

# Non-perturbative constraints on perturbation theory at finite temperature

(PL, O. Philipsen, 2405.02009)

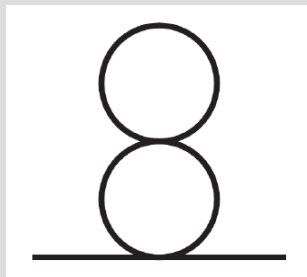
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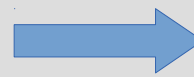
# 1. Perturbation theory subtleties for $T > 0$

- It has long been understood that the inclusion of temperature  $T > 0$  introduces various complications:
  - QCD perturbation theory suffers inconsistencies above a fixed loop order (“Linde problem” [PLB 96, 1980])
  - Perturbative series have poor convergence properties: need to reorganise the expansion (screened perturbation theory, infinite resummations, functional techniques, ...)
- All of these problems stem from the fact that  $T > 0$  perturbation theory introduces a class of diagrams that are not present in the vacuum theory, and these are highly sensitive to the infrared dynamics

e.g. 2-loop order  $\phi^4$  theory:



Includes sub-component



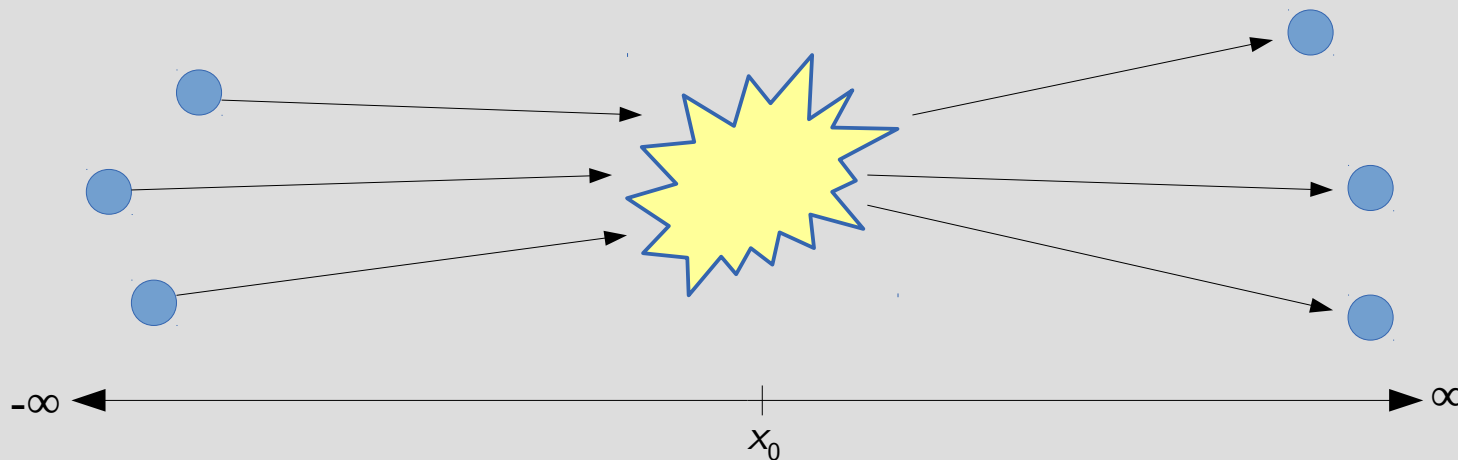
$$T \sum_n \int \frac{d^3 p}{(2\pi)^3} \frac{1}{(\omega_n^2 + |\vec{p}|^2 + m^2)^2}$$

Diverges for  $m \rightarrow 0$ , even after renormalisation

# 1. Perturbation theory subtleties for $T > 0$

- These features are often viewed as a purely perturbative feature of finite-temperature QFT, but they may in fact be a symptom of a more fundamental (non-perturbative) constraint:

“Narnhofer-Requardt-Thirring Theorem” [Commun. Math. Phys. 92, 247 (1983)]



Idea: construct thermal “quasi-particle” in/out scattering states with real dispersion relations  $\omega = E(\mathbf{p}) \rightarrow$  Conclusion: *the S-matrix must be trivial!*

- So quasi-particles can exist, but only when there are no interactions...

# 1. Perturbation theory subtleties for $T > 0$

→ **What's the reason for this?**

- QFT: *thermal states satisfy the KMS condition, and this gives rise to very different spectral constraints than in the vacuum case*
- Physics: *Dissipative effects of the thermal medium are everywhere-present, so need to take these into account in the definition of scattering states!*
- This has significant implications for perturbation theory: *neither free field, nor quasi-particle propagators with real poles  $\omega = E(\mathbf{p})$ , can form the basis of finite-temperature perturbative expansions* [Landsman, *Ann. Phys.* 186, 141 (1988)]
- Note: these constraints are not restricted to infrared regimes, but also affect systems at low temperatures, or with non-vanishing mass scales (i.e.  $T/m \leq 1$ )

# 1. Perturbation theory subtleties for $T > 0$

→ Is there evidence for these constraints?

Yes! Weldon [*PRD* 65 (2002)] showed that the standard perturbative procedure for scalar theories has inconsistencies

- Can classify the singularities occurring in the n-loop perturbative propagator given the singularities of the basic propagator
- If the basic propagators have real poles  $\omega = E(\mathbf{p})$ , at some loop order the perturbative propagator will develop a branch point singularity at these poles, e.g. in  $\Phi^4$  theory this occurs at 2-loop order
- This prohibits perturbative corrections to the propagator pole from being computed, hence the standard approach breaks down!

This is a generic feature of *any* perturbative computation that uses propagators with real poles

$$G_R(p) = -\frac{1}{(\omega + i\epsilon)^2 - E(\vec{p})^2}$$

## 2. Investigation of $\phi^4$ theory on the lattice

- In principle, the implications of these constraints can also be looked for by comparing perturbative predictions with lattice calculations

Idea [PL, O. Philipsen, 2405.02009]: perform lattice simulations of basic field correlators in  $\phi^4$  theory and compare these with the predictions of lattice perturbation theory (LPT) → see: Montvay & Münster

- In LPT one uses the discretised propagator

$$G(p; a, L, L_\tau) = \frac{1}{\sum_\mu \frac{4}{a^2} \sin^2\left(\frac{ap_\mu}{2}\right) + m_0^2}$$

- Theory is defined on a finite  $N_s^3 \times N_\tau$  lattice → internal loop integrals are replaced with sums over first Brillouin zone

$$\mathcal{B}_a = \left\{ p \in \mathbb{R}^4 : -\frac{\pi}{a} < p_\mu \leq \frac{\pi}{a} \right\}$$

- In  $\phi^4$  theory the lattice action is:

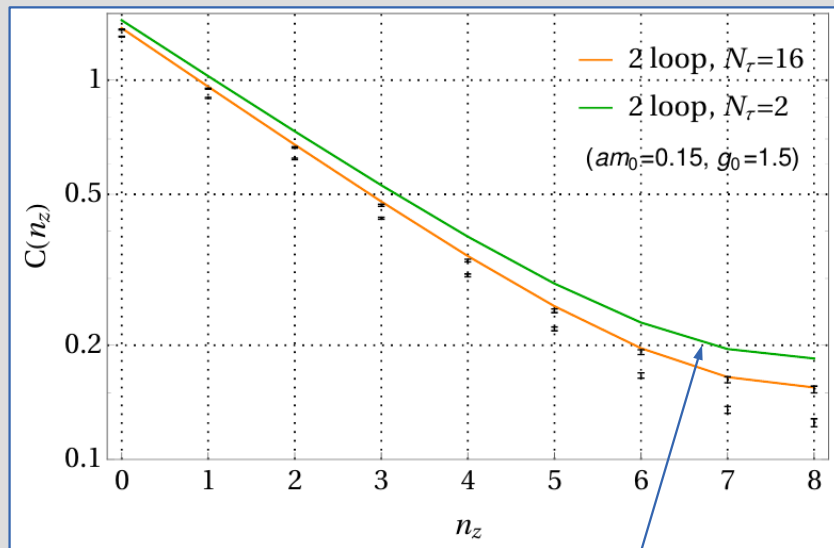
$$S = a^4 \sum_{x \in \Lambda} \left[ \frac{1}{2} \Delta_\mu^f \phi_0(x) \Delta_\mu^f \phi_0(x) + \frac{m_0^2}{2} \phi_0(x)^2 + \frac{g_0}{4!} \phi_0(x)^4 \right]$$

- Here we focus on perturbative calculations of the *spatial correlator*, which is defined:

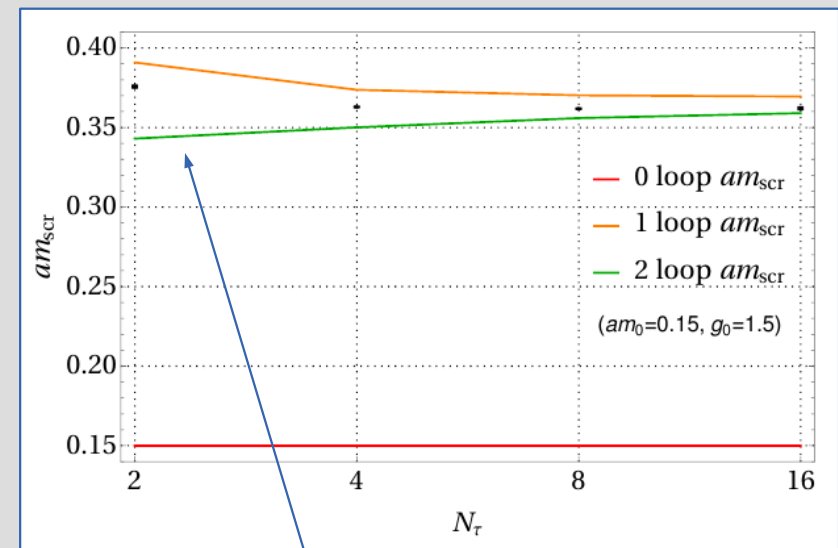
$$C(z; a, N_s, N_\tau) = a^3 \sum_{\tau, x, y} \langle \phi(\tau, \vec{x}) \phi(0) \rangle.$$

## 2. Investigation of $\phi^4$ theory on the lattice

- How do the 2-loop LPT predictions compare with the lattice data as one varies the (lattice) temperature  $T=1/(aN_\tau)$ ?
  1. For sufficiently small  $g_0$  the perturbative predictions are consistent with the data for all temperatures
  2. For non-negligible  $g_0$  the perturbative predictions for  $C(z)$  increasingly deteriorate as  $T/m$  increases



For  $T/m \sim 1$  the 2-loop predictions deviate drastically from the data



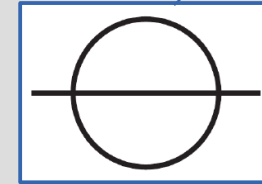
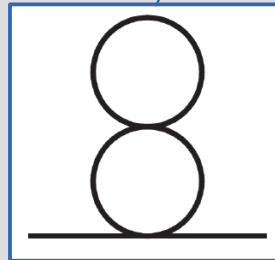
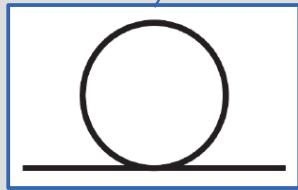
These deviations are also reflected by the screening mass  $m$ , defined:  $C(z) \sim e^{-mz}$ , having the *opposite*  $T$  dependence

## 2. Investigation of $\Phi^4$ theory on the lattice

- To establish the origin of these deviations one needs to understand how the 2-loop perturbative predictions are computed

$$C(z; a, N_s, N_\tau) = \frac{1}{N_s} \sum_{k_z=0}^{N_s-1} e^{\frac{2\pi i k_z}{a N_s} z} \frac{a}{4 \sin^2\left(\frac{\pi k_z}{N_s}\right) + (am_0)^2 + a^2 \Pi(\omega_E = p_x = p_y = 0, p_z = \frac{2\pi k_z}{a N_s}; a, N_s, N_\tau)}.$$

$$a^2 \Pi_{(2)}(p; a, N_s, N_\tau) = \frac{g_0}{2} J_1(am_0; N_s, N_\tau) - \frac{g_0^2}{4} J_1(am_0; N_s, N_\tau) J_2(am_0; N_s, N_\tau) - \frac{g_0^2}{6} I_3(p, am_0; N_s, N_\tau)$$



The functions  $J_n$   
and  $I_3$  are defined:

$$J_n(am_0; N_s, N_\tau) = \frac{1}{N_s^3 N_\tau} \sum_{p \in \mathcal{B}_a} \frac{1}{\left[ \sum_{\mu} 4 \sin^2\left(\frac{ap_{\mu}}{2}\right) + (am_0)^2 \right]^n}, \quad n = 1, 2$$

$$I_3(p, am_0; N_s, N_\tau) = \frac{1}{(N_s^3 N_\tau)^2} \sum_{q \in \mathcal{B}_a} \sum_{r \in \mathcal{B}_a} \frac{1}{\left[ \sum_{\mu} 4 \sin^2\left(p_{\mu} - \frac{aq_{\mu}}{2} - \frac{ar_{\mu}}{2}\right) + (am_0)^2 \right]} \times \frac{1}{\left[ \sum_{\nu} 4 \sin^2\left(\frac{aq_{\nu}}{2}\right) + (am_0)^2 \right]} \frac{1}{\left[ \sum_{\Lambda} 4 \sin^2\left(\frac{ar_{\Lambda}}{2}\right) + (am_0)^2 \right]}$$



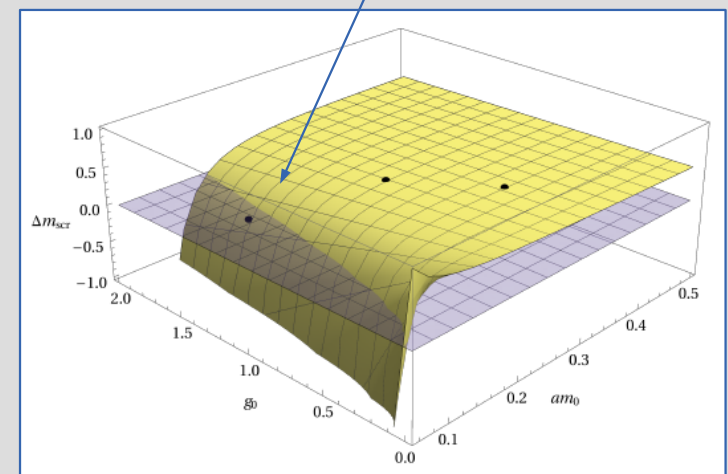
## 2. Investigation of $\phi^4$ theory on the lattice

- Numerically, one finds that the sunset diagram contribution is sub-dominant, and that the behaviour of  $C(z)$  is largely controlled by the interplay between the tadpole and cactus diagrams
- As one varies  $(am_0, g_0)$  there are certain choices for which the self-energy at any fixed  $N_\tau$  is **smaller** than when  $N_\tau \rightarrow \infty$  (i.e.  $T = 0$ )
  - This implies that the screening mass  $m_{\text{scr}}$ , which is defined via the (continuum) large- $z$  behaviour:  $C(z) \sim \exp(-m_{\text{scr}} z)$ , **decreases** with  $T$

$$\Delta m_{\text{scr}} = \frac{\frac{g_0}{2} J_1(am_0; N_s, N_\tau = 1) - \frac{g_0^2}{4} J_1(am_0; N_s, N_\tau = 1) J_2(am_0; N_s, N_\tau = 1)}{\frac{g_0}{2} J_1(am_0; N_s, N_\tau = \infty) - \frac{g_0^2}{4} J_1(am_0; N_s, N_\tau = \infty) J_2(am_0; N_s, N_\tau = \infty)} - 1,$$

- When  $\Delta m_{\text{scr}} < 0$ , and the sunset is sufficiently small, then:  $m_{\text{scr}}(T=0) > m_{\text{scr}}(T>0)$
- This behaviour is **opposite** to physical expectations (and the lattice data!)

Predictions with  $(am_0, g_0)$  closer to the region  $\Delta m_{\text{scr}} < 0$  deviate further from the data



## 2. Investigation of $\phi^4$ theory on the lattice

- The proximity to the region  $\Delta m_{\text{scr}} < 0$  depends on the structure of the propagators entering the loop calculations
- From the plot one can see that for fixed  $g_0$  one will *always* move into this region when  $am_0$  is sufficiently small
  - This can be understood by the fact that for small  $am_0$  the zero-mode contributions dominate the sums in  $J_n$ , and these have the form:

$$J_1(am_0; N_s, N_\tau) \sim \frac{1}{N_s^3 N_\tau} \frac{1}{(am_0)^2}, \quad J_2(am_0; N_s, N_\tau) \sim \frac{1}{N_s^3 N_\tau} \frac{1}{(am_0)^4}.$$

- The cactus diagram outcompetes the tadpole for small enough  $am_0$

Conclusion: the deviations in the two-loop perturbative predictions are driven by the analytic structure of the free field propagators → this is consistent with the constraints of the NRT theorem and the findings of Weldon [PRD 65 (2002)]

### 3. A potential resolution

- The NRT constraints and previous results suggest that one needs an alternative approach to scalar perturbation theory when  $T > 0$

Idea: Start with propagators that consistently take thermal effects into account from the outset, i.e. non-real singularities

→ *But what form should these propagators take?*

- There's theoretical evidence to suggest that the thermal medium contains discrete particle-like excitations [Bros, Buchholz, *NPB* 627 (2002)]

$$\rho(\omega, \vec{p}) = \int_0^\infty ds \int \frac{d^3 \vec{u}}{(2\pi)^2} \epsilon(\omega) \delta(\omega^2 - (\vec{p} - \vec{u})^2 - s) \tilde{D}_\beta(\vec{u}, s) \longrightarrow \tilde{D}_{m,\beta}(\vec{u}) \delta(s - m^2)$$

“Thermoparticle”

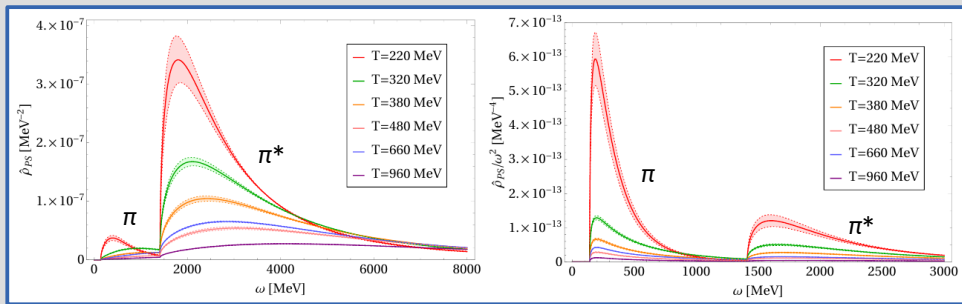
- Thermoparticles reduce to those of a vacuum particle with mass  $m$  when  $T \rightarrow 0$
- “Damping factor”  $\tilde{D}_\beta(\mathbf{u})$  results in thermally-broadened peaks in the spectral function: this parametrises the effects of collisional broadening

# 3. A potential resolution

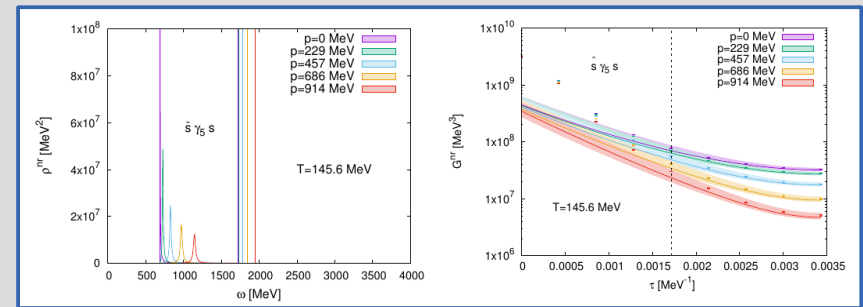
- Thermoparticle components have distinctive properties:

1. They dominate the large-time behaviour of  $\langle \phi(x_0, \vec{x}) \phi(0) \rangle$
2. Along the direction  $\mathbf{x} = \mathbf{v} x_0$  they behave like:  $D_{m,\beta}(\vec{v}x_0) |x_0|^{-3/2}$  so their damped behaviour depends on their velocity [Bros, Ann. H. Poincare 4, (2003)]
3. They evade the consequences of the NRT theorem since they no longer have a sharp dispersion law

→ There's also (mounting) evidence from lattice QCD data that these excitations are present in various (meson) spectral functions



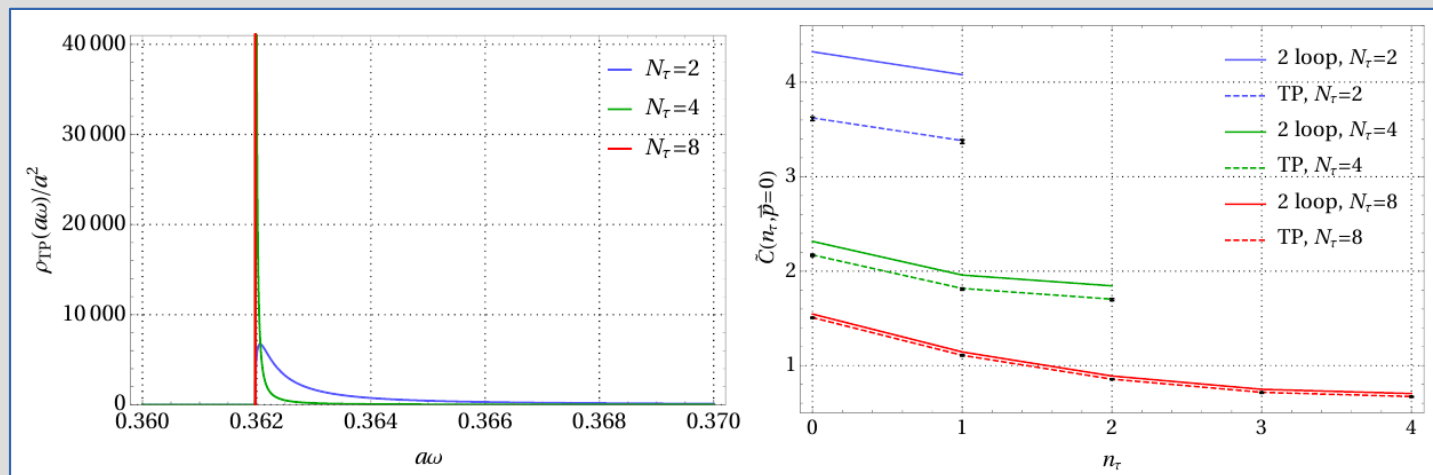
Light-light pseudo-scalar meson (pion) channel  
[P.L., O. Philipsen, *JHEP* 10, 161 (2022)]



Light-strange (kaon) and strange-strange (eta) pseudo-scalar meson channels [D. Bala, O. Kaczmarek, P. L., O. Philipsen, and T. Ueding, *JHEP* 05, 332 (2024)]

# 3. A potential resolution

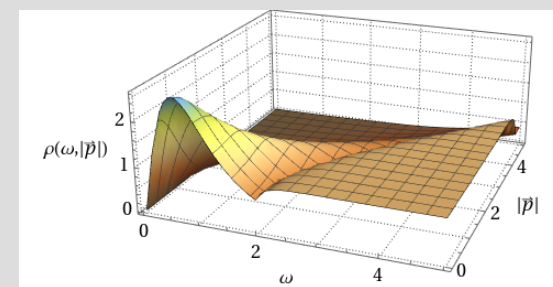
- Thermoparticles dominate the large-time behaviour of correlators and are therefore natural candidates for describing scattering states when  $T > 0$
- Scattering states determine the form of the basic field propagators used in any perturbative QFT expansion  $\rightarrow$  *Thermoparticle propagators could form the basis of a consistent  $T > 0$  perturbative expansion*
- Note: It has been shown that these propagators can be *uniquely* fixed by the dynamical equations of the theory [Bros, Buchholz, 2002]
- For  $\Phi^4$  theory there appear to be signatures of these excitations in the data



See: [PL, O. Philipsen, 2405.02009]

# Summary & outlook

- Perturbation theory has several known issues related to the infrared structure of certain classes of diagrams when  $T > 0$
- This seems to be a realisation of a more fundamental constraint: non-trivial scattering at  $T > 0$  is incompatible with real dispersion laws (*NRT theorem*)
- There is now both analytic (Weldon, 2002) and numerical (PL, Philipsen, 2024) evidence for this, even for massive theories, and at small  $T$
- A potential resolution of these issues is to use propagators which don't have real dispersion laws  $\rightarrow$  thermoparticles are a natural candidate
- The analytic structure of thermoparticle propagators can be derived for specific QFTs (Bros, Buchholz, 2002), and therefore might provide a consistent perturbative framework when  $T > 0$ . *Stay tuned!*



# Backup: (Axiomatic) QFT in a nutshell

- QFT framework defined by a core set of axioms:

**Axiom 1 (Hilbert space structure).** *The states of the theory are rays in a Hilbert space  $\mathcal{H}$  which possesses a continuous unitary representation  $U(a, \alpha)$  of the Poincaré spinor group  $\overline{\mathcal{P}}_+^\uparrow$ .*

**Axiom 2 (Spectral condition).** *The spectrum of the energy-momentum operator  $P^\mu$  is confined to the closed forward light cone  $\mathcal{V}^+ = \{p^\mu \mid p^2 \geq 0, p^0 \geq 0\}$ , where  $U(a, 1) = e^{iP^\mu a_\mu}$ .*

**Axiom 3 (Uniqueness of the vacuum).** *There exists a unit state vector  $|0\rangle$  (the vacuum state) which is a unique translationally invariant state in  $\mathcal{H}$ .*

**Axiom 4 (Field operators).** *The theory consists of fields  $\varphi^{(\kappa)}(x)$  (of type  $(\kappa)$ ) which have components  $\varphi_l^{(\kappa)}(x)$  that are operator-valued tempered distributions in  $\mathcal{H}$ , and the vacuum state  $|0\rangle$  is a cyclic vector for the fields.*

**Axiom 5 (Relativistic covariance).** *The fields  $\varphi_l^{(\kappa)}(x)$  transform covariantly under the action of  $\overline{\mathcal{P}}_+^\uparrow$ :*

$$U(a, \alpha)\varphi_l^{(\kappa)}(x)U(a, \alpha)^{-1} = S_{ij}^{(\kappa)}(\alpha^{-1})\varphi_j^{(\kappa)}(\Lambda(\alpha)x + a)$$

*where  $S(\alpha)$  is a finite dimensional matrix representation of the Lorentz spinor group  $\overline{\mathcal{L}}_+^\uparrow$ , and  $\Lambda(\alpha)$  is the Lorentz transformation corresponding to  $\alpha \in \overline{\mathcal{L}}_+^\uparrow$ .*

**Axiom 6 (Local (anti-)commutativity).** *If the support of the test functions  $f, g$  of the fields  $\varphi_l^{(\kappa)}, \varphi_m^{(\kappa')}$  are space-like separated, then:*

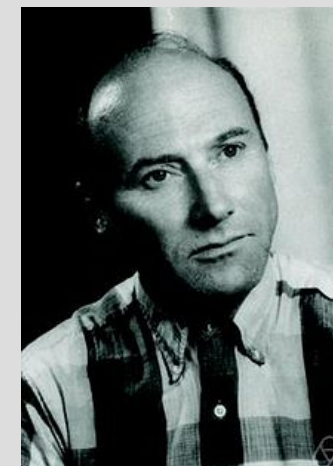
$$[\varphi_l^{(\kappa)}(f), \varphi_m^{(\kappa')}(g)]_\pm = \varphi_l^{(\kappa)}(f)\varphi_m^{(\kappa')}(g) \pm \varphi_m^{(\kappa')}(g)\varphi_l^{(\kappa)}(f) = 0$$

*when applied to any state in  $\mathcal{H}$ , for any fields  $\varphi_l^{(\kappa)}, \varphi_m^{(\kappa')}$ .*



A. Wightman

[R. F. Streater and A. S. Wightman, *PCT, Spin and Statistics, and all that* (1964).]



R. Haag

[R. Haag, *Local Quantum Physics*, Springer-Verlag (1996).]

# Backup: *Particle signatures in vacuum*

- Correlation functions  $\langle 0 | \varphi_{l_1}^{(\kappa_1)}(x_1) \cdots \varphi_{l_n}^{(\kappa_n)}(x_n) | 0 \rangle$  encode **all** of the dynamical information, including the properties of particles!
- For the purpose of this talk I will focus on (scalar) 2-point correlation functions, in particular the *spectral function*, defined by

$$\rho(\omega, \vec{p}) = \int d^4x e^{i(\omega x_0 - \vec{p} \cdot \vec{x})} \langle [\phi(x), \phi(0)] \rangle$$

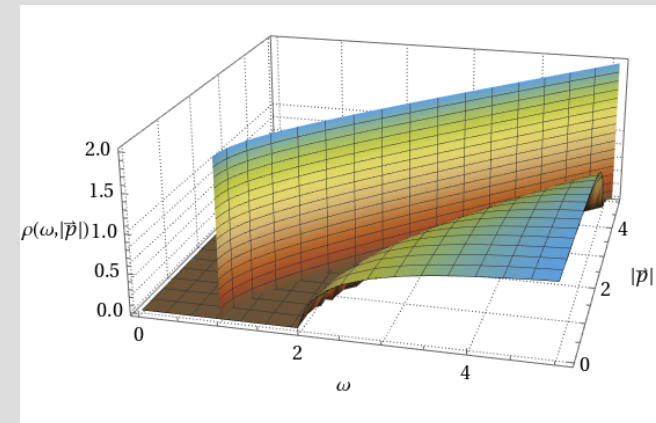
- The simplest case: a non-interacting particle with mass  $m$

$$\rho(\omega, \vec{p}) = 2\pi \epsilon(\omega) \delta(\omega^2 - |\vec{p}|^2 - m^2)$$

... and when we add interactions

Delta peak  $\rightarrow$  stable particle component

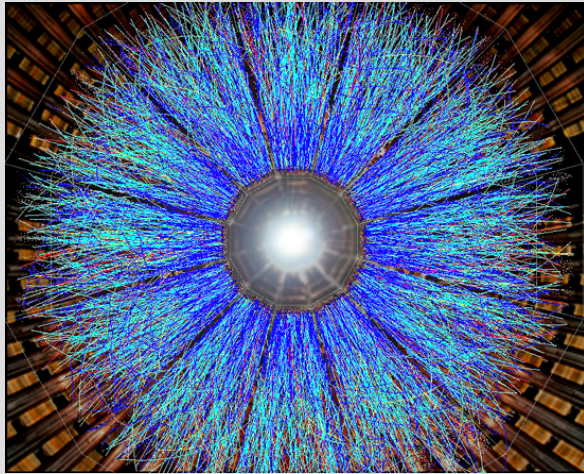
Continuum onset  $\rightarrow$  multi-particle states



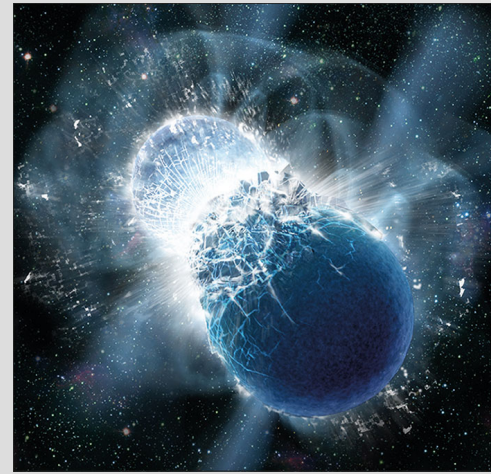


# Backup: *Beyond the vacuum*

- To describe physical phenomena in “extreme environments” one must understand of how QFT applies to systems that are hot, dense, or both



[Brookhaven National Lab]



[Skyworks Digital Inc.]

- Therefore need to figure out how the inclusion of temperature  $T = 1/\beta$  or density modifies the standard QFT assumptions, and what effect this has on the correlation functions
  - Let's stick to the case  $T > 0$  and vanishing density

# Backup: *Beyond the vacuum*

- **Idea:** Look for a generalisation of the standard axioms which is compatible with  $T > 0$ , and approaches the vacuum case for  $T \rightarrow 0$

**Axiom 1 (Hilbert space structure).** *The states of the theory are rays in a Hilbert space  $\mathcal{H}$  which possesses a continuous unitary representation  $U(a, \alpha)$  of the Poincaré spinor group  $\overline{\mathcal{P}}_+^\uparrow$ .*

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$$[\varphi_l^{(\kappa)}(f), \varphi_m^{(\kappa')}(g)]_\pm = \varphi_l^{(\kappa)}(f)\varphi_m^{(\kappa')}(g) \pm \varphi_m^{(\kappa')}(g)\varphi_l^{(\kappa)}(f) = 0$$

*when applied to any state in  $\mathcal{H}$ , for any fields  $\varphi_l^{(\kappa)}, \varphi_m^{(\kappa')}$ .*



$H_\beta$  is defined for fixed  $\beta=1/T$



Replaced by the KMS condition

$$\begin{aligned} & \langle \Omega_\beta | \phi(x_1) \cdots \phi(x_k) \phi(x_{k+1}) \cdots \phi(x_n) | \Omega_\beta \rangle \\ &= \langle \Omega_\beta | \phi(x_{k+1}) \cdots \phi(x_n) \phi(x_1 + i(\beta, \vec{0})) \cdots \phi(x_k + i(\beta, \vec{0})) | \Omega_\beta \rangle \end{aligned}$$



Instead, thermal background state  $|\Omega_\beta\rangle$



Fields are still distributions



The fields no longer transform under general unitary Lorentz transformations



Causality is unaffected by the properties of the background state.  
*This is important!*

# Backup: *Beyond the vacuum*

- For  $T > 0$  there are some immediate modifications:
  - **Lorentz invariance** ✗ → but can retain rotational invariance
  - **Spectral condition** ✗ → replaced by equilibrium (KMS) condition
  - **Field locality (causality)** ✓ → this is important!
- Taking all of these  $T > 0$  constraints into account implies that the spectral function has the representation \*

$$\rho(\omega, \vec{p}) = \int_0^\infty ds \int \frac{d^3 \vec{u}}{(2\pi)^2} \epsilon(\omega) \delta(\omega^2 - (\vec{p} - \vec{u})^2 - s) \tilde{D}_\beta(\vec{u}, s)$$

This is the  $T > 0$  generalisation of the textbook *Källén-Lehmann* representation!

$$\rho(\omega, \vec{p}) = 2\pi\epsilon(\omega) \int_0^\infty ds \delta(p^2 - s) \varrho(s)$$

“Thermal spectral density”

- $T > 0$  effects amount to understanding:  $\rho(s) \rightarrow \tilde{D}_\beta(\mathbf{u}, s)$ , which tell us about the possible excitations that can exist in a thermal medium

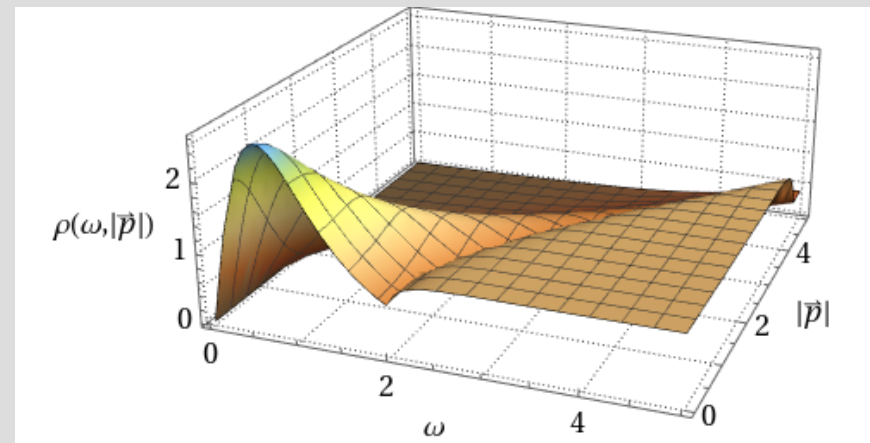
\* See: *J. Bros and D Buchholz, Z. Phys. C 55 (1992), Ann. Inst. H.Poincaré Phys.Theor. 64 (1996)*

# Backup: A thermal particle? $\rightarrow$ “Thermoparticle”

Proposition: the medium contains “Thermoparticles”: particle-like excitations which differ from collective quasi-particle modes, and show up as **discrete** contributions to  $\tilde{D}_\beta(\mathbf{u}, s)$  [Bros, Buchholz, *NPB* 627 (2002)]

$$\tilde{D}_\beta(\vec{u}, s) = \tilde{D}_{m,\beta}(\vec{u}) \delta(s - m^2) + \tilde{D}_{c,\beta}(\vec{u}, s)$$

- $\rightarrow$  Thermoparticle components  $\tilde{D}_\beta(\mathbf{u})\delta(s-m^2)$  reduce to those of a vacuum particle with mass  $m$  in the limit  $T \rightarrow 0$
- $\rightarrow$  Non-trivial “Damping factor”  $\tilde{D}_\beta(\mathbf{u})$  results in thermally-broadened peaks in the spectral function: this parametrises the effects of collisional broadening
- $\rightarrow$  Component  $\tilde{D}_{c,\beta}(\mathbf{u}, s)$  contains all other types of excitations, including those that are *continuous* in  $s$



# Backup: *Perturbation theory setup*

- Perturbation theory for  $T = 1/\beta > 0$  can be performed using either real or imaginary-time propagators. Both have their advantages and disadvantages (see: Kapusta & Gale, Le Bellac)
- Here we focus on the imaginary time  $\tau$  formalism. The basic idea is that one uses imaginary time “Matsubara” propagators which have discrete energies  $\omega_n = 2\pi n/\beta$

$$G(\omega_n, \vec{p}) = \frac{1}{\omega_n^2 + |\vec{p}|^2 + m^2}$$

- This discretisation comes about because the Euclidean 2-point function  $W_E(\tau, \mathbf{x})$  is  $\beta$ -periodic in  $\tau$  (due to the KMS condition)

→ At the end one analytically continues the result:  $i\omega_n \rightarrow \omega + i\varepsilon$

- The perturbation theory rules are analogous to the  $T=0$  case, except now each internal line is integrated with the measure:

$$T \sum_n \int \frac{d^3 p}{(2\pi)^3}$$

# Backup: *Perturbation theory results*

- Results for  $L=16$ :

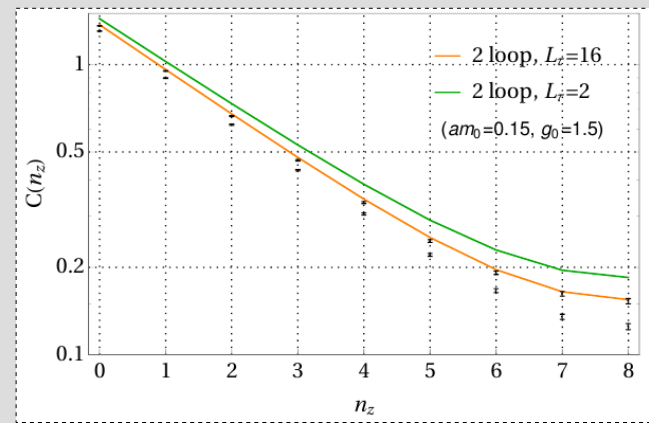
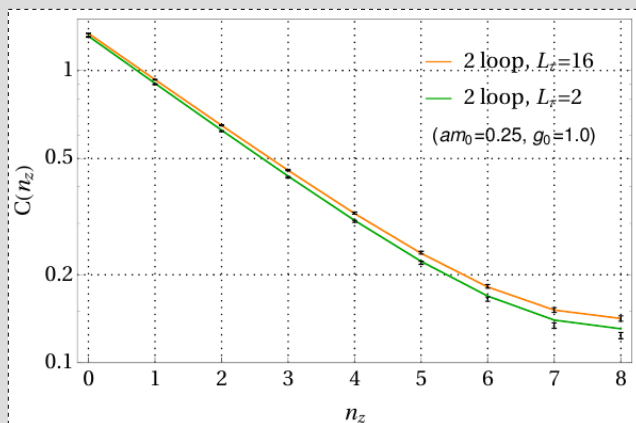
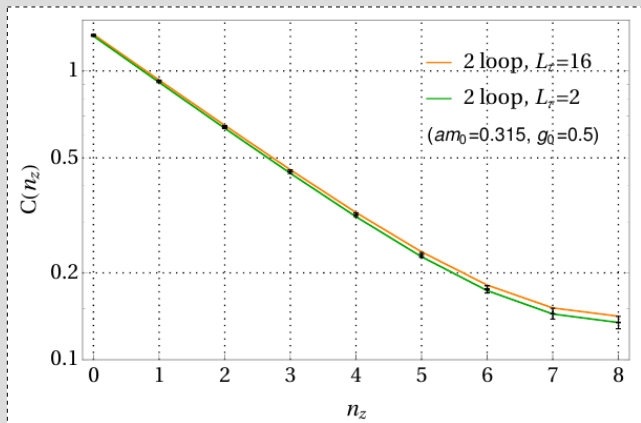
For small coupling the results agree with data for all  $T/m$

For large coupling the vacuum-like data differs from prediction

→ Higher-loop corrections needed!

$L^3 \times L_\tau$	$(am_0, g_0)$	$m/\Lambda_c$	$T/m$	$(\chi^2/\text{d.o.f.})_{2\text{-loop}}$	$\Delta_{1\text{-loop}}^\ell$ [%]	$\Delta_{2\text{-loop}}^\ell$ [%]
$16^3 \times 16$	(0.315, 0.5)	0.118	0.17	1.7	0.4 (0.6)	0.8 (0.6)
$16^3 \times 8$	(0.315, 0.5)	0.118	0.34	0.4	0.2 (0.6)	0.2 (0.6)
$16^3 \times 4$	(0.315, 0.5)	0.118	0.68	0.3	1.0 (0.6)	0.4 (0.6)
$16^3 \times 2$	(0.315, 0.5)	0.118	1.35	0.5	1.2 (0.6)	0.4 (0.6)
$16^3 \times 16$	(0.25, 1.0)	0.117	0.17	0.1	2.0 (0.6)	0.2 (0.6)
$16^3 \times 8$	(0.25, 1.0)	0.117	0.34	0.7	1.1 (0.6)	0.8 (0.6)
$16^3 \times 4$	(0.25, 1.0)	0.117	0.68	0.1	2.4 (0.6)	0.1 (0.6)
$16^3 \times 2$	(0.25, 1.0)	0.117	1.36	3.9	3.3 (0.6)	0.9 (0.6)
$16^3 \times 16$	(0.15, 1.5)	0.115	0.17	10.3	4.1 (0.6)	2.3 (0.6)
$16^3 \times 8$	(0.15, 1.5)	0.115	0.35	49.7	4.6 (0.6)	4.2 (0.6)
$16^3 \times 4$	(0.15, 1.5)	0.115	0.69	188.5	6.0 (0.6)	8.7 (0.7)
$16^3 \times 2$	(0.15, 1.5)	0.115	1.38	1262.3	8.7 (0.5)	22.2 (0.7)

→ In all cases where the interactions are non-negligible, the predictions deteriorate as  $T/m$  increases



# Backup: *Spectral properties from QCD data*

**Goal:** Extract information about the finite  $T$  spectral function  $\rho_r(\omega, \mathbf{p})$  from data of *Euclidean* correlator  $C_\Gamma(\tau, \vec{x}) = \langle O_\Gamma(\tau, \vec{x}) O_\Gamma(0, \vec{0}) \rangle_T$   $O_r =$  scalar operator

- Standard approach: extract  $\rho_r(\omega, \mathbf{p})$  from temporal correlator  $\tilde{C}_r(\tau, \mathbf{p})$

$$\tilde{C}_\Gamma(\tau, \vec{p}) = \int_0^\infty \frac{d\omega}{2\pi} \frac{\cosh \left[ \left( \frac{\beta}{2} - |\tau| \right) \omega \right]}{\sinh \left( \frac{\beta}{2} \omega \right)} \rho_\Gamma(\omega, \vec{p}) \rightarrow \text{Problem is ill-conditioned, need more information!}$$

- Instead, one can use the **spatial** correlator, where one integrates  $C_r(\tau, \mathbf{x})$  over  $\{\tau, x, y\}$  and fixes a spatial direction  $z$

$$C_\Gamma(z) = \int_{-\infty}^{\infty} \frac{dp_z}{2\pi} e^{ip_z z} \int_0^\infty \frac{d\omega}{\pi\omega} \rho_\Gamma(\omega, p_x = p_y = 0, p_z)$$

- It turns out that *if* thermoparticles exist, then they will give a distinct contribution to  $C(z)$

$$C(z) \approx \frac{1}{2} \int_{|z|}^{\infty} dR e^{-mR} D_{m,\beta}(R)$$

[P.L., *PRD* 106 (2022); P.L., O. Philipsen, *JHEP* 10, 161 (2022)]

→ This component can be extracted *directly* from data

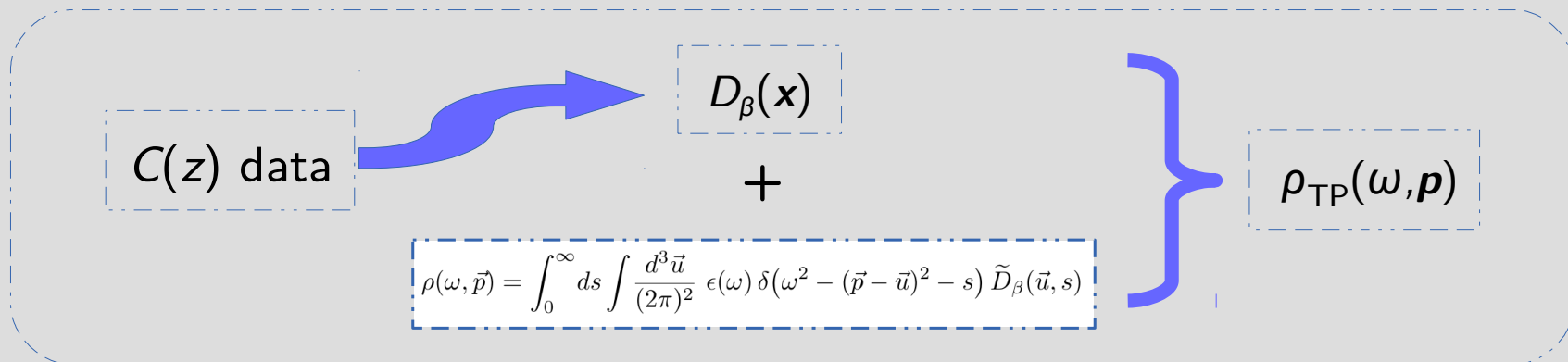
# Backup: *Spectral properties from QCD data*

- But what impact do **thermoparticles** have on the behaviour of  $C(z)$ ?

→ They give contributions that *dominate* for large  $z$

$$C(z) \approx \frac{1}{2} \int_{|z|}^{\infty} dR e^{-mR} D_{m,\beta}(R)$$

- If they exist, one can extract their damping factors  $D_{\beta}(\mathbf{x})$  *directly* from  $C(z)$  and then determine the spectral contribution  $\rho_{\text{TP}}(\omega, \mathbf{p})$



- Apply approach to QCD lattice data → simple case is the spatial correlator  $C_{\text{PS}}(z)$  of the pseudo-scalar meson operator  $\mathcal{O}_{\text{PS}}^a = \bar{\psi} \gamma_5 \frac{\tau^a}{2} \psi$

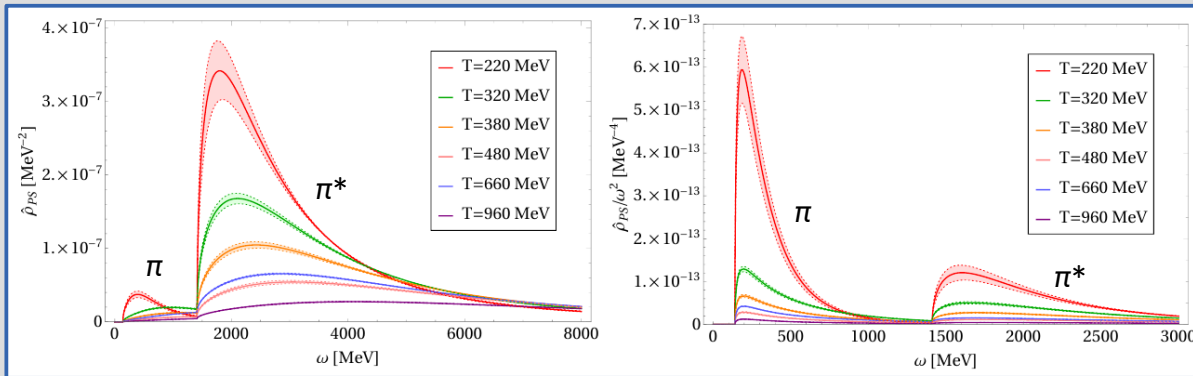
→ Done for light-light (pion), light-strange (kaon), and strange-strange (eta)

quarks [P.L., O. Philipsen, 2022; D. Bala, O. Kaczmarek, P.L., O. Philipsen, and T. Ueding, 2310.13476]



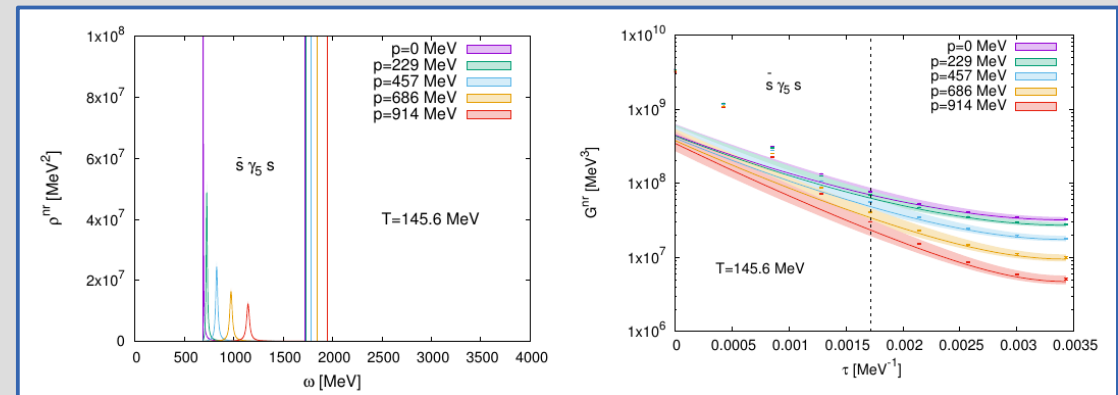
# Backup: Spectral properties from QCD data

- Apply approach to lattice QCD data → simple case is the spatial correlator  $C_{PS}(z)$  of the pseudo-scalar meson operator  $\mathcal{O}_{PS}^a = \bar{\psi}\gamma_5\frac{\tau^a}{2}\psi$
- Analysis performed for various pseudo-scalar meson operators:



Light-light pseudo-scalar meson (pion) channel [P.L., O. Philipsen, *JHEP* 10, 161 (2022)]

Light-strange (kaon) and strange-strange (eta) pseudo-scalar meson channels [D. Bala, O. Kaczmarek, P. L., O. Philipsen, and T. Ueding, *JHEP* 05, 332 (2024)]



Data in *all* channels consistent with a thermoparticle-type ground state: suggests light pseudo-scalar mesons (pions, kaons,..) still have a bound-state-like structure, even at high  $T$

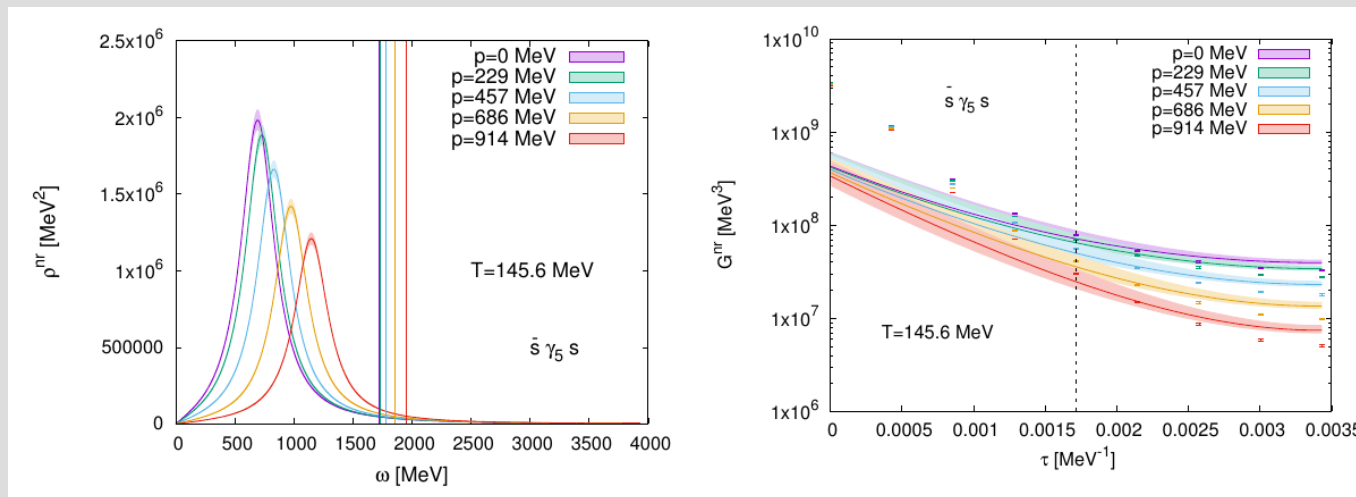
# Backup: *Spectral properties from QCD data*

- The robustness of the thermoparticle hypothesis can also be tested by comparing with different causal models, e.g. a Breit Wigner

$$\rho_{\text{BW}}(\omega, \vec{p}) = \frac{4\omega\Gamma}{(\omega^2 - |\vec{p}|^2 - m^2 - \Gamma^2)^2 + 4\omega^2\Gamma^2},$$

$$C_{\text{BW}}(z) = \frac{e^{-\sqrt{m^2 + \Gamma^2}|z|}}{2\sqrt{m^2 + \Gamma^2}}.$$

- Same procedure as with the thermoparticle case: (i) extract the width parameter  $\Gamma$  and coefficient from the spatial lattice data (ii) use this to predict the corresponding temporal correlator



→ Data is *not* consistent with a Breit-Wigner-type ground state!

# Backup: An alternative approach

→ But given a specific QFT, what form should these components take?

Idea: thermal scattering states are defined by imposing an asymptotic field condition [Bros, Buchholz, (2002)]:

Asymptotic fields  $\Phi_0$  are assumed to satisfy dynamical equations, but only at large  $x_0$

In  $\Phi^4$  theory

$$(\partial^2 + m^2)\phi_0(x) + \frac{\lambda}{3!}\phi_0^3(x) \xrightarrow{|x_0| \rightarrow \infty} 0$$

- Since thermoparticles dominate the large-time behaviour of correlators, they are natural candidates for describing such states. It turns out that their damping factors  $\tilde{D}_{m,\beta}(\mathbf{u})$  are **uniquely fixed** by the asymptotic condition
- In  $\Phi^4$  theory one finds (where  $\kappa$  is a thermal width):

For  $g < 0$ :

$$\tilde{D}_{m,\beta}^{(-)}(\vec{u}) = \frac{2\pi^2}{\kappa^2} \delta(|\vec{u}| - \kappa),$$

For  $g > 0$ :

$$\tilde{D}_{m,\beta}^{(+)}(\vec{u}) = \frac{4\pi}{\kappa_0 (|\vec{u}|^2 + \kappa^2)},$$

$$\tilde{G}_{\beta}^{(-)}(k_0, \vec{p}) = \frac{1}{4|\vec{p}|\kappa} \ln \left[ \frac{-k_0^2 + m^2 + (|\vec{p}| + \kappa)^2}{-k_0^2 + m^2 + (|\vec{p}| - \kappa)^2} \right]$$

$$\tilde{G}_{\beta}^{(+)}(k_0, \vec{p}) = \frac{i}{2|\vec{p}|\kappa_0} \ln \left[ \frac{\sqrt{-k_0^2 + m^2 - i|\vec{p}| + \kappa}}{\sqrt{-k_0^2 + m^2 + i|\vec{p}| + \kappa}} \right]$$