

# Fluid Dynamics and Kinetic theory

## 1. Basic elements of fluid dynamics

- Basic elements: energy and momentum density and fluxes

↔ Energy-momentum tensor  $T^{\mu\nu}$ :

$T^{00}$  energy density

$T^{0i}$   $i=(x,y,z)$  momentum density

$T^{i0}$  energy flux in  $i$ -direction (or flux density)

$T^{ij}$   $j$ -momentum flux in  $i$ -direction

- Energy conservation:  $\partial_t T^{00} = -\partial_x T^{x0} - \partial_y T^{y0} - \partial_z T^{z0}$

infinitesimal  
cube: →



$$\underbrace{\partial_t \int T^{00} dV}_{\text{change of } E \text{ within the cube}} = - \underbrace{\oint T^{i0} dA_i}_{\text{energy flux through the boundary}} \quad \xrightarrow{V \rightarrow 0}$$

- Momentum conservation  $\partial_t T^{0i} = -\partial_x T^{xi} - \partial_y T^{yi} - \partial_z T^{zi}$

$\Rightarrow$  Energy-momentum conservation:

$$\boxed{\partial_\mu T^{\mu\nu} = 0} \quad (1.1)$$

- This is the most fundamental equation in fluid dynamics

- Requires only continuously distributed energy and momentum

- Conserved charges: Charge 4-current  $N_q^\mu$

$N_q^0$ : charge density

$N_q^i$ : charge flux in  $i$ -direction

- Charge conservation  $\partial_t \int N_q^0 dV = -\int N_q^i dA_i \xrightarrow{V \rightarrow 0} \partial_t N_q^0 = -\partial_i N_q^i$

$$\boxed{\partial_\mu N_q^\mu = 0} \quad (1.2)$$

- Angular momentum tensor:  $\mathcal{M}^{\alpha\beta\mu} = x^\alpha T^{\beta\mu} - x^\beta T^{\alpha\mu}$

- angular momentum conservation

$$\begin{aligned}
 0 = \partial_\mu \mathcal{M}^{\alpha\beta\mu} &= \underbrace{(\partial_\mu x^\alpha)}_{=\delta_\mu^\alpha} T^{\beta\mu} - \underbrace{(\partial_\mu x^\beta)}_{=\delta_\mu^\beta} T^{\alpha\mu} + x^\alpha \underbrace{\partial_\mu T^{\beta\mu}}_{=0} - x^\beta \underbrace{\partial_\mu T^{\alpha\mu}}_{=0} \\
 &= T^{\beta\alpha} - T^{\alpha\beta} \quad \Rightarrow \quad T^{\alpha\beta} = T^{\beta\alpha} \quad (T^{\mu\nu} \text{ symmetric})
 \end{aligned}$$

- The basic quantities and equations of motion are now

$$\partial_\mu T^{\mu\nu} = 0 \quad \partial_\mu N_g^M = 0$$

- Energy-momentum conservation for continuous matter

- This defines fluid dynamics, but not yet very useful:

- 5 equations but 14 unknowns  $\left( \begin{array}{l} T^{\mu\nu} - 10 \text{ components (independent)} \\ N^M - 4 \text{ components} \end{array} \right)$

- Continuous matter : motion described by fluid 4-velocity  $u^\mu$
- Position of fluid element  $x^\mu(\tau)$ ,  $\tau$  is a proper time (in comoving frame)
 
$$u^\mu = \frac{dx^\mu}{d\tau} = \left( \frac{dt}{d\tau}, \frac{d\vec{x}}{d\tau} \right) = \gamma (1, \vec{v}) \quad \text{where } \vec{v} = \frac{d\vec{x}}{dt}, \quad \gamma = \frac{1}{\sqrt{1-v^2}} = \frac{dt}{d\tau}$$
- $u^\mu$  normalized :  $u_\mu u^\mu = \gamma^2 (1 - \vec{v}^2) = 1$
- $u^\mu = (1, \vec{0})$  defines Local Rest Frame (LRF), where the fluid element is at rest.
- Using  $u^\mu$  we can define

• co-moving time derivative (material derivative) of quantity  $A$

$$\dot{A} = u^\mu \partial_\mu A \stackrel{\text{LRF: } u^\mu = (1, \vec{0})}{=} \partial_t A = \frac{d}{d\tau} A \quad (\text{note } u^\mu \partial_\mu \text{ is scalar } \rightarrow \text{same in any frame}) \quad (1.3)$$

↑  
now time in LRF

$$\| \text{also } \underbrace{u^\mu \partial_\mu x^\nu}_{= \delta_\mu^\nu} = u^\nu \quad (\text{as it should})$$

$$\text{note : } u^\mu \partial_\mu A = (\gamma \partial_t + \gamma \vec{v} \cdot \nabla) A$$

$$\| \text{non-relativistic : } \frac{dA}{d\tau} = (\partial_t + \vec{v} \cdot \nabla) A$$

▷ energy and charge density in LRF

$$u_\mu u_\nu T^{\mu\nu} = T_{\text{LRF}}^{00} = e \quad (1.4)$$

$$u_\mu N_q^\mu = N_{q, \text{LRF}}^0 = n_q \quad (1.5)$$

• In fluid dynamics important distinction between convection and diffusion

▷ Convection: energy / charge / momentum transport with flow  $u^\mu$

▷ Diffusion:  $\text{---} \quad \parallel \quad \text{---}$  without / orthogonal to flow

• Decomposition w.r.t.  $u^\mu$

$$N_q^\mu = N_q^\alpha \delta_\alpha^\mu = N_q^\alpha \left( u_\alpha u^\mu + \underbrace{\delta_\alpha^\mu - u_\alpha u^\mu}_{\equiv \Delta_\alpha^\mu} \right) = \underbrace{n_q u^\mu}_{\text{convection}} + \underbrace{V_q^\mu}_{\text{diffusion}} \quad V_q^\mu = \Delta_\nu^\mu N_q^\nu$$

▷  $\Delta^{\mu\nu} = g^{\mu\nu} - u^\mu u^\nu$  is projection operator (orthogonal to  $u^\mu$ )

$$\parallel \Delta^{\mu\nu} u_\mu = \Delta^{\mu\nu} u_\nu = 0 \quad \& \quad \Delta_\alpha^\mu \Delta_\nu^\alpha = \Delta_\nu^\mu$$

•  $T^{\mu\nu}$  similarly

$$T^{\mu\nu} = T^{\alpha\beta} \delta_{\alpha}^{\mu} \delta_{\beta}^{\nu} = T^{\alpha\beta} (\delta_{\alpha}^{\mu} - u^{\mu} u_{\alpha} + u^{\mu} u_{\alpha}) (\delta_{\beta}^{\nu} - u^{\nu} u_{\beta} + u^{\nu} u_{\beta})$$

$$= T^{\alpha\beta} \Delta_{\alpha}^{\mu} \Delta_{\beta}^{\nu} + T^{\alpha\beta} \Delta_{\alpha}^{\mu} u^{\nu} u_{\beta} + T^{\alpha\beta} \Delta_{\beta}^{\nu} u^{\mu} u_{\alpha} + \underbrace{T^{\alpha\beta} u_{\alpha} u_{\beta} u^{\mu} u^{\nu}}_{= e}$$

Define  $W^{\mu} = T^{\alpha\beta} \Delta_{\alpha}^{\mu} u_{\beta}$

$$= e u^{\mu} u^{\nu} + W^{\mu} u^{\nu} + W^{\nu} u^{\mu} + T^{\alpha\beta} \underbrace{\Delta_{\alpha}^{\mu} \Delta_{\beta}^{\nu}}_{\text{symmetrize \& separate trace}}$$

$$\Delta_{\alpha}^{\mu} \Delta_{\beta}^{\nu} = \frac{1}{2} \left( \Delta_{\alpha}^{\mu} \Delta_{\beta}^{\nu} + \Delta_{\beta}^{\mu} \Delta_{\alpha}^{\nu} - \frac{2}{3} \Delta^{\mu\nu} \Delta_{\alpha\beta} \right) + \frac{1}{3} \Delta^{\mu\nu} \Delta_{\alpha\beta} + \text{antisymmetric}$$

$= \Delta_{\alpha\beta}^{\mu\nu}$  projection operator  
symmetric, orthogonal to  $u^{\mu}$   
& traceless

→  $\odot$ , because  $T^{\mu\nu}$  symmetric  
↓

$$\Rightarrow T^{\mu\nu} = e u^\mu u^\nu - P \Delta^{\mu\nu} + 2 W^{(\mu} u^{\nu)} + \pi^{\mu\nu}$$

$$e = T^{\mu\nu} u_\mu u_\nu \quad (\text{LRF energy density})$$

$$P = -\frac{1}{3} T^{\alpha\beta} \Delta_{\alpha\beta} \quad (\text{Isotropic pressure})$$

$$W^\mu = T^{\alpha\beta} \Delta_\alpha^\mu u_\beta \quad (\text{Energy diffusion current})$$

$$\pi^{\mu\nu} = \Delta_{\alpha\beta}^{\mu\nu} T^{\alpha\beta} \equiv T^{\langle\mu\nu\rangle} \quad \left( \begin{array}{l} \text{Shear-stress tensor} \\ \text{"momentum diffusion"} \end{array} \right)$$

$$N_q^\mu = n_q u^\mu + V_q^\mu$$

$$n_q = N_q^\mu u_\mu \quad (\text{LRF charge density})$$

$$V_q^\mu = \Delta_\nu^\mu N_q^\nu \quad (\text{Charge diffusion current})$$

- These are now the basic fluid dynamical quantities
- So far no help in closing the equations of motion (we have only introduced 3 new components of  $u^\mu$ )

(1.6)

(1.7)

- A second ingredient (besides energy, momentum and charge conservation) is an existence of a local equilibrium state,
- In equilibrium we can relate temperature  $T$ , and thermodynamical pressure  $P_{eq}$  to  $e_{eq}$  and  $n_{eq}$  through equation of state (EoS)

$$P_{eq} = P_{eq}(e_{eq}, n_{eq}) \quad T = T(e_{eq}, n_{eq}) \quad (1.8)$$

- In equilibrium: matter is locally isotropic (in LRF)  $\Rightarrow$

$$T_{eq}^{\mu\nu} = e_{eq} u^\mu u^\nu - P_{eq} \Delta^{\mu\nu} \quad N_{q,eq}^\mu = n_q u^\mu \quad (1.9)$$

- We can then split  $T^{\mu\nu}$  into equilibrium and viscous parts

$$T^{\mu\nu} = (e_{eq} + \delta e) u^\mu u^\nu - (P_{eq} + \Pi) \Delta^{\mu\nu} + 2W^{(\mu} u^{\nu)} + \pi^{\mu\nu} \quad (1.10)$$

$$N^\mu = (n_{eq} + \delta n) u^\mu + V_q^\mu$$

here  $\delta e = e - e_{eq}$  and  $\delta n = n - n_{eq}$  are differences between actual densities and densities of an equilibrium state.

$\Pi = P - P_{eq}$  is difference btw. total isotropic pressure and eq. pressure (bulk viscous pressure)



- For a given state  $T^{\mu\nu}$ , equilibrium state is not unique, but need to be defined.

- A usual choice is the equilibrium state for which

$$e_{eq} = e \quad \text{and} \quad n_{eq} = n \quad (\text{i.e. } \delta e = \delta n = 0) \quad (1.11)$$

- These are often referred as (Landau) matching conditions.

- In equilibrium  $e_{eq} = e_{eq}(T, \mu) \quad n_{eq} = n_{eq}(T, \mu)$

$\Rightarrow$  matching conditions can be thought as a definition of temperature and chemical potential for a non-equilibrium state,

$$\Rightarrow T^{\mu\nu} = e u^\mu u^\nu - (p_{eq} + \Pi) \Delta^{\mu\nu} + 2W^{\langle\mu} u^{\nu\rangle} + \Pi^{\mu\nu} \quad (1.12)$$

$$N_q^\mu = n_q u^\mu + V_q^\mu$$

- Dissipative quantities are now  $\Pi, W^\mu, \Pi^{\mu\nu}$  and  $V_q^\mu$

Note: only definition of bulk viscous pressure  $\Pi$  need explicit reference to equilibrium, other dissipative quantities can be obtained by projecting  $T^{\mu\nu}$  and  $N_q^\mu$  with  $u^\mu$ .

- Similarly as equilibrium state, fluid velocity is not unique (only in equilibrium), but needs to be defined.
- In order to define  $u^\mu$  we need to decide what "flows"
- Two common choices: net-charge flow (Eckart frame), or total energy flow (Landau frame)

• Eckart frame:  $u^\mu = \frac{N_q^\mu}{\sqrt{N_q^\alpha N_q^\alpha}} \Rightarrow \Delta^{\mu\nu} = g^{\mu\nu} - \frac{N_q^\mu N_q^\nu}{N_q \cdot N_q}$

$\Rightarrow V_q^\mu = N_q^\nu \Delta_{\nu}^{\mu} = N_q^\mu - N_q^\mu \frac{N_q \cdot N_q}{N_q \cdot N_q} = 0$

$\Rightarrow$  Charge diffusion vanishes  
(no flow of charge  $\perp u^\mu$ )

• Landau frame:  $T^{\mu\nu} u_\nu = e u^\mu$

In general:  $T^{\mu\nu} u_\nu = e u^\mu + W^\mu$

$\Rightarrow$  In Landau frame energy diffusion vanishes

$W^\mu = 0$

Let's use this from now on. Eckart frame is sometimes ill defined, e.g. when  $n_q = 0$ .

- Energy-momentum tensor and charge  $U$ -current are now expressed as

$$T^{\mu\nu} = e u^\mu u^\nu - (p_{eq}(e, n_f) - \Pi) \Delta^{\mu\nu} + \pi^{\mu\nu}$$

$$N_f^\mu = n_f u^\mu + V_f^\mu$$

- Equilibrium state defined through matching  $e_{eq}(\tau, \mu) = e$   $n_{eq}(\tau, \mu) = n_f$
- Fluid  $U$ -velocity defined through  $T^{\mu\nu} u_\nu = e u^\mu$  (Landau frame)
- We can now cast the conservation laws into a more understandable form

$$0 = u_\nu \partial_\mu T^{\mu\nu} = u_\nu \partial_\mu [e u^\mu u^\nu - (p_{eq} + \Pi) \Delta^{\mu\nu} + \pi^{\mu\nu}]$$

$$\begin{aligned} \parallel \text{ use } & u^\mu \partial_\mu e = \dot{e} \\ & u^\mu \partial_\nu u_\mu = 0 \quad (u^\mu u_\mu = 1 \Rightarrow \partial_\nu (u^\mu u_\mu) = 0 = 2u^\mu \partial_\nu u_\mu) \\ & \partial_\mu g^{\mu\nu} = 0 \\ & u_\nu \partial_\mu \pi^{\mu\nu} = -\pi^{\mu\nu} \partial_\mu u_\nu \end{aligned}$$

$$\Rightarrow \dot{e} = - \underbrace{(e + p_{eq} + \Pi)}_{\text{}} \partial_\mu u^\mu + \pi^{\mu\nu} \partial_\mu u_\nu$$

- Evolution of energy density driven by gradients of  $u^\mu$

- Decompose  $\partial_\mu u_\nu$

• Define 3-gradient:  $\nabla'_\mu = \Delta_{\mu\nu} \partial_\nu \stackrel{\text{LRF}}{=} (0, \vec{\nabla}) \quad || \quad u^\mu \nabla'_\mu = 0$

$$\partial_\mu = \delta_\mu^\alpha \partial_\alpha = \underbrace{(u^\alpha u_\mu + \delta_\mu^\alpha - u^\alpha u_\mu)}_{= \Delta_\mu^\alpha} \partial_\alpha = u_\mu \underbrace{u^\alpha \partial_\alpha}_{\frac{d}{d\tau}} + \nabla'_\mu = u_\mu \frac{d}{d\tau} + \nabla'_\mu$$

$$\Rightarrow \partial_\mu u_\nu = u_\mu \frac{d}{d\tau} u_\nu + \underbrace{\nabla'_\mu u_\nu}$$

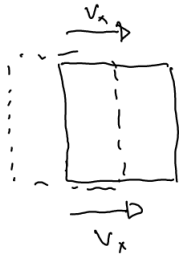
$$= \underbrace{\frac{1}{2} (\nabla'_\mu u_\nu + \nabla'_\nu u_\mu)}_{\text{symmetric}} + \underbrace{\frac{1}{2} (\nabla'_\mu u_\nu - \nabla'_\nu u_\mu)}_{\text{antisymmetric}}$$

$$= u_\mu \dot{u}_\nu + \underbrace{\frac{1}{2} (\nabla'_\mu u_\nu + \nabla'_\nu u_\mu - \frac{2}{3} \Delta_{\mu\nu} \nabla'_\alpha u^\alpha)}_{\substack{\downarrow \\ = \sigma_{\mu\nu}}} + \frac{1}{3} \Delta_{\mu\nu} \underbrace{\nabla'_\alpha u^\alpha}_{= \Theta} + \underbrace{\frac{1}{2} (\nabla'_\mu u_\nu - \nabla'_\nu u_\mu)}_{= \omega_{\mu\nu}}$$

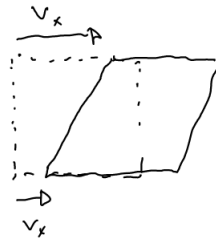
$\uparrow$  add and remove trace

$$\Rightarrow \partial_\mu u_\nu = u_\mu \dot{u}_\nu + \sigma_{\mu\nu} + \frac{1}{3} \Delta_{\mu\nu} \Theta + \omega_{\mu\nu}$$

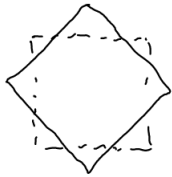
- Different parts here correspond different deformations of fluid element



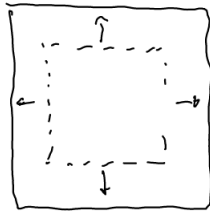
$$u_\mu \dot{u}_\nu = \text{acceleration}$$



$$\sigma_{\mu\nu} = \Delta_{\mu\nu}^{\alpha\beta} \partial_\alpha u_\beta = \text{shear deformation (do not change volume, that is why we removed trace)}$$



$$\omega_{\mu\nu} = \text{rotations (vorticity tensor)}$$



$$\Delta_{\mu\nu} \Theta = \text{expansion} \quad \Theta = \nabla_\mu u^\mu = \partial_\mu u^\mu = \text{volume expansion rate}$$

- Different type of deformations associated with different dissipative processes



- Charge conservation

$$\partial_\mu N_q^\mu = \partial_\mu (n_q u^\mu + V_q^\mu) = \dot{n}_q + n_q \theta + \partial_\mu V_q^\mu = 0$$

$$\Rightarrow \dot{n}_q = - \underbrace{n_q \theta}_{\text{expansion}} - \underbrace{\partial_\mu V_q^\mu}_{\text{diffusion}}$$

- We have now written the conservation laws in terms of our new fluid variables

$$\dot{e} = -(e + p_{eq} + \pi) \theta + \pi^{\mu\nu} \sigma_{\mu\nu}$$

$$(e + p_{eq} + \pi) \dot{u}^\mu = \nabla^\mu (p_{eq} + \pi) + \nabla_\alpha \pi^{\mu\alpha}$$

$$\dot{n}_q = -n_q \theta - \partial_\mu V_q^\mu$$

- Doesn't solve the problem that we have 14 unknowns and only 5 equations

• Need 9 additional equations to close the system :

$$\begin{aligned} \pi^{\mu\nu} &= \pi^{\mu\nu}(e, n_q, \nabla u, \nabla e, \nabla n_q) & V_q^\mu &= V_q^\mu(e, n_q, \nabla u, \nabla e, \nabla n) & \theta &= \theta(e, n_q, \nabla u, \nabla n, \nabla e) \\ (5) & & (3) & & (1) \end{aligned}$$

- Usually we think fluid dynamics in situation, where densities, and velocity change slowly in space and time

⇒ Include only 1st-order gradients

$$\begin{cases} \pi^{\mu\nu} = 2\eta\sigma^{\mu\nu} \\ V_q^\mu = \kappa\nabla^\mu\left(\frac{\mu}{T}\right) \\ \Pi = -\zeta\Theta \end{cases}$$

- $\eta$ ,  $\kappa$ , and  $\zeta$  positive : material properties (like EoS)

$\eta(T, \mu)$	shear viscosity
$\kappa(T, \mu)$	diffusion constant
$\zeta(T, \mu)$	bulk viscosity

- This is relativistic generalization of Newtonian fluids

(usually referred as relativistic Navier-Stokes theory)



- $$\frac{\pi^{\mu\nu}}{p_{eq}} = \frac{2\eta}{p_{eq}} \underbrace{\sigma^{\mu\nu}}_{\text{dimension fm}} \rightarrow \nabla u \text{ dimension fm}^{-1} \text{ (macroscopic scale over which velocity varies)}$$

$$\frac{2\eta}{p_{eq}} \sim \lambda_{mfp} \text{ (microscopic scale)}$$

$$\Rightarrow \frac{2\eta}{p_{eq}} \sigma^{\mu\nu} \sim O(Kn)$$

- Relativistic Navier-Stokes theory is 1st order in Knudsen number  $Kn$

$$Kn = \frac{\lambda_{micr}}{L_{macr}}$$

- $Kn$  quantifies degree of separation between microscopic and macroscopic scales

- Fluid dynamics when  $Kn \lesssim 1$

- As presented here, relativistic Navier-Stokes theory is not a good relativistic theory:
  - signal propagation speed can exceed speed of light  $\rightarrow$  acausal theory

- Problem is that in NS theory the microscopic state of matter (e.g.  $\pi^{\mu\nu}$ ) reacts immediately to changes in external conditions e.g.  $\nabla u$  or  $\sigma^{\mu\nu}$ . In reality this takes time  $\tau_{mic}$  between particle collisions.

- We can incorporate  $\tau_{mfc}$  into fluid dynamics, and resolve it explicitly  
 → Transient fluid dynamics (e.g. Israel-Stewart theory)
- $\Pi^{\mu\nu} = 2\eta\sigma^{\mu\nu} \implies \dot{\Pi}^{\langle\mu\nu\rangle} = \frac{1}{\tau_{\pi}} (2\eta\sigma^{\mu\nu} - \Pi^{\mu\nu}) + \text{higher order terms}$   
 + similarly for  $V_q^\mu$  and  $\Pi$
- Next our goal is to derive this type of fluid dynamics from kinetic theory of relativistic gases.

## 2. Elements of kinetic theory

- The basic quantity in kinetic theory is a single-particle distribution function

$$f(x, k) \equiv f_k$$

- This is probability density for observing particle at spacetime point  $[x, x+dx]$  with momentum  $[k, k+dk]$
- Macroscopic densities and fluxes can be obtained by integrating over  $k$

- Particle density  $N^0 = \int \underbrace{\frac{d^3\vec{k}}{(2\pi)^3}}_{\text{density of states}} f_k$

- Particle flux  $N^i = \int \frac{d^3\vec{k}}{(2\pi)^3} \underbrace{\frac{k^i}{k^0}}_{\text{particle velocity}} f_k$ ,  $k^0 = \sqrt{\vec{k}^2 + m^2}$

combine both:

$\Rightarrow$  Particle 4-current

$$N^\mu = \int \underbrace{\frac{d^3\vec{k}}{(2\pi)^3 k^0}}_{\text{scalar}} \underbrace{k^\mu}_{\text{4-vector}} f_k$$

(2.1)

- Energy density  $T^{00} = \int \frac{d^3\vec{k}}{(2\pi)^3} k^0 f_k$
- Energy flux  $T^{i0} = \int \frac{d^3\vec{k}}{(2\pi)^3} \frac{k^i}{k^0} k^0 f_k$
- Momentum density  $T^{0i} = \int \frac{d^3\vec{k}}{(2\pi)^3} k^i f_k$
- $\hat{j}$ -momentum flux in  $\hat{i}$ -direction  $T^{ij} = \int \frac{d^3\vec{k}}{(2\pi)^3} \frac{k^i}{k^0} k^j f_k$

combine  $\Rightarrow T^{\mu\nu} = \int \frac{d^3\vec{k}}{(2\pi)^3 k^0} k^\mu k^\nu f_k \quad (2.2)$

notation  $dK = \frac{d^3\vec{k}}{(2\pi)^3 k^0} \quad (2.3)$

- We have now our basic fluid quantities expressed as integrals over  $f_k$

$$T^{\mu\nu} = \int dK k^\mu k^\nu f_k \quad N^\mu = \int dK k^\mu f_k \quad (2.4)$$

- We can now decompose  $T^{\mu\nu}$  and  $N^\mu$  as before, so that definitions (1.6) remain the same (still use Landau velocity)

$$T^{\mu\nu} = e u^\mu u^\nu - (p_{eq} + \Pi) \Delta^{\mu\nu} + \pi^{\mu\nu} \quad N^\mu = n u^\mu + V^\mu$$

$$e = T^{\mu\nu} u_\mu u_\nu = \int dK E_k^2 f_k, \quad \text{where } E_k = k^\mu u_\mu \stackrel{\text{LRF}}{=} k_{\text{LRF}}^0 \quad (\text{particle energy in LRF})$$

$$p_{eq} + \Pi = -\frac{1}{3} \Delta_{\mu\nu} T^{\mu\nu} = -\frac{1}{3} \int dK (\Delta_{\mu\nu} k^\mu k^\nu) f_k$$

$$\Pi^{\mu\nu} = T^{\langle\mu\nu\rangle} = \int dK k^{\langle\mu} k^{\nu\rangle} f_k$$

(2.5)

$$n = N^\mu u_\mu = \int dK E_k f_k$$

$$V^\mu = N^\alpha \Delta_\alpha^\mu = \int dK k^{\langle\mu} \rangle f_k, \quad \text{where } k^{\langle\mu} \rangle = \Delta_\alpha^\mu k^\alpha \quad k^{\langle\mu} \rangle u_\mu = 0$$

- Equilibrium distribution  $f_{eq}$  is given by Bose-Einstein or Fermi-Dirac distr.

$$f_{eq}(T, \mu) = \left[ e^{(E_k - \mu)/T} + a \right]^{-1} \quad a = \begin{cases} 1, & \text{fermions} \\ -1, & \text{bosons} \\ 0, & \text{classical particles} \end{cases} \quad (2.6)$$

- As before, we wish to include equilibrium state explicitly, and write

$$f_k = f_{eq} + \delta f \quad (2.7)$$

- $\delta f$  is now deviation from equilibrium distribution, in equilibrium  $\delta f = 0$
- For a general  $f_k$  equilibrium state is not unique: invoke matching conditions

$$\left. \begin{aligned} e &= \int dK E_k^2 f_k = e_{eq} = \int dK E_k^2 f_{eq} \\ n &= \int dK E_k f_k = n_{eq} = \int dK E_k f_{eq} \end{aligned} \right\} \Rightarrow \begin{aligned} \delta e &= \int dK E_k^2 \delta f = 0 \\ \delta n &= \int dK E_k \delta f = 0 \end{aligned} \quad (2.8)$$

- $\delta f$  does not contribute energy nor particle density
- These conditions define  $T$  and  $\mu$  for a general off-equilibrium state  $f_k$
- In equilibrium:

↙ integrand: only tensors  $u^\mu$  and  $g^{\mu\nu}$  ( $k^\mu$  integrated)

$$T_{eq}^{\mu\nu} = \int dK k^\mu k^\nu f_{eq}(E_k = u \cdot k) = A u^\mu u^\nu + B g^{\mu\nu} = \frac{e_{eq} u^\mu u^\nu - p_{eq} \Delta^{\mu\nu}}{\quad} \quad (2.9)$$

$$N_{eq}^\mu = \int dK k^\mu f_{eq} = \underline{n u^\mu}$$

$$\begin{aligned} \Pi_{eq}^{\mu\nu} &= \Delta_{\alpha\beta}^{\mu\nu} T_{eq}^{\alpha\beta} = \int dK k^{\langle\mu} k^{\nu\rangle} f_{eq} = \Delta_{\alpha\beta}^{\mu\nu} (e_{eq} u^\alpha u^\beta - p_{eq} \Delta^{\alpha\beta}) = 0 \\ V_{eq}^\mu &= \Delta_\alpha^\mu N_{eq}^\alpha = \int dK k^{\langle\mu\rangle} f_{eq} = \Delta_\alpha^\mu (n u^\alpha) = \underline{0} \end{aligned} \quad (2.10)$$

- These are not restricted to equilibrium, but any  $F_k(E_k)$  that is only function of  $E_k$  give

$$\underbrace{\int dK k^{\langle\mu} k^{\nu\rangle} F_k}_{\text{directional information}} = \int dK k^{\langle\mu} F_k = 0 \quad (2.11)$$

( $F_k$  isotropic in LRF)

- Our expression of  $f_k$  is now

$$f_k = f_{eq} + \delta f \quad (2.12)$$

↑  
depends on  $u^\mu, T, \mu \rightarrow$  defined by velocity frame  
and matching conditions

- Obviously  $f_k$  has still infinitely many degrees of freedom, and 5 constraints from the conservation laws  $\partial_\mu T^{\mu\nu} = 0$  and  $\partial_\mu N^\mu = 0$  do not constrain it much.

- Full dynamics of  $f_k$  is given by the relativistic Boltzmann equation

$$k^\mu \partial_\mu f_k = \underbrace{C[f_k]}_{\text{collision integral}} \quad (2.13)$$

- Collision integral for binary collision  $k+k' \rightarrow p+p'$

$$C[f_k] = \frac{1}{2} \int dK' dP dP' W_{kk' \rightarrow pp'} (f_p f_{p'} \tilde{f}_k \tilde{f}_{k'} - f_k f_{k'} \tilde{f}_p \tilde{f}_{p'}) \quad (2.14)$$

$$W_{kk' \rightarrow pp'} \quad \text{transition rate}, \quad \tilde{f}_k = \underbrace{1 - a f_k}_{\text{quantum statistics}}$$

- Question now is: can we reduce the complicated dynamics given by the Boltzmann equation, and write the dynamics in terms of just few macroscopic quantities like  $T^{\mu\nu}$  and  $N^\mu$

↗  
i.e. can we write fluid dynamical limit of Boltzmann equation

- For this purpose we first write the Boltzmann equation entirely in terms of macroscopic quantities that are moments of  $f_k$ .

$$\text{e.g. } \pi^{\mu\nu} = \int dK k^{\langle\mu} k^{\nu\rangle} f_k, \quad \text{but also } \rho_n^{\mu\nu} = \int dK E_k^n k^{\langle\mu} k^{\nu\rangle} f_k$$

There are infinitely many such quantities  $\leftrightarrow$  infinitely many d.o.f. in  $f_k$



### 3. Expansion of $f_k$

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- $f_k = f_{eq} + \delta f = f_{eq} \left( 1 + \frac{\delta f}{f_{eq}} \right)$  (3.1)  
 $\hookrightarrow 1 - a f_{eq}$

- Expand  $f_k$  around equilibrium (in momentum space)

- Expansion basis  $1, k^\mu, k^\mu k^\nu, k^\mu k^\nu k^\lambda, \dots$  (like Israel and Stewart)

$$\Rightarrow \phi = \varepsilon + \varepsilon^\mu k_\mu + \varepsilon^{\mu\nu} k_\mu k_\nu + \varepsilon^{\mu\nu\lambda} k_\mu k_\nu k_\lambda + \dots \quad (3.2)$$

•  $\varepsilon$ 's are the expansion coefficients

- One way to reduce the degrees of freedom is to directly truncate the expansion.

- Israel & Stewart 14-moment approximation

$$\phi = \varepsilon + \varepsilon^\mu k_\mu + \varepsilon^{\mu\nu} k_\mu k_\nu \quad (3.3)$$

- The coefficients can be determined by requiring

$$T^{\mu\nu} = \int dK k^\mu k^\nu f_k = \int dK k^\mu k^\nu f_{eq} + \int dK k^\mu k^\nu f_{eq} \tilde{f}_{eq} (\varepsilon + \varepsilon^\mu k_\mu + \varepsilon^{\mu\nu} k_\mu k_\nu)$$

$$N^\mu = \int dK k^\mu f_k = \int dK k^\mu f_{eq} + \int dK k^\mu f_{eq} \tilde{f}_{eq} (\varepsilon + \varepsilon^\mu k_\mu + \varepsilon^{\mu\nu} k_\mu k_\nu)$$

$\Rightarrow$  express  $\varepsilon$ 's in terms of  $T^{\mu\nu}$  and  $N^\mu$

Example:

$$\underline{\underline{\Pi^{\mu\nu}}} = \Delta_{\alpha\beta}^{\mu\nu} \int dK k^\alpha k^\beta f_k = \Delta_{\alpha\beta}^{\mu\nu} \int dK k^\alpha k^\beta f_{eq} + \Delta_{\alpha\beta}^{\mu\nu} \int dK f_{eq} \tilde{f}_{eq} k^\alpha k^\beta (\varepsilon + \varepsilon^\gamma k_\gamma + \varepsilon^{\gamma\delta} k_\gamma k_\delta)$$

$$= \Delta_{\alpha\beta}^{\mu\nu} \int dK f_{eq} \tilde{f}_{eq} k^\alpha k^\beta k^\gamma \varepsilon_\gamma + \Delta_{\alpha\beta}^{\mu\nu} \int dK f_{eq} \tilde{f}_{eq} k^\alpha k^\beta k^\gamma k^\delta \varepsilon_{\gamma\delta}$$

$$= \varepsilon_\gamma \left[ \underbrace{A u^\alpha u^\beta u^\gamma}_{\downarrow 0} + \underbrace{B u^{(\alpha} \Delta^{\beta\gamma)}}_{\downarrow \& \text{trace}} \right] = \varepsilon_{\gamma\delta} \left[ \underbrace{A' u^\alpha u^\beta u^\gamma u^\delta}_{\downarrow 0} + \underbrace{B' u^{(\alpha} u^\beta \Delta^{\gamma\delta)}}_{\downarrow 0} + \underbrace{C' \Delta^{(\alpha\beta} \Delta^{\gamma\delta)}}_{\downarrow 0} \right]$$

$$= \Delta_{\alpha\beta}^{\mu\nu} \varepsilon_{\gamma\delta} C' \Delta^{(\alpha\beta} \Delta^{\gamma\delta)} = \varepsilon_{\gamma\delta} C' \Delta_{\alpha\beta}^{\mu\nu} \frac{1}{3} \left( \Delta^{\alpha\beta} \Delta^{\gamma\delta} + \Delta^{\alpha\gamma} \Delta^{\beta\delta} + \Delta^{\alpha\delta} \Delta^{\beta\gamma} \right) = \frac{2}{3} \varepsilon_{\gamma\delta} C' \Delta^{\mu\nu \gamma\delta}$$

$$= \frac{2}{3} C' \underline{\underline{\varepsilon^{\langle\mu\nu\rangle}}}$$

- Coefficient  $C'$  can be computed from

$$J^{\alpha\beta\gamma\delta} = \int dK f_{eq} \tilde{f}_{eq} k^\alpha k^\beta k^\gamma k^\delta = A' u^\alpha u^\beta u^\gamma u^\delta + B' u^{(\alpha} u^{\beta} \Delta^{\gamma\delta)} + C' \Delta^{(\alpha\beta} \Delta^{\gamma\delta)}$$

by taking projection  $\Delta_{\alpha\beta} \Delta_{\gamma\delta} J^{\alpha\beta\gamma\delta} \rightarrow C' = \frac{1}{5} \int dK (\Delta_{\alpha\beta} k^\alpha k^\beta)^2 f_{eq} \tilde{f}_{eq}$

- This is usually written as

$$C' = 3 J_{42}$$

- Definition  $J_{nq} = \frac{1}{(2q+1)!!} \int dK E_k^{n-2q} (-\Delta^{\alpha\beta} k_\alpha k_\beta)^q f_{eq} \tilde{f}_{eq} \quad (3.4)$

$\Rightarrow \pi^{\mu\nu} = 2 J_{42} \varepsilon^{\langle\mu\nu\rangle}$   $\hookrightarrow$  if  $\tilde{f}_{eq} = 1$   $J_{nq} \rightarrow I_{nq}$

- If bulk viscosity and diffusion can be neglected  $\rightarrow$  often used approximation in heavy-ion phenomenology

$$\delta f = \frac{f_{eq}}{2 J_{42}} \pi^{\mu\nu} k_{\langle\mu} k_{\nu\rangle} \quad (3.5)$$

- We introduced 14-moment approximation

$$\phi = \varepsilon + \varepsilon^\mu k_\mu + \varepsilon^{\mu\nu} k_\mu k_\nu$$

- By matching to  $T^{\mu\nu}$  and  $N^\mu$  we can write  $\varepsilon$ 's in terms of dissipative quantities  $\Pi, V^\mu, \pi^{\mu\nu}$
- So far no dynamics, but we can see that time-evolution of  $T^{\mu\nu}$  and  $N^\mu$  give also time-evolution of  $F_k$  (in 14 moment approximation)
- We can generalize 14-moment approximation to any number of moments.
- For this purpose we modify the expansion basis somewhat

$$1, k^\mu, k^\mu k^\nu, k^\mu k^\nu k^\lambda, \dots \rightarrow 1, k^{\langle\mu}, k^{\langle\mu\nu}, k^{\langle\mu\nu\lambda}, \dots \quad (3.6)$$

where tensors  $k^{\langle\mu_1 \dots \mu_m\rangle} = \Delta^{\mu_1 \dots \mu_m \nu_1 \dots \nu_m} k_{\nu_1} \dots k_{\nu_m}$  are symmetric, orthogonal, and traceless that satisfy orthogonality condition

$$\int dK F_k k^{\langle\mu_1 \dots \mu_m\rangle} k^{\langle\nu_1 \dots \nu_n\rangle} = \frac{m! \delta_{mn}}{(2m+1)!!} \Delta^{\mu_1 \dots \mu_m \nu_1 \dots \nu_m} \int dK F_k (\Delta^{\alpha\beta} k_\alpha k_\beta)^m \quad (3.7)$$

where  $F_k$  is an arbitrary function of  $E_k$  ( $F_k$  isotropic in LRF)

- Using this basis we can write  $f_k$  as

$$f_k = f_{eq} \left[ 1 + \frac{\tilde{f}_{eq}}{f_{eq}} \sum_{l=0}^{\infty} \sum_{n=0}^{N_l} \mathcal{H}_n^{(l)} \rho_n^{\mu_1 \dots \mu_l} k_{\langle \mu_1} \dots k_{\mu_l \rangle} \right] \quad (3.8)$$

where  $\mathcal{H}_n^{(l)}$  is polynomial in  $E_k$

- Irreducible tensors  $\rho_n^{\mu_1 \dots \mu_l}$  are defined as

$$\rho_n^{\mu_1 \dots \mu_l} = \int dK E_k^n k^{\langle \mu_1} \dots k^{\mu_l \rangle} \delta f \quad (3.9)$$

- These are independent of momentum  $\vec{k}$  and some of them can be identified as dissipative fluid variables

$$\rho_0 = -\frac{3\pi}{m^2} \quad \rho_0^\mu = V^\mu \quad \rho_1^\mu = W^\mu \stackrel{\text{Landau } u^\mu}{=} 0, \quad \rho_0^{\mu\nu} = \pi^{\mu\nu} \quad \left\| \begin{array}{l} e_1 n \text{-matching:} \\ \rho_1 = \rho_2 = 0 \end{array} \right. \quad (3.10)$$

- Once we have truncated the expansion (3.8) for any number of terms, the polynomials  $\mathcal{H}_n^{(l)} = \sum_{i=0}^l a_{ni}^{(l)} E_k^i$  can be determined by requiring consistency between (3.8) and (3.9).

- Alternatively it is possible to use orthogonal polynomials  $\rho_n^{\mu_1 \dots \mu_l} \rightarrow C_n^{\mu_1 \dots \mu_l} = \int dK P_n^{(l)} k^{\langle \mu_1} \dots k^{\mu_l \rangle} \delta f$   
 $\mathcal{H}_n^{(l)} \rightarrow P_n^{(l)}$ , where  $P_n^{(l)}$  is set of orthogonal polynomials (not used here)

- What we have now achieved is that we have written  $f_k$  in terms of (infinite) set of macroscopic quantities  $\int_n^{\mu_1 \dots \mu_\ell}$
- Once we know dynamics of  $\left\{ \int_n^{\mu_1 \dots \mu_\ell} \right\} \Rightarrow$  dynamics of  $f_k$
- Remaining task is to write equations of motion for  $\int_n^{\mu_1 \dots \mu_\ell}$
- Write co-moving derivative of  $\int_n^{\mu_1 \dots \mu_\ell}$  :

$$\Delta_{\nu_1 \dots \nu_\ell}^{\mu_1 \dots \mu_\ell} \frac{d}{d\tau} \int_n^{\nu_1 \dots \nu_\ell} = \dot{\int}^{\langle \mu_1 \dots \mu_\ell \rangle} = \Delta_{\nu_1 \dots \nu_\ell}^{\mu_1 \dots \mu_\ell} \frac{d}{d\tau} \int dK E_k^n k^{\langle \nu_1 \dots \nu_n \rangle} \delta f_k \quad (3.11)$$

- Using decompositions  $f_k = f_{eq} + \delta f$ ,  $k^\mu = E_k u^\mu + k^{\langle \mu \rangle}$ , and  $\partial_\mu = u_\mu \frac{d}{d\tau} + \nabla_\mu$  we can write the Boltzmann equation as

$$\dot{\delta f} = -\dot{f}_{eq} - E_k^{-1} k_\nu \nabla^\nu f_{eq} - E_k^{-1} k_\nu \nabla^\nu \delta f_k + E_k^{-1} C[f_k] \quad (3.12)$$

- Substituting this to (3.11) we obtain exact equations for  $\dot{\int}_n^{\langle \mu_1 \dots \mu_\ell \rangle}$
- This is somewhat long computation and here are only some steps as an example for rank-2 tensors.

$$\dot{S}_r^{(\mu\nu)} = \Delta_{\alpha\beta}^{\mu\nu} \frac{d}{d\tau} \int dK E_k^r k^{\langle\alpha} k^{\beta\rangle} \delta f = \Delta_{\alpha\beta}^{\mu\nu} \int dK \frac{d}{d\tau} (E_k^r k^{\langle\alpha} k^{\beta\rangle}) \delta f + \Delta_{\alpha\beta}^{\mu\nu} \int dK E_k^r k^{\langle\alpha} k^{\beta\rangle} \dot{\delta f} \quad (3.13)$$

\* substitute Boltzmann equation

$$= \Delta_{\alpha\beta}^{\mu\nu} \int dK \frac{d}{d\tau} (E_k^r k^{\langle\alpha} k^{\beta\rangle}) \delta f + \Delta_{\alpha\beta}^{\mu\nu} \int dK E_k^r k^{\langle\alpha} k^{\beta\rangle} \left[ \underset{1^{\circ}}{-\dot{f}_{eq}} - \underset{2^{\circ}}{E_k^{-1} k_{\lambda} \nabla^{\lambda} f_{eq}} - \underset{3^{\circ}}{E_k^{-1} k_{\lambda} \nabla^{\lambda} \delta f_k} + E_k^{-1} G[f] \right]$$

$$\underset{1^{\circ}}{=} \Delta_{\alpha\beta}^{\mu\nu} \int dK E_k^r k^{\alpha} k^{\beta} \dot{f}_{eq}$$

$$\dot{f}_{eq} = \frac{d}{d\tau} \left[ e^{\frac{E_k - \mu}{T} + a} \right]^{-1} = \frac{d}{d\tau} \left( \frac{E_k - \mu}{T} \right) \underbrace{e^{\frac{E_k - \mu}{T}} \left[ e^{(E_k - \mu)/T + a} \right]^{-2}}_{\lambda = k^{\lambda} u_{\lambda}}$$

$$= \left[ \frac{\dot{u}_{\gamma} k^{\gamma}}{T} - \frac{E_k}{T^2} \dot{T} - \frac{\dot{\mu}}{T} \right] f_{eq} \tilde{f}_{eq}$$

$$\begin{aligned} e^x [e^x + a]^{-2} &= \frac{e^x + a - a}{[e^x + a]^2} = \\ &= [e^x + a]^{-1} [1 - a [e^x + a]^{-1}] \\ &= f_{eq} \tilde{f}_{eq} \end{aligned}$$

$$\Delta_{\alpha\beta}^{\mu\nu} \dot{u}_{\gamma} \frac{1}{T} \int dK E_k^r k^{\gamma} k^{\alpha} k^{\beta} f_{eq} \tilde{f}_{eq} - \Delta_{\alpha\beta}^{\mu\nu} \frac{\dot{T}}{T^2} \int dK (-)$$

$$\underbrace{A' u^{\alpha} u^{\beta} u^{\gamma} + B' u^{\gamma} \Delta^{\alpha\beta}}_{\downarrow \quad \downarrow}$$

$\hookrightarrow 0$  (similarly)

$$\Rightarrow \Delta_{\alpha\beta}^{\mu\nu} \int dK E_k^r k^{\alpha} k^{\beta} \dot{f}_{eq} = 0$$

(3.14)

2<sup>o</sup> :  $\Delta_{\alpha\beta}^{\mu\nu} \int dK E_k^r k^{\langle\alpha} k^{\beta\rangle} \left[ E_k^{-1} k_\lambda \nabla^\lambda f_{eq} \right]$  (3.15)

↗ take this outside of integral

$$= \Delta_{\alpha\beta}^{\mu\nu} \nabla_\lambda \left( \underbrace{\int dK E_k^{r-1} k^\alpha k^\beta k^\lambda f_{eq}}_{\text{equilibrium moments}} - (r-1) \Delta_{\alpha\beta}^{\mu\nu} \nabla_\lambda u_\gamma \underbrace{\int dK E_k^{r-2} k^\alpha k^\beta k^\lambda k^\rho f_{eq}}_{\text{equilibrium moments}} \right)$$

decompose

$$\left\{ \begin{aligned} \int dK E_k^{r-1} k^\alpha k^\beta k^\gamma f_{eq} &= l_{r+2,0} u^\alpha u^\beta u^\gamma - 3 l_{r+2,1} u^{\langle\alpha} \Delta^{\beta\gamma\rangle} \\ \int dK E_k^{r-2} k^\alpha k^\beta k^\lambda k^\rho &= l_{r+2,0} u^\alpha u^\beta u^\lambda u^\rho - 6 l_{r+2,1} u^{\langle\alpha} u^\rho \Delta^{\beta\lambda\rangle} + 3 l_{r+2,2} \Delta^{\langle\beta} \Delta^{\lambda\rho\rangle} \end{aligned} \right.$$

• Only the last terms with  $l_{r+2,1}$  and  $l_{r+2,2}$  will survive  $\Delta_{\alpha\beta}^{\mu\nu}$

$$\Rightarrow \Delta_{\alpha\beta}^{\mu\nu} \int dK E_k^r k^{\langle\alpha} k^{\beta\rangle} \left[ E_k^{-1} k_\lambda \nabla^\lambda f_{eq} \right] = -2 \left[ l_{r+2,1} + (r-1) l_{r+2,2} \right] \underbrace{\Delta_{\alpha\beta}^{\mu\nu} \nabla^\alpha u^\beta}_{= \sigma^{\mu\nu}}$$

3<sup>o</sup>  $\Delta_{\alpha\beta}^{\mu\nu} \int dK E_k^{r-1} k^\alpha k^\beta G[f]$  (3.16)

$$= \Delta_{\alpha\beta}^{\mu\nu} \int dK E_k^{r-1} k^\alpha k^\beta \frac{1}{2} \int dK' dP dP' W_{kk' \rightarrow pp'} (f_p f_{p'} \tilde{f}_k \tilde{f}_{k'} - f_k f_{k'} \tilde{f}_p \tilde{f}_{p'})$$



- Linearizing the collision integral

$$\frac{1}{2} \int dK' dP dP' W_{kk' \rightarrow pp'} (f_p f_{p'} \tilde{f}_k \tilde{f}_{k'} - f_k f_{k'} \tilde{f}_p \tilde{f}_{p'}) \quad (3.17)$$

• use  $f_k = f_{eq}(1 + \tilde{f}_{eq} \phi_k)$  and keep only terms 1st order in  $\phi$

$$f_p f_{p'} = f_{eq,p} f_{eq,p'} (1 + \tilde{f}_{eq,p'} \phi_{p'} + \tilde{f}_{eq,p} \phi_p) + O(\phi^2)$$

$$\tilde{f}_p \tilde{f}_{p'} = \tilde{f}_{eq,p} \tilde{f}_{eq,p'} (1 - \alpha f_{eq,p'} \phi_{p'} - \alpha f_{eq,p} \phi_p) + O(\phi^2)$$

- Using further

$$\tilde{f}_{eq,p} = f_{eq,p} \exp\left[\frac{E_p - \mu}{T}\right] \quad (3.18)$$

$$f_{eq,p} f_{eq,p'} \tilde{f}_{eq,k} \tilde{f}_{eq,k'} = f_{eq,k} f_{eq,k'} \tilde{f}_{eq,p} \tilde{f}_{eq,p'}$$

we can write the linearized collision integral in the form

$$C[f] = \frac{1}{v} \int dK' dP dP' W_{kk' \rightarrow pp'} f_{eq,k} f_{eq,k'} \tilde{f}_{eq,p} \tilde{f}_{eq,p'} (\phi_p + \phi_{p'} - \phi_k - \phi_{k'}) + O(\phi^2) \quad (3.19)$$

- The collision term in the rank-2 e.o.m. is now

$$C_{r-1}^{<\mu\nu>} = \Delta_{\alpha\beta}^{\mu\nu} \int dK E_k^{r-1} k^\alpha k^\beta \frac{1}{2} \int dK' dP dP' W_{kk' \rightarrow pp'} f_{eq,k} f_{eq,k'} \tilde{f}_{eq,p} \tilde{f}_{eq,p'} (\phi_p + \phi_{p'} - \phi_k - \phi_{k'}) \quad (3.20)$$

- It turns out (details: arXiv:1202.4551) that we can write

$$C_{r-1}^{<\mu_1 \dots \mu_\ell>} = - \sum_{m=0}^{\infty} \mathcal{A}_{rn}^{(\ell)} \mathcal{S}_n^{\mu_1 \dots \mu_\ell} \quad (3.21)$$

where

$$\mathcal{A}_{rn}^{(\ell)} = \frac{1}{v(2\ell+1)} \int dK dK' dP dP' W_{kk' \rightarrow pp'} f_{eq,k} f_{eq,k'} \tilde{f}_{eq,p} \tilde{f}_{eq,p'} E_k^{r-1} k^{<\mu_1 \dots \mu_\ell>} \\ \times \left( \mathcal{H}_{k,n}^{(\ell)} k_{<\mu_1 \dots \mu_\ell>} + \mathcal{H}_{k',n}^{(\ell)} k'_{<\mu_1 \dots \mu_\ell>} - \mathcal{H}_{p,n}^{(\ell)} p_{<\mu_1 \dots \mu_\ell>} - \mathcal{H}_{p',n}^{(\ell)} p'_{<\mu_1 \dots \mu_\ell>} \right) \quad (3.22)$$

- Essential for our purpose is that the coefficients  $\mathcal{A}_{rn}^{(\ell)}$  contains all the details of the particle interactions

- In the rank-2 e.o.m. the collision term takes then a form

$$C_{r-1}^{<\mu\nu>} = - \sum_{m=0}^{\infty} \mathcal{A}_{rn}^{(2)} \mathcal{S}_n^{\mu\nu} \quad (3.23)$$

- Combining all the terms we finally get

$$\dot{S}_r^{\langle\mu\nu\rangle} = 2 [l_{r+2,1} + (r-1) l_{r+2,2}] \sigma^{\mu\nu} - \sum A_{rn}^{(2)} S_n^{\mu\nu} + \underbrace{\Delta_{\alpha\beta}^{\mu\nu} \int dk \frac{d}{dt} \left( E_k^r k^{\langle\alpha} k^{\beta\rangle} \right) \delta f - \Delta_{\alpha\beta}^{\mu\nu} \int dk E_k^n k^{\langle\alpha} k^{\beta\rangle} \frac{1}{E_k} k_\lambda \nabla^\lambda \delta f}_{\text{these lead to non-linear terms such as } S_r^{\mu\nu} \theta}$$

- This already resembles the Israel-Stewart equations for  $\pi^{\mu\nu}$ :

$$\dot{\pi}^{\langle\mu\nu\rangle} + \frac{1}{\tau} \pi^{\mu\nu} = \frac{2\eta}{\tau} \nabla^{\mu\nu} + \text{higher order terms,}$$

but is still a coupled equation for moments  $S_n^{\mu_1 \dots \mu_n}$

- In it's full glory the rank-2 equation is

$$\begin{aligned} \dot{S}_r^{\langle\mu\nu\rangle} - C_{r-1}^{\langle\mu\nu\rangle} &= 2\alpha_r^{(2)} \sigma^{\mu\nu} + \frac{2}{15} [(r-1)m^4 S_{r-2} - (2r+3)m^2 S_r + (r+4) S_{r+2}] \sigma^{\mu\nu} + \frac{2}{5} [r m^2 S_{r-1}^{\langle\mu} - (r+5) S_{r+1}^{\langle\mu}] \dot{u}^{\nu\rangle} \\ &+ r S_{r-1}^{\mu\nu\lambda} \dot{u}_\lambda - \frac{2}{5} \nabla^{\langle\mu} (m^2 S_{r-1}^{\nu\rangle} - S_{r+1}^{\nu\rangle}) + \frac{1}{3} [(r-1)m^2 S_{r-2}^{\mu\nu} - (r+4) S_r^{\mu\nu}] \theta \\ &+ \frac{2}{7} [(2r-2)m^2 S_{r-2}^{\lambda\langle\mu} - (2r+5) S_r^{\lambda\langle\mu}] \sigma_\lambda^{\nu\rangle} + 2 S_r^{\lambda\langle\mu} \omega_\lambda^{\nu\rangle} - \Delta_{\alpha\beta}^{\mu\nu} \nabla_\lambda S_{r-1}^{\alpha\beta\lambda} \\ &+ (r-1) S_{r-2}^{\mu\nu\lambda k} \nabla_{\lambda k} \end{aligned} \quad (3.24)$$

where  $\alpha_r^{(2)} = l_{r+2,1} - (r-1) l_{r+2,2}$

- Similar equations can be written for all  $\rho_n^{\mu_1 \dots \mu_\ell}$

- If we recall the expansion of  $f_k$ :

$$f_k = f_{eq} \left[ 1 + \frac{\tilde{z}}{f_{eq}} \sum_{\ell=0}^{\infty} \sum_{n=0}^{N_\ell} \mathcal{H}_n^{(\ell)} \rho_n^{\mu_1 \dots \mu_\ell} \langle k_{\mu_1} \dots k_{\mu_\ell} \rangle \right] \quad (3.25)$$

- Equations of motion for  $\rho_n^{\mu_1 \dots \mu_\ell}$  together with the expansion gives full dynamics of  $f_k$

- We are still no closer to fluid dynamics, but we have re-written the Boltzmann equation in easier form (although it may not appear so)

- The main complication is that the moment e.o.m. couple all the moments: Eventually the goal is to reduce this set of equations only to the dissipative quantities, e.g.  $\pi^{\mu\nu} = \rho_0^{\mu\nu}$

- For this purpose we need to be able to relate general moments  $\rho_n^{\mu_1 \dots \mu_\ell}$  to the dissipative quantities  $\pi^{\mu\nu}, v^M, \Pi$

## 14-moment approximation

• If we now recall the 14-moment approximation (truncation of expansion)

• Let's simplify a bit and assume that scalar (bulk viscosity) and vector moment (diffusion) can be neglected

$$\Rightarrow f_k = f_{eq} + f_{eq} \tilde{f}_{eq} \frac{1}{2J_{42}} \pi^{\mu\nu} k_{\langle\mu} k_{\nu\rangle} \quad (3.26)$$

$$\Rightarrow S_r = \int dK E_k^r \delta f = \underbrace{\int dK E_k^r f_{eq} \tilde{f}_{eq} k_{\langle\mu} k_{\nu\rangle}}_{=0} \frac{\pi^{\mu\nu}}{2J_{42}} = 0$$

$$S_r^\alpha = \int dK E_k^r \underbrace{k^{\langle\alpha\rangle} k_{\langle\mu} k_{\nu\rangle}}_{\rightarrow 0 \text{ orthogonality}} f_{eq} \tilde{f}_{eq} \frac{\pi^{\mu\nu}}{2J_{42}} = 0$$

$$S_r^{\alpha\beta} = \int dK E_k^r k^{\langle\alpha} k^{\beta\rangle} k_{\langle\mu} k_{\nu\rangle} f_{eq} \tilde{f}_{eq} \frac{\pi^{\mu\nu}}{2J_{42}} = \frac{J_{4+r,2}}{J_{42}} \pi^{\mu\nu} \leftarrow \text{every rank-2 moment is proportional to } \pi^{\mu\nu}$$

$$S_r^{\alpha\beta\gamma} = S_r^{\alpha\beta\gamma\delta} = \dots = 0$$

- If we now recall the eom, and take  $r=0$  (and assume further massless gas,  $m=0$ )

This is actually by Denicol, Koide & Rischke,  $r=2$  is original Israel & Stewart

$$\begin{aligned}
 \dot{S}_r^{\langle\mu\nu\rangle} - C_{r-1}^{\langle\mu\nu\rangle} &= 2\alpha_r^{(2)} \sigma^{\mu\nu} + \frac{2}{15} \left[ (r-1)m^4 \dot{S}_{r-2} - (2r+3)m^2 \dot{S}_r + (r+4) \dot{S}_{r+2} \right] \sigma^{\mu\nu} + \frac{2}{5} \left[ r m^2 \dot{S}_{r-1}^{\langle\mu} - (r+5) \dot{S}_{r+1}^{\langle\mu} \right] \dot{u}^{\nu\rangle} \\
 &+ r \dot{S}_{r-1}^{\mu\nu\lambda} \dot{u}_\lambda - \frac{2}{5} \nabla^{\langle\mu} \left( m^2 \dot{S}_{r-1}^{\nu\rangle} - \dot{S}_{r+1}^{\nu\rangle} \right) + \frac{1}{3} \left[ (r-1)m^2 \dot{S}_{r-2}^{\mu\nu} - (r+4) \dot{S}_r^{\mu\nu} \right] \theta \\
 &+ \frac{2}{7} \left[ (2r-2)m^2 \dot{S}_{r-2}^{\lambda\langle\mu} - (2r+5) \dot{S}_r^{\lambda\langle\mu} \right] \sigma_\lambda^{\nu\rangle} + 2 \dot{S}_r^{\lambda\langle\mu} \omega_\lambda^{\nu\rangle} - \Delta_{\alpha\beta}^{\mu\nu} \nabla_\lambda \dot{S}_{r-1}^{\alpha\beta\lambda} \\
 &+ (r-1) \dot{S}_{r-2}^{\mu\nu\lambda k} \nabla_{\lambda k}
 \end{aligned}$$

- We are left with

$$\dot{S}_0^{\langle\mu\nu\rangle} - C_{-1}^{\langle\mu\nu\rangle} = 2\alpha_0^{(2)} \sigma^{\mu\nu} - \frac{4}{3} \dot{S}_0^{\mu\nu} \theta - \frac{10}{7} \dot{S}_0^{\lambda\langle\mu} \sigma_\lambda^{\nu\rangle} + 2 \dot{S}_0^{\lambda\langle\mu} \omega_\lambda^{\nu\rangle} \quad (3.27)$$

- further  $\dot{S}_0^{\mu\nu} = \pi^{\mu\nu}$ ,  $C_{-1}^{\langle\mu\nu\rangle} = -\mathcal{A}_{00}^{(2)} \pi^{\mu\nu}$  ← only one term in  $f_k$  expansion

$$\Rightarrow \pi^{\langle\mu\nu\rangle} + \mathcal{A}_{00}^{(2)} \pi^{\mu\nu} = 2\alpha_0^{(2)} \sigma^{\mu\nu} - \frac{4}{3} \pi^{\mu\nu} \theta - \frac{10}{7} \pi^{\lambda\langle\mu} \sigma_\lambda^{\nu\rangle} + 2 \pi^{\lambda\langle\mu} \omega_\lambda^{\nu\rangle} \quad (3.28)$$

- These are the Israel-Stewart equations for  $\pi^{\mu\nu}$  with  $\tau = \left( \mathcal{A}_{00}^{(2)} \right)^{-1}$ ,  $\eta = \alpha_0^{(2)} \tau$   
↗  
massless

- If take full 14-moment approximation

$$p_n = \gamma_n^{(0)} \Pi$$

$$j_n^\mu = \gamma_n^{(1)} V^\mu$$

$$j_n^{\mu\nu} = \gamma_n^{(2)} \Pi^{\mu\nu}$$

(3.29)

where  $\gamma_n^{(i)}$ 's are thermodynamical coefficients (depend on  $T$  and  $\mu$ )

- Substituting these into the moment equations ( $j^{\langle \mu_1 \dots \mu_k \rangle}$ ) gives the full coupled Israel-Stewart equations of motion

$$\dot{\Pi} = -\frac{1}{\tau_\Pi} (\Pi + \theta) + \text{non-linear \& couplings}$$

$$\dot{V}^{\langle \mu \rangle} = -\frac{1}{\tau_V} \left( V^\mu - \kappa \nabla^\mu \left( \frac{\mu}{T} \right) \right) + \text{"}$$

(3.30)

$$\dot{\Pi}^{\langle \mu\nu \rangle} = -\frac{1}{\tau_\sigma} \left( \Pi^{\mu\nu} - 2\eta \nabla^{\langle \mu} \nabla^{\nu \rangle} \right) + \text{"}$$

• Is this good enough? What we want from fluid dynamical limit

◦ Evolution described by the conserved currents  $T^{\mu\nu}$  and  $N^\mu$  alone

◦ Fluid dynamics applicable when separation between microscopic and macroscopic scales

$$Kn = \frac{\lambda_{\text{micr}}}{L_{\text{macr}}} \lesssim 1 \quad (\text{Knudsen number}) \quad (3.31)$$

◦ microscopic scales  $\tau_\pi, \tau_v, \tau_\pi$  macroscopic scales  $\theta, \nabla^\mu \left( \frac{\mu}{T} \right), \frac{|\nabla^\mu e|}{e}$

$$\Rightarrow \text{e.g. } \tau_\pi \theta \lesssim 1$$

◦ Fluid dynamics applicable when when close to equilibrium.  
Quantify by inverse Reynolds numbers

$$R_\pi^{-1} = \frac{|\pi|}{P_{\text{eq}}} \ll 1 \quad R_v^{-1} = \frac{|v^\mu|}{n} \ll 1 \quad R_\pi^{-1} = \frac{|\pi^{\mu\nu}|}{P_{\text{eq}}} \ll 1 \quad (3.32)$$

◦ In Israel-Stewart type of fluid dynamics these are two independent type of quantities (related by e.o.m.)

◦ Want well-defined expansion in  $Kn$  and  $Re^{-1}$



- 14-moment approximation is a direct truncation of  $f_2^1$ 's expansion

e.g.  $\frac{\delta f}{\delta c_4} = \frac{1}{2J_{H_2}} \pi^{\mu\nu} k_\mu k_\nu$  + other dissipative quantities  $V^M$  &  $\Pi$  + neglect  $(\rho_1^{\mu\nu}, \rho_2^{\mu\nu}, \dots)$

- 14-moment approximation reduces the independent degrees of freedom to  $T^{\mu\nu}$  and  $N^M$ , but it is not truncation in  $K_n$  nor in  $\mathcal{R}^{-1}$

▷ in 14-mom :  $\rho_1^{\mu\nu} = \frac{J_{S_2}}{J_{H_2}} \pi^{\mu\nu} \Rightarrow$  same order in  $\mathcal{R}^{-1}$   
⏟  
thermodynamic function

- Let's look at how this can be done better.

- Back to the moment equation (rank-2 as an example)

$$\underbrace{\dot{\rho}_n^{(\mu\nu)} + \sum_{n=0}^{N_2} A_{rn}^{(2)} \rho_n^{\mu\nu}}_{\text{most important part when sufficiently close to equilibrium \& gradients small}} = 2\alpha_n^{(2)} \sigma^{\mu\nu} + \text{non-linear terms}$$

↳ gradient  $\times \rho$   
⏟  
these will lead to higher-order terms  $O(2)$  + higher

$O(kn)$  theory (Navier-Stokes):

- If we can neglect time-derivative & non-linear terms  $\rightarrow$

$$\sum_{n=0}^{N_L} \mathcal{A}_{nn}^{(2)} g_n^{\mu\nu} = 2\alpha_r^{(2)} \sigma^{\mu\nu} \quad (3.33)$$

- Define inverse of  $\mathcal{A}_{nn}^{(2)}$  :  $\sum_r \tau_{mr}^{(2)} \mathcal{A}_{nn}^{(2)} = \delta_{mn}$  (3.34)

$$\Rightarrow \sum_{rn} \tau_{mr}^{(2)} \mathcal{A}_{nn}^{(2)} g_n^{\mu\nu} = \sum_r 2\alpha_r \tau_{mr}^{(2)} \sigma^{\mu\nu} \quad (3.35)$$

$$\Rightarrow \delta_{mn} g_n^{\mu\nu} = \sum_r 2\alpha_r \tau_{mr}^{(2)} \sigma^{\mu\nu} \quad (3.36)$$

$$\stackrel{m=0}{\Rightarrow} g_0^{\mu\nu} = \pi^{\mu\nu} = 2 \underbrace{\sum_r \tau_{0r}^{(2)} \alpha_r^{(2)}} \sigma^{\mu\nu} \quad (3.37)$$

This can be identified as shear viscosity  $\eta$

- Shear viscosity given by the inverse of the collision matrix  $\mathcal{A}_{nn}^{(2)}$

$\triangle$  compare to 14-mom. approximation  $\eta = (\mathcal{A}_{00}^{(2)})^{-1} \alpha_0^{(2)}$   
 $\hookrightarrow$  just one term from the full matrix

- In order to get the full 1st-order theory we need to sum all moments  $\rho_n^{\mu\nu}$

△ 14-mom. approximation neglects infinitely many  $O(K^n)$  terms

- Same applies to the transient fluid dynamics (where  $2\eta\sigma^{\mu\nu}$  is one of the terms)

### Transient fluid dynamics

- Restore time derivative to moment eqs.

$$\dot{\rho}_r^{\langle\mu\nu\rangle} + \sum_{n=0}^{N_e} \mathcal{A}_{rn}^{(2)} \rho_n^{\mu\nu} = 2\alpha_r^{(2)} \sigma^{\mu\nu} + \underbrace{\text{non-linear terms}}_{\text{ignore these for now}} \quad (3.38)$$

- Linearized moment equation is now

$$\dot{\rho}_r^{\langle\mu\nu\rangle} + \sum_n \mathcal{A}_{rn}^{(2)} \rho_n^{\mu\nu} = 2\alpha_r^{(2)} \sigma^{\mu\nu} \quad (3.39)$$

- Diagonalize this system

- Define matrix  $\Omega^{(2)}$  that diagonalizes the collision matrix  $\mathcal{A}^{(2)}$

$$(\Omega^{(2)})^{-1} \mathcal{A}^{(2)} \Omega^{(2)} = \text{diag}(x_0^{(2)}, x_1^{(2)}, \dots, x_{N_e}^{(2)}) \quad (3.40)$$

- $x_i$  are the eigenvalues of  $\mathcal{A}^{(2)}$

- Define further rank-2 tensors

$$X_i^{\mu\nu} = \sum_j (\Omega^{(2)})^{-1}_{ij} g_j^{\mu\nu} \quad (3.41)$$

- these are the eigenmodes of (linearized) Boltzmann equation

$\Rightarrow$  Linearized moment equation:

$$\dot{g}_r^{\langle\mu\nu\rangle} + \sum_n \mathcal{A}_{rn}^{(2)} g_n^{\mu\nu} = 2\alpha_r^{(2)} \sigma^{\mu\nu} \quad \Big| \Omega^{-1} \times \quad (3.42)$$

$\wedge$   
 $\Omega \Omega^{-1}$

$$\Rightarrow \dot{X}_i^{\langle\mu\nu\rangle} + x_i^{(2)} X_i^{\mu\nu} = \beta_i^{(2)} \sigma^{\mu\nu} + \text{higher-order terms} \quad (3.43)$$

$\swarrow$  time derivative of  $\Omega$

$$\beta_i^{(2)} = 2 \sum_j (\Omega^{(2)})^{-1}_{ij} \alpha_j^{(2)}$$

$\Rightarrow$  Equations of motion for  $X_i^{\mu\nu}$  decouple (in linear regime)

$1/\chi_i^{(2)}$  are now relaxation times for tensors  $X_i^{\mu\nu}$  (eigenmodes of  $\mathcal{A}^{(2)}$ )

• Note that if the gradients  $\sigma^{\mu\nu} = 0$

$$\Rightarrow \dot{X}_i^{(\mu\nu)} + \chi_i^{(2)} X_i^{\mu\nu} = 0 \quad \Rightarrow \quad X_i^{\mu\nu} \xrightarrow{1/\chi_i^{(2)}} 0 \quad (3.44)$$

$\Rightarrow 1/\chi_i^{(2)}$  are thermalization times of the system

• If we wait long enough  $\Rightarrow$  slowest thermalization/relaxation time dominates!

• Order  $\chi_i^{(2)}$ 's such that  $1/\chi_0^{(2)}$  is the slowest time scale

• If we now assume that only the slowest mode is fully dynamical

$$\dot{X}_0^{(\mu\nu)} + \chi_0^{(2)} X_0^{\mu\nu} = \beta_0^{(2)} \sigma^{\mu\nu} + \text{higher-order} \quad (3.45)$$

and rest of the modes can be approximated as

$$X_r^{\mu\nu} = \frac{\beta_r^{(2)}}{\chi_r^{(2)}} \sigma^{\mu\nu} + \text{higher-order} \quad (\text{for } r \neq 0) \quad (3.46)$$

$\Rightarrow$  We have reduced the independent degrees of freedom to  $X_i^{\mu}$

• Good, but we still need to express everything in terms of  $\pi^{\mu\nu}$ ,  $V^{\mu}$  and  $\Pi$

• We can first invert  $X_i^{\mu\nu} = \sum_j (\Omega^{(2)})_{ij}^{-1} \rho_j^{\mu\nu}$

$$\Rightarrow \rho_i^{\mu\nu} = \sum_j \Omega_{ij}^{(2)} X_j^{\mu\nu} = \Omega_{i0}^{(2)} X_0^{\mu\nu} + \sum_{j=1} \Omega_{ij}^{(2)} \frac{\beta_j^{(2)}}{\chi_j^{(2)}} \sigma^{\mu\nu} = \Omega_{i0}^{(2)} X_0^{\mu\nu} + O(k^n) \quad (3.47)$$

• take  $i=0$ , so that  $\rho_0^{\mu\nu} = \pi^{\mu\nu}$ , and set  $\Omega_{00}^{(2)} = 1$

$$\Rightarrow X_0^{\mu\nu} = \pi^{\mu\nu} - \sum_{j=1} \Omega_{0j}^{(2)} \frac{\beta_j^{(2)}}{\chi_j^{(2)}} \sigma^{\mu\nu}$$

substitute

$$\Rightarrow \rho_i^{\mu\nu} \approx \Omega_{i0}^{(2)} \pi^{\mu\nu} - \sum_{j=1} \Omega_{i0}^{(2)} \Omega_{ij}^{(2)} \frac{\beta_j^{(2)}}{\chi_j^{(2)}} \sigma^{\mu\nu} + \sum_{j=1} \Omega_{ij}^{(2)} \frac{\beta_j^{(2)}}{\chi_j^{(2)}} \sigma^{\mu\nu} + \text{H.O.} \quad (3.48)$$

$$\parallel \text{ use: } \beta_i^{(2)} = 2 \sum_j (\Omega^{(2)})_{ij}^{-1} \alpha_j^{(2)}$$

$$\& \quad \tau_{in}^{(2)} = \sum_{m=0}^{N_L} \Omega_{im}^{(2)} \frac{1}{\chi_m^{(2)}} (\Omega^{-1})_{mn}^{(2)}$$

$\uparrow$   
invers of diagonal matrix  $(a \ b \ c \ \dots)$

$$= \begin{pmatrix} 1/a & & & \\ & 1/b & & \\ & & 1/c & \\ & & & \dots \end{pmatrix}$$

- Finally we get the relation

$$\beta_i^{\mu\nu} = \Omega_{i0}^{(2)} \Pi^{\mu\nu} + (2\eta_i - \Omega_{i0}^{(2)} \eta_0) \sigma^{\mu\nu} = \Omega_{i0}^{(2)} \Pi^{\mu\nu} + O(kn) \quad (3.49)$$

where  $\eta_i = \sum_{r=0}^i \tau_{ir}^{(2)} \alpha_r^{(2)}$ , so that  $\eta_0 = \eta = \text{shear viscosity}$

- These are now relations that can be substituted into the original moment equations (similar relations can be derived for  $\beta_i$  and  $\beta_i^T$  as well)

$$\begin{aligned} \dot{\beta}_r^{\langle\mu\nu\rangle} - C_{r-1}^{\langle\mu\nu\rangle} &= 2\alpha_r^{(2)} \sigma^{\mu\nu} + \frac{2}{15} [(r-1)m^4 \beta_{r-2} - (2r+3)m^2 \beta_r + (r+4)\beta_{r+2}] \sigma^{\mu\nu} + \frac{2}{5} [r m^2 \beta_{r-1}^{\langle\mu} - (r+5)\beta_{r+1}^{\langle\mu} ] \dot{u}^{\nu\rangle} \\ &+ r \beta_{r-1}^{\mu\nu\lambda} \dot{u}_\lambda - \frac{2}{5} \nabla^{\langle\mu} (m^2 \beta_{r-1}^{\nu\rangle} - \beta_{r+1}^{\nu\rangle}) + \frac{1}{3} [(r-1)m^2 \beta_{r-2}^{\mu\nu} - (r+4)\beta_r^{\mu\nu}] \theta \\ &+ \frac{2}{7} [(2r-2)m^2 \beta_{r-2}^{\lambda\langle\mu} - (2r+5)\beta_r^{\lambda\langle\mu} ] \sigma_\lambda^{\nu\rangle} + 2\beta_r^{\lambda\langle\mu} \omega_\lambda^{\nu\rangle} - \Delta_{\alpha\beta}^{\mu\nu} \nabla_\lambda \beta_{r-1}^{\alpha\beta\lambda} \\ &+ (r-1) \beta_{r-2}^{\mu\nu\lambda k} \nabla_{\lambda k} \end{aligned} \quad (3.50)$$

• Note that there are some moments  $g_i^{\mu\nu}$  in the equations for which  $i \neq 0$ , and above relation is only for  $i=0$

$\Rightarrow$  We can use  $g_i^{\mu\nu} = \int d^4k E^{-i} k^{(\mu} k^{\nu)} \delta f$  (3.51)

↑  
substitute expansion  
with  $g_i^{\mu\nu}$ 's,  $i \neq 0$  into here

• The form of the final equation for  $\pi^{\mu\nu}$  is now

$$\tau_\pi \dot{\pi}^{\langle\mu\nu\rangle} + \pi^{\mu\nu} = 2\eta \sigma^{\mu\nu} + \mathcal{J}^{\mu\nu} + K^{\mu\nu} + R^{\mu\nu}, \quad (3.52)$$

where  $\tau_\pi = 1/\chi_{(0)}$ ,  $\eta = \sum_r \tau_{or}^{(2)} \alpha_r^{(2)}$

•  $\mathcal{J}^{\mu\nu}$  contains terms of order  $K_n \times R^{-1}$

$$\mathcal{J}^{\mu\nu} = 2\pi_{\lambda}^{\langle\mu} \omega^{\nu\rangle\lambda} - \delta_{\pi\pi} \pi^{\mu\nu} \Theta - \tau_{\pi\pi} \pi^{\lambda\langle\mu} \sigma^{\nu\rangle\lambda} + \lambda_{\pi\pi\pi} \pi \sigma^{\mu\nu} - \tau_{\pi\nu} V^{\langle\mu} \nabla^{\nu\rangle} p_{eq} + \eta_{\pi\nu} \nabla^{\langle\mu} V^{\nu\rangle} + \lambda_{\pi\pi} V^{\langle\mu} \nabla^{\nu\rangle} \frac{\mu}{\tau}$$

•  $K^{\mu\nu}$  terms  $K_n^2$  (absent completely in Israel-Stewart theory)

$$K^{\mu\nu} = \eta_1 \omega_{\lambda}^{\langle\mu} \omega^{\nu\rangle\lambda} + \eta_2 \theta \sigma^{\mu\nu} + \eta_3 \sigma^{\lambda\langle\mu} \sigma^{\nu\rangle\lambda} + \eta_4 \sigma_{\lambda}^{\langle\mu} \omega^{\nu\rangle\lambda} + \eta_5 \nabla^{\langle\mu} \left(\frac{\mu}{\tau}\right) \nabla^{\nu\rangle} \left(\frac{\mu}{\tau}\right) + \eta_6 \nabla^{\langle\mu} p_{eq} \nabla^{\nu\rangle} p_{eq} + \eta_7 \nabla^{\langle\mu} \left(\frac{\mu}{\tau}\right) \nabla^{\nu\rangle} p_{eq} + \eta_8 \nabla^{\langle\mu} \nabla^{\nu\rangle} \left(\frac{\mu}{\tau}\right) + \eta_9 \nabla^{\langle\mu} \nabla^{\nu\rangle} p_{eq}$$



- $R^{\mu\nu}$  contains  $(R^{-1})^2$  that come from the non-linear part of the collision integral

$$R^{\mu\nu} = g_6 \Pi \sigma^{\mu\nu} + g_7 \Pi^{\lambda \langle \mu} \Pi_{\lambda}^{\nu \rangle} + g_8 V^{\langle \mu} V^{\nu \rangle}$$

- Every type of term that can appear will appear
- These are the DNMR equations. (arXiv:1202.4551)

(all the moments summed into coefficients, slowest microscopic time identified as relaxation time,  $k_n$  and  $R^{-1}$  identified as separate quantities)

• 14-mom. approximation

$$\frac{\delta f}{f_k} = \varepsilon + \varepsilon^\mu k_\mu + \varepsilon^{\mu\nu} k_\mu k_\nu$$

$$\text{full ex.} \implies f_k = f_{eq} \left[ 1 + \sum_{l=0}^{\infty} \sum_{n=0}^{N_L} \mathcal{H}_n^{(l)} \left( \frac{k_{\mu_1} \dots k_{\mu_l}}{k_{\mu_1} \dots k_{\mu_l}} \right) \right]$$

e.g.  $\mathcal{H}_n^{\mu\nu} = \Omega_{io}^{(2)} \Pi^{\mu\nu} + O(k_n)$

- Note:  $\frac{g_i^{\mu\alpha}}{g_i^{\mu\nu\alpha\beta}} \sim O(2)$ , first order term doesn't exist

- Problem are  $Kn^2$  terms & contains e.g.  $(\partial_t)^2$  terms  
 $\Rightarrow$  Potentially ruin causality of the theory

### Other approaches

- In DNMR approach

$$S_r^{\mu\nu} = \Omega_{rs}^{(2)} \pi^{\mu\nu} + O(Kn)$$

$\uparrow$   
 This is the reason for  $O(Kn^2)$  terms

- Instead of trying to identify relaxation time as the real slowest timescale of BE, assume that we are near asymptotic regime where

$$S_i^{\mu\nu} = 2\eta_i \sigma^{\mu\nu} + O(2)$$

$\hat{=}$  "shear viscosity of moment  $S_i^{\mu\nu}$ ,  $\eta_0 = \eta$

$$\Rightarrow \left. \begin{array}{l} \pi^{\mu\nu} = 2\eta \sigma^{\mu\nu} + O(2) \\ S_i^{\mu\nu} = 2\eta_i \sigma^{\mu\nu} + O(2) \end{array} \right\} \Rightarrow \underline{S_i^{\mu\nu} = \frac{\eta_i}{\eta} \pi^{\mu\nu} + O(2)}$$

- Order-of-magnitude approximation (Fotakis talk)

- We can then substitute  $\rho_i^\mu = \frac{\eta_i}{\eta} \pi^\mu + O(2)$  into moment equations  
 $\downarrow$   
 Can be neglected as gives at least  $O(3)$  terms in eom's
- Nice feature is that  $O(Kn^2)$  terms are now absent.
- Relaxation time no longer slowest timescale of DE, but some effective timescale
- $Re^{-1}$  and  $Kn$  are not anymore completely independent, but are of the same order

$$\underbrace{\frac{\pi^\mu}{\rho_{eq}}}_{R^{-1}} \sim \underbrace{\frac{2\eta}{\rho_{eq}} \sigma^\mu}_{Kn} + O(2)$$

- In principle there is no such restriction in DNMR, as long as  $Kn$  and  $R^{-1}$  sufficiently small

- It is possible to have both identification of slowest timescale and absence of explicit  $Kn^2$  terms by doubling d.o.f

$$S_r^{\mu\nu} = \#_{\alpha_1} \Pi^{\mu\nu} + \#_{\alpha_1} \rho_1^{\mu\nu} + O(Kn^2)$$

↑  
new degree of freedom in fluid dynamics

$$\Rightarrow \begin{cases} \dot{\Pi}^{\langle\mu\nu\rangle} + \tau_{00} \Pi^{\mu\nu} + \tau_{\alpha_1} \rho_1^{\mu\nu} = 2\alpha^{(0)} \nabla^{\mu\nu} \\ \dot{\rho}_1^{\langle\mu\nu\rangle} + \tau_{\phi 0} \Pi^{\mu\nu} + \tau_{11} \rho_1^{\mu\nu} = 2\alpha^{(1)} \nabla^{\mu\nu} \end{cases} \left. \begin{array}{l} + \text{higher-order} \\ + \text{higher-order} \end{array} \right\} \rightarrow \text{no } O(Kn^2) \text{ terms}$$

△ This is in good agreement with direct numerical solutions of BE (mere transport coef. & mere initial conditions)

Bouras et al, arXiv: 1207.6811



$$\bullet N_B^M = \sum_i B_i N_i^M = n_B u^M + V_B^M \quad \& N_S^M \quad \text{and} \quad N_Q^M$$

$\uparrow$   
 matching  $n_{B,eq} = n_B = \sum_i B_i (n_i + \delta n_i)$  ,  $\sum_i B_i \delta n_i = 0$  ,  $\delta n_i \neq 0$

$$\bullet n_{B,eq}(T, \mu_B, \mu_S, \mu_Q) = n_B \quad n_{S,eq}(T, \mu_B, \mu_S, \mu_Q) = n_S \quad n_{Q,eq}(T, \mu_B, \mu_S, \mu_Q) = n_Q$$

$$\& \mathcal{E}_{eq}(T, \mu_B, \mu_S, \mu_Q) = \mathcal{E}$$



$$\bullet \text{Single fluid with } \underbrace{u^M, T, \{\mu_i = B_i \mu_B + S_i \mu_S + Q_i \mu_Q\}}_{\text{common } u^M \& T}$$