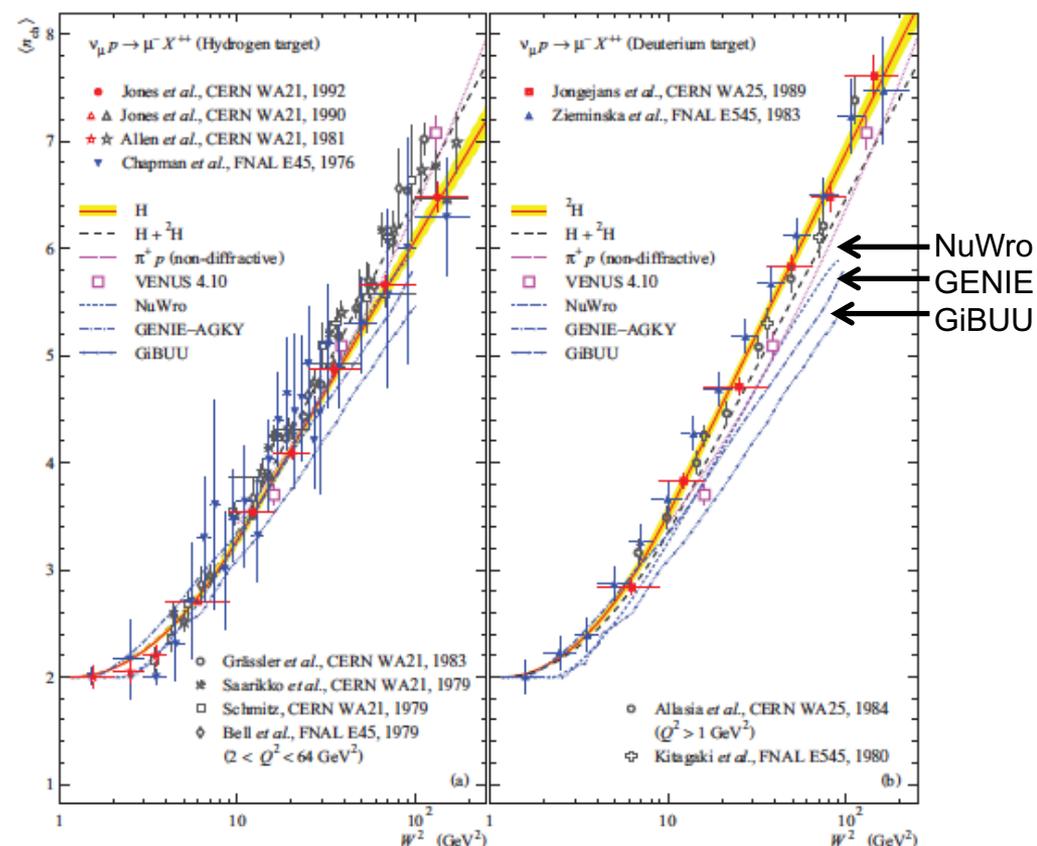
7. Conclusion

6. High- W hadronization model

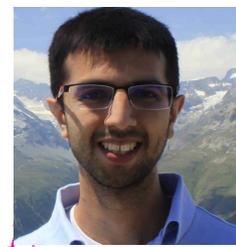
Kuzmin-Naumov fit

- They systematically analysed all bubble chamber data
- Difference of hydrogen and deuterium data
- Presence of kinematic cuts
- Better parameterization

All PYTHIA-based models underestimate averaged charged hadron multiplicity data (GiBUU, GENIE, NuWro, NEUT)



Average charged hadron multiplicity with function of W^2



6. High-W hadronization model

Averaged charged hadron multiplicity $\langle n_{\text{ch}} \rangle$

- PYTHIA6 with tuned Lund string function can reproduce $\langle n_{\text{ch}} \rangle$ data both neutrino and antineutrino.

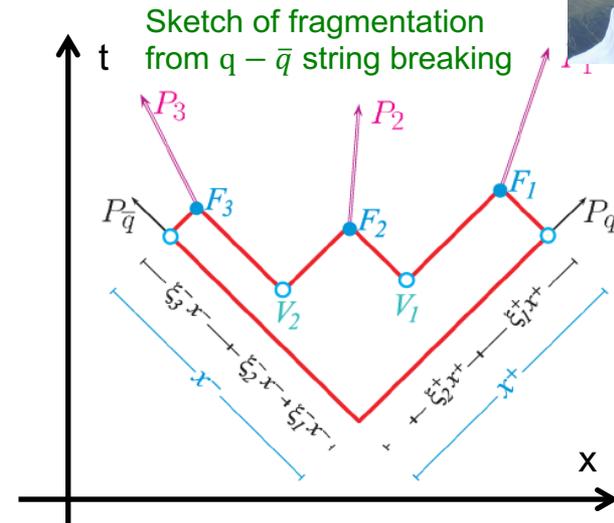
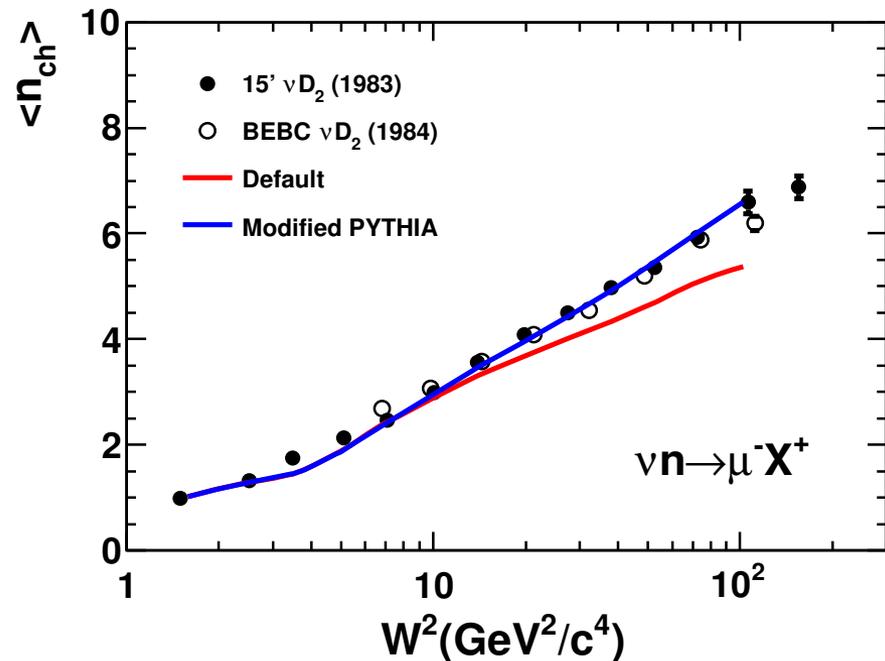
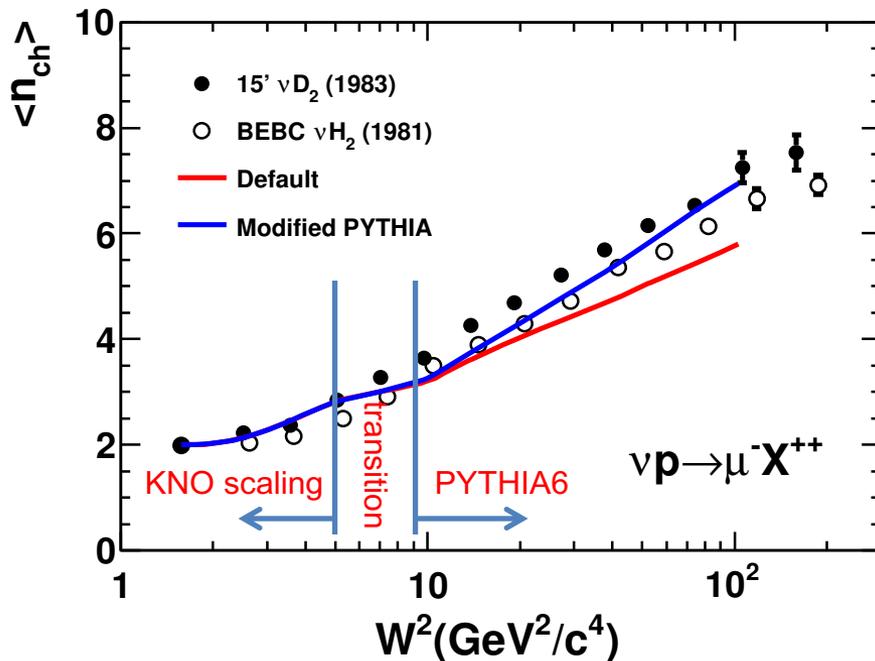
Lund string
function

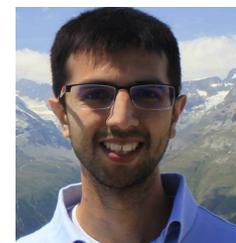
hadron energy distribution
from iterative process

tunnelling probability

$$f(z) \propto z^{-1} (1-z)^a \cdot \exp\left(\frac{-bm_{\perp}^2}{z}\right)$$

Neutrino average charged hadron multiplicity





6. High- W hadronization model error

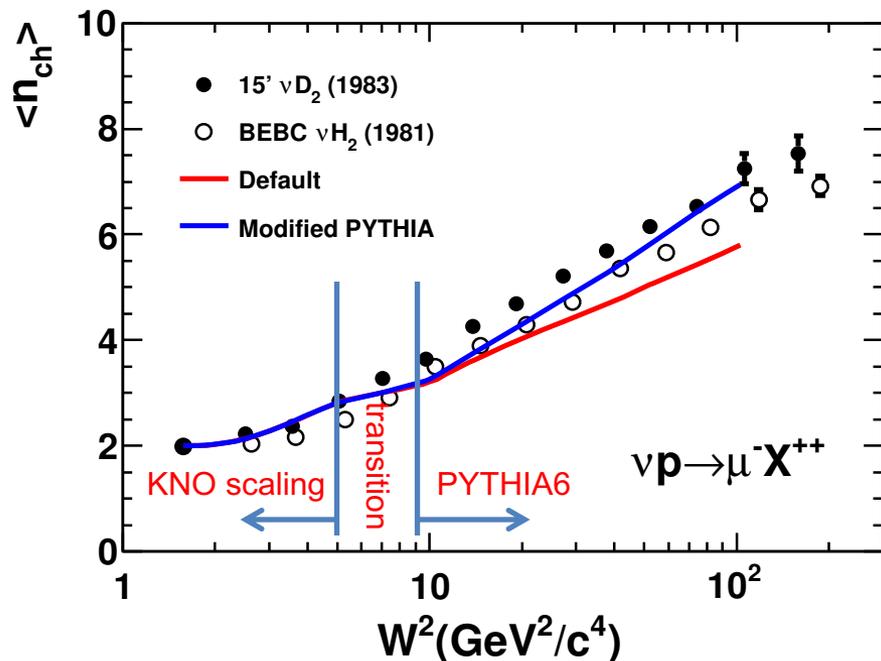
Averaged charged hadron multiplicity $\langle n_{\text{ch}} \rangle$

- PYTHIA6 with tuned Lund string function can reproduce $\langle n_{\text{ch}} \rangle$ data both neutrino and antineutrino.

Hadronization error propagation

- Difference of averaged charged hadron multiplicity is translated to visible hadron energy, then effective inelasticity. This is applied to variation of inelasticity error in simulation. Impact of hadronization error is small for experiments which only measure hadron shower

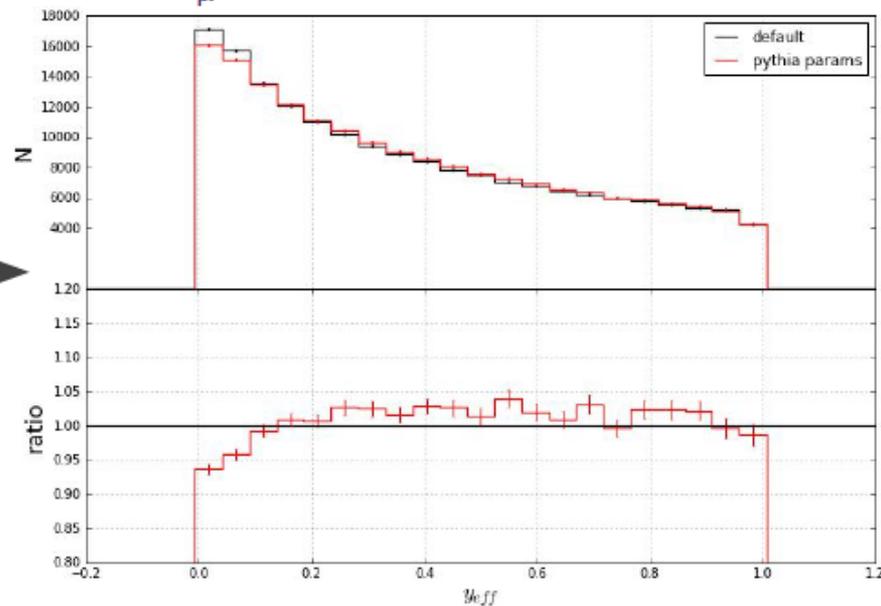
Neutrino average charged hadron multiplicity

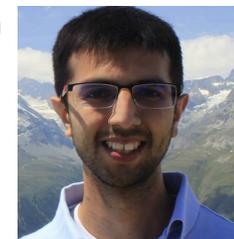


$$E_h^{\text{vis}} = \sum_{E_k^i > E_h^i} T_h^i + \sum E_\gamma^i$$

$$y^{\text{eff}} = \frac{E_h^{\text{vis}}}{E_h^{\text{vis}} + E_\mu}$$

ν_μ CC DIS GENIE v2.10.0





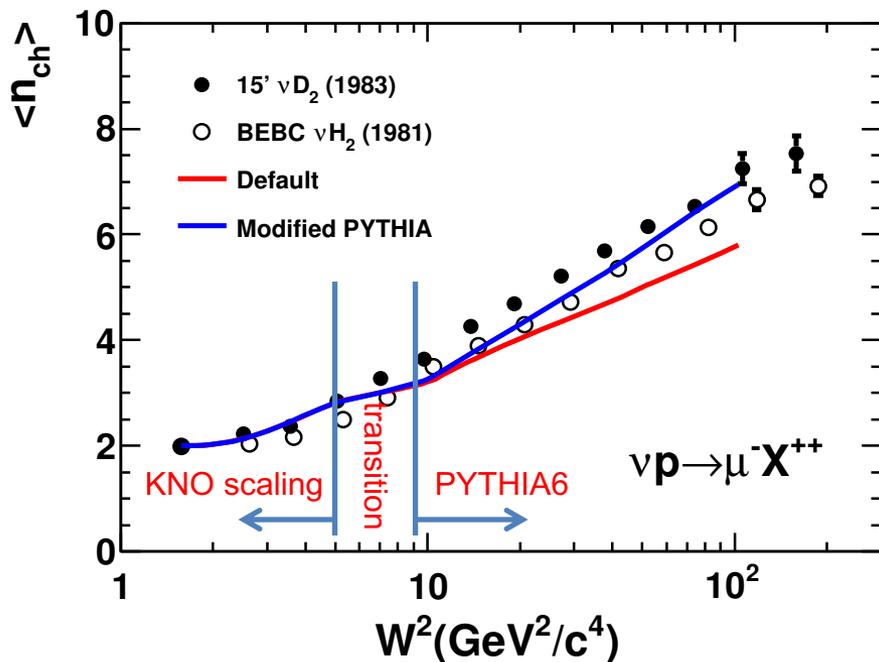
6. High-W hadronization dispersion error?

Bubble chamber topological cross section data

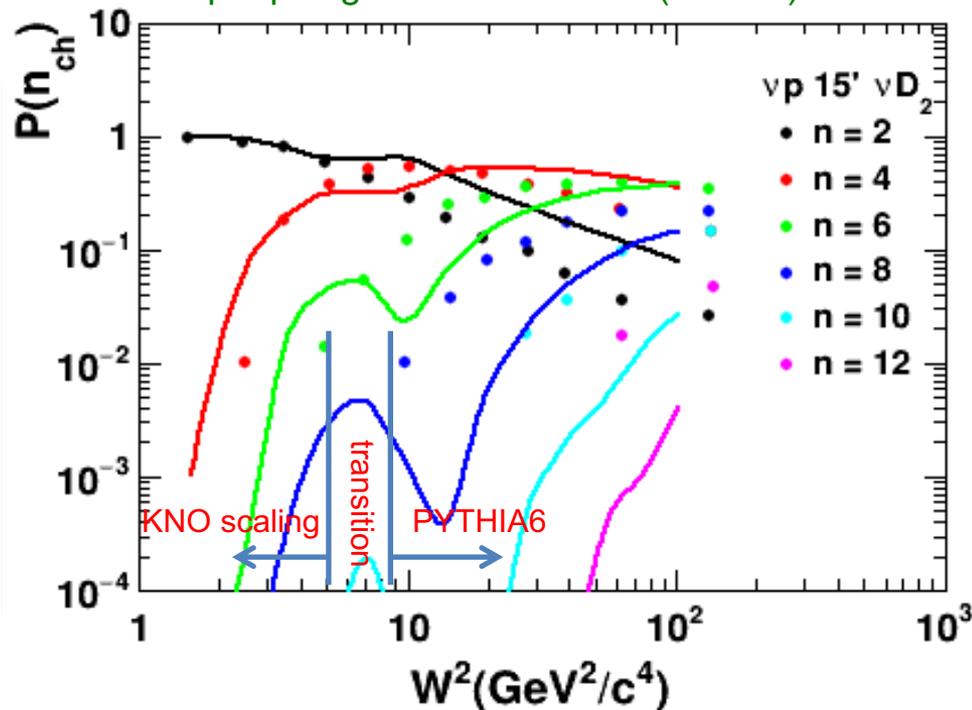
Although averaged charged hadron multiplicity makes continuous curve, topological cross sections are discontinuous, because multiplicity dispersion by PYTHIA6 is much narrower than bubble chamber data.

If the experiment is sensitive to hadron counting, you need to re-think how to propagate hadronization error...

Neutrino average charged hadron multiplicity



ν -p topological cross section (GENIE)





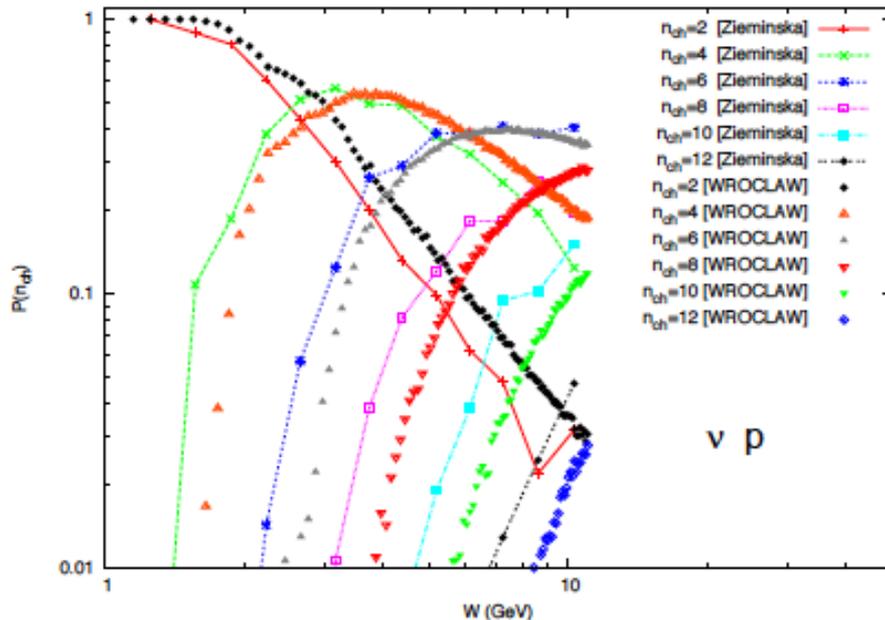
6. High-W hadronization dispersion error?

Bubble chamber topological cross section data

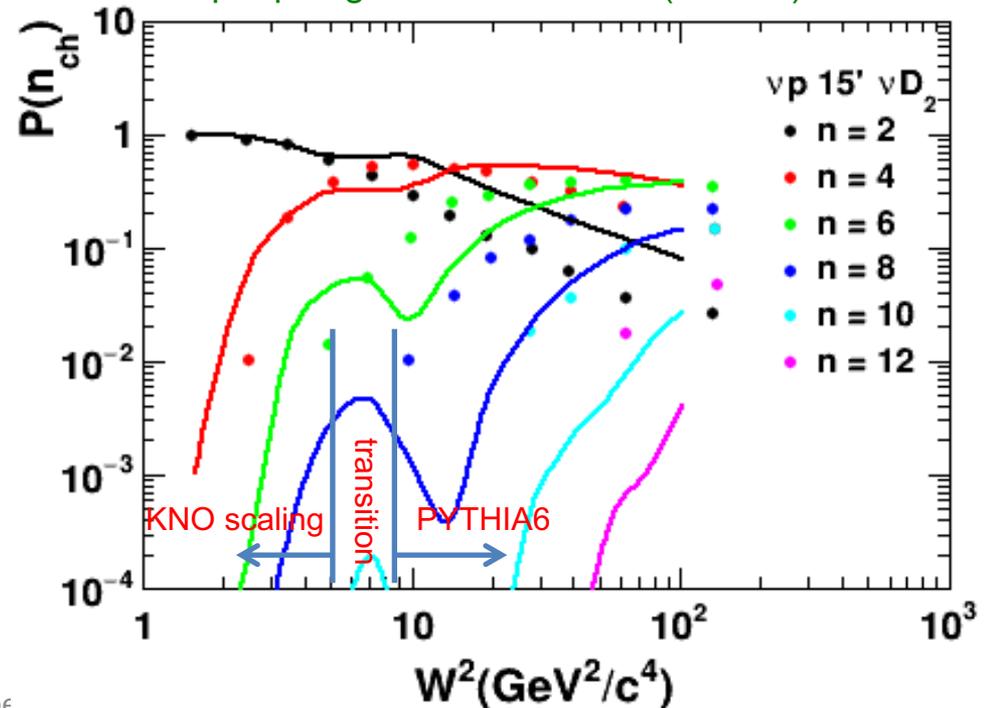
Although averaged charged hadron multiplicity makes continuous curve, topological cross sections are discontinuous, because multiplicity dispersion by PYTHIA6 is much narrower than bubble chamber data.

If the experiment is sensitive to hadron counting, you need to re-think how to propagate hadronization error...

v-p topological cross section (NuWro)



v-p topological cross section (GENIE)



7. DIS-hadronization errors, summary

- Goal is to make event weight with function of E_ν , x , y , etc, for IceCube oscillation program
- Some of systematic errors are identified to be dangerous

	DIS or Hadronization	type of error	approach	size
some study (MSU)	DIS	Bodek-Yang correction	play with Bodek-Yang parameters (by eyes)	maybe large?
done	DIS	differential xs	NuTeV-GENIE comparison (bottom-up)	1-2% by GENIE study
under investigation	DIS	A-scaling	MINERvA-GENIE (bottom-up)	maybe large?
some study (MSU)	DIS	PDF	From nuclear PDF, CT10? nCTEQ? (top-down)	expected to be tiny
under investigation	Hadronization	low W averaged charged hadron multiplicity	play with KNO parameters (by eyes)	maybe large?
done JPhysG42(2015)115004	Hadronization	high W averaged charged hadron multiplicity	bubble chamber-PYTHIA comparison (bottom-up)	1-2% by GENIE study

7. Conclusion

- Goal is to make event weight with function of E_ν , x , y , etc, for IceCube oscillation program
- Some of systematic errors are identified to be dangerous

Bodek-Yang correction : We must update it in the near future. It is out-of-date, and systematic error estimation is impossible.

DIS differential cross section error : Systematic error is under control.

DIS A-dependent error : It may be large, need a careful study. MINERvA data is a starting point but a dedicated argon-DIS cross section measurement may remove this error (NuMI beam at SBN?)

DIS PDF error : It looks error can be estimated easily.

Low W hadronization error : It may be large, need a careful study. Incoherent treatment of resonances (RES) and non-resonant background (DIS) may be a problem.

High W hadronization error : Systematic error is under control.

Thank you for your attention!



Back up

1. Neutrino cross section overview

GENIE uses “Frankenstein” model..., there are 2 transtions for both cross section and hadronization

Cross section

$W^2 < 2.9 \text{ GeV}^2$: RES

$W^2 > 2.9 \text{ GeV}^2$: DIS

Hadronization (AGKY model)

$W^2 < 5.3 \text{ GeV}^2$: KNO scaling based model

$5.3 \text{ GeV}^2 < W^2 < 9.0 \text{ GeV}^2$: transition

$9.0 \text{ GeV}^2 < W^2$: PYTHIA6

ν_μ CC on H_2O target with atmospheric neutrino flux in W_{recon}

