### Karpacz Winter School



in collaboration with Oscar-Garcia Montero, Marco Müller and Hannah Elfner







### June 20<sup>TH</sup> 2022

# Nils Sass



# **Signals of Polarization**



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L. Adamczyk et al. [STAR], Nature 548 (2017), 62-65 Snellings 2011 New J. Phys. 13 055008 (adoption)

### **STAR Measurement**

- Global spin-polarization of Λ hyperons in non-central Au+Au collisions at  $\sqrt{s_{NN}}$  = 3 - 200 GeV and mid-rapidity
- Hyperon polarization is a direct hint of vorticity of the QGP
- Polarization decreases for lower beam energies



- collisions by inducing vorticity
- transfer to the fireball is crucial
- signals

### Angular momentum of the nuclei is the driver of spin polarization in heavy-ion

To model hyperon polarization, the dynamic description of angular momentum

Detailed understanding of the dynamics is important to identify phase transition





**Pb-Pb** collision at  $E_{lab} = 40 \text{ GeV}$ 







t = 6 fm

© Justin Mohs



t = 12 fm

http://smash-transport.github.io

- \*Simulating Many Accelerated Strongly-Interacting Hadrons
- Dynamical non-equilibrium description of HICs at low beam energies (FAIR) and late stage rescattering at high beam energies (RHIC/LHC)
- Includes all hadrons up from the PDG to m ~ 2.35 GeV

### **SMASH Setup**

Geometric collision criterion 

$$d_{trans} < d_{int} = \sqrt{\frac{\sigma_{tot}}{\pi}} \qquad d_{trans}^2 = (\vec{r}_a - \vec{r}_b)^2 - \frac{((\vec{r}_a - \vec{r}_b) \cdot (\vec{p}_a - \vec{p}_b))^2}{(\vec{p}_a - \vec{p}_b)^2}$$

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### **SMASH**



Weil et al., Phys. Rev. C 94 (2016)

Effective solution of the relativistic Boltzmann equation

$$p^{\mu}\partial_{\mu}f_{i}(x,p) + m_{i}F^{\alpha}\partial_{\alpha}^{p}f_{i}(x,p) = C_{coll}^{i}$$

Test particle method  $\sigma \rightarrow \sigma \cdot N_{test}^{-1}$ ,  $N \rightarrow N \cdot N_{test}$ 









# **Angular Momentum & Fermi Motion in SMASH**

- SMASH accesses the full phase-space information of every particle and thus its angular momentum
- Fermi motion: nucleonic momentum distribution due to Pauli exclusion principle
- In the ground state nucleus, nucleon momentum distribution corresponds to a uniformly filled Fermi sphere with radius

$$p_F(\vec{r}) = \hbar c \left(3\pi^2 \rho(\vec{r})\right)^{\frac{1}{3}}$$





# Angular Momentum & Fermi Motion in SMASH



- Energy dependence of impact parameter b<sub>max</sub> for which the remaining angular momentum L<sub>r</sub> becomes maximal
- Fermi motion induces additional angular momentum into the system

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#### N. Sass et al, in preparation

- System size dependence of the ratio of the fireball's angular momentum over the initial angular momentum  $L_r / L_0$
- More angular momentum is deposited at mid rapidity in more central collisions and at lower beam energies



# **Angular Momentum Evolution**



- Broken angular momentum conservation

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#### N. Sass et al, in preparation

Secondary collisions at higher beam energies shift the flattening of L<sub>sp</sub> and L<sub>r</sub> to later times

We observe a kink in the total angular momentum at the time when both nuclei collide

For higher beam energies the kink occurs at smaller times due to faster time evolution



### **Angular Momentum Conservation**

- Additional momenta by Fermi motion slightly increase non-conservation of angular momentum
- Geometrical Interpretation of the cross section breaks angular momentum conservation in binary in/elastic collisions



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• Collective "loss" of angular momentum in Au-Au collisions amounts to 3.5% for small impact parameters



## **Treating Conservation in Elastic Scatterings**

Calculate the change in angular momentum  $\Delta L$  and mass moment  $\Delta K$  for each binary collision:

$$\Delta \overrightarrow{L} = (\overrightarrow{x}_1 - \overrightarrow{x}_2) \times \overrightarrow{p} - (\overrightarrow{x}_1' - \overrightarrow{x}_2') \times \overrightarrow{p'}$$
  
$$\Delta \overrightarrow{K} = (t_1 - t_2)(\overrightarrow{p} - \overrightarrow{p'}) - (\overrightarrow{x}_1 - \overrightarrow{x}_1')E_1 - (\overrightarrow{x}_2 - \overrightarrow{x}_2')E_2$$

Get the corrected positions of the scattered particles restricting collision trajectories to a plane and solving

$$\Delta \vec{L} = 0, \quad \Delta \vec{K} = 0$$

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## **Conclusion & Outlook**

- Impact parameter b<sub>max</sub> for which the angular momentum of the fireball becomes maximal is nearly energy independent for a broad energy range →  $b_{max} \in [4.5 \text{ fm}, 6.6 \text{ fm}] \text{ for } \sqrt{s_{NN}} \in [2.41 \text{ GeV}, 200 \text{ GeV}]$
- Higher relative transfer of initial angular momentum to the fireball in more central collisions
- Fermi motion induces additional angular momentum
- Predictions for expectation of high angular momentum is important for future experimental measurements
- Detailed understanding of the dynamics is crucial for extracting phase transition signals
- Long-term goal: Implementing spin degrees of freedom to describe hadronic polarization within the transport approach

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# **Backup Slides**

# **Spin Polarization**



#### L. Adamczyk et al. [STAR], Nature 548 (2017), 62-65

- Spin polarization: Alignment of particle spins with system's angular momentum
- polarization (Chiral Magnetic Effect)

Global polarization of quarks & anti-quarks due to spin-orbit coupling

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Deng and Huang, Phys. Rev. C 93 (2016) no.6, 064907

Non-central heavy-ion collisions induce vorticity and orbital angular momentum in the QGP Studying polarization observables might give hints of magnetic fields as another source for





### **Initial Conditions**

Sampling of the initial nuclei in coordinate space according to the Woods-Saxon distribution

$$\frac{dN}{d^3r} = \frac{\rho_0}{\exp\left(\frac{r-r_0}{d}+1\right)}$$

- d: diffusiveness of the nucleus
- $\rho_0$ : nuclear ground state density

• Hard sphere limit:  $d \rightarrow 0$ 



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## SMASH

### **Fermi Motion**

- Nuclei get additional momenta
- Nuclei are "stable" if additional potentials are turned on
- "Frozen" Fermi motion only considered for collision and turned off for propagation

 $\Gamma_{R \to ab} = \Gamma^0_{R \to ab} \frac{\rho_{ab}(m)}{\rho_{ab(M_0)}}$ 



### **\_\_ Resonances**

 $\mathscr{A}(m) = \frac{2\mathscr{N}}{-\!\!-\!\!-}$ 

- Particles with widths < 10 keV treated as stable</li>
- Unstable particles assigned a relativistic Breit-Wigner spectral function

$$\frac{m^2 \Gamma(m)}{(m^2 - M_0^2)^2 + m^2 \Gamma(m)^2}$$

$$\frac{m : \text{ resonance m}}{M_0 : \text{ pole mass}}$$

$$\Gamma(m) : \text{ width function}$$

$$\mathcal{N} : \text{ normalization}$$

• Decay width of two body decay  $R \rightarrow ab$  by treatment of Manley et al.

$$\rho_{ab}(m)$$
: mass integrals over resonance spectral fur

$$\Gamma^0_{R \to ab} = \Gamma_{R \to ab}(M_0)$$

### **ELEMENTS General Assembly**







## **Angular Momentum And Impact Parameter**

- Angular Momentum as function of the impact parameter shows a distinct maximum at a single b<sub>max</sub>
- Qualitative agreement with predictions from geometrical Glauber model



Becattini et al, Phys. Rev. C77:024906, 2008



